



Dust estimation via the
triple window IR technique
($8.7\mu\text{m}$, $10.8\mu\text{m}$, $12.0\mu\text{m}$)



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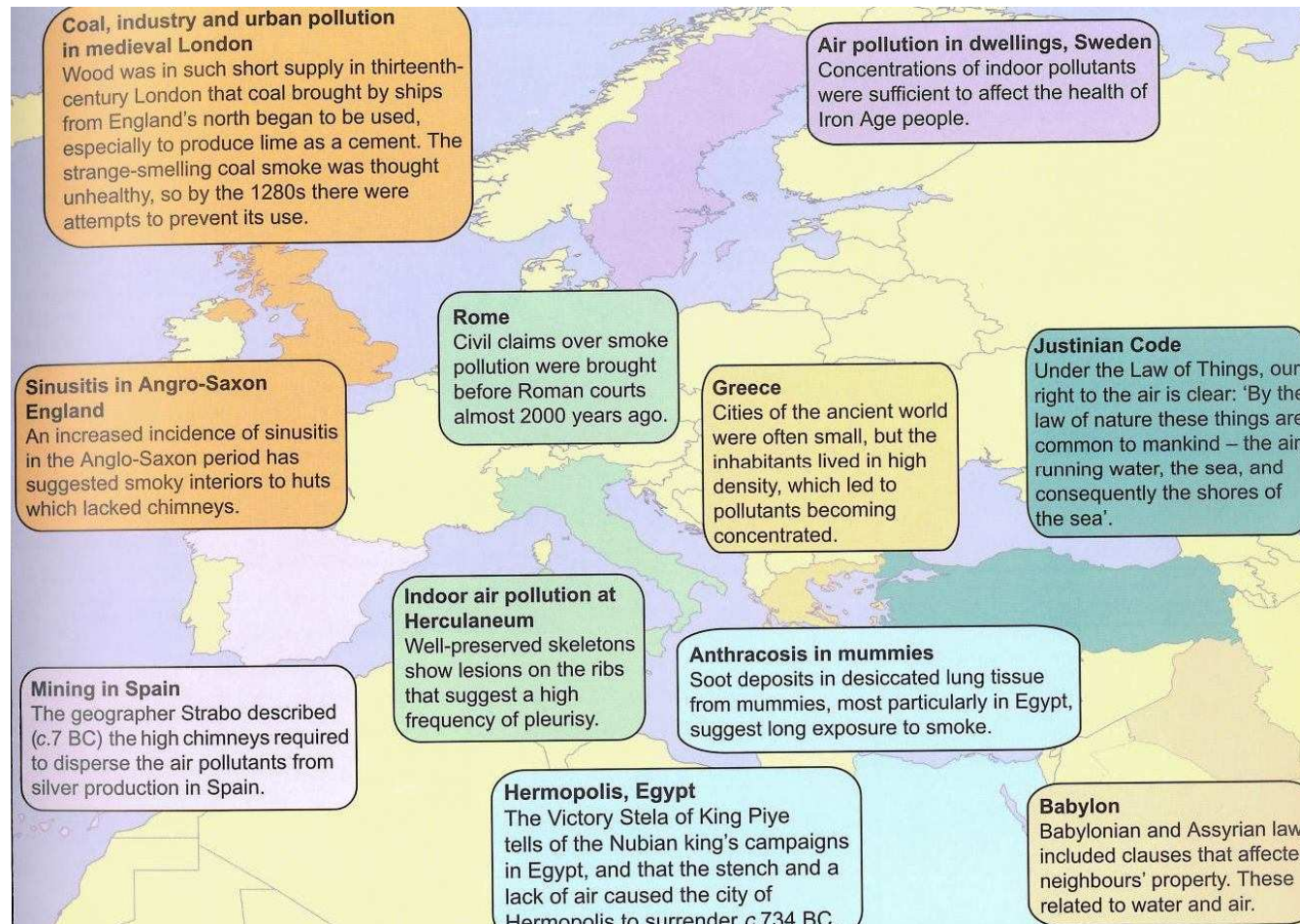


Satellite remote sensing of mineral dust

Dust estimation via infrared

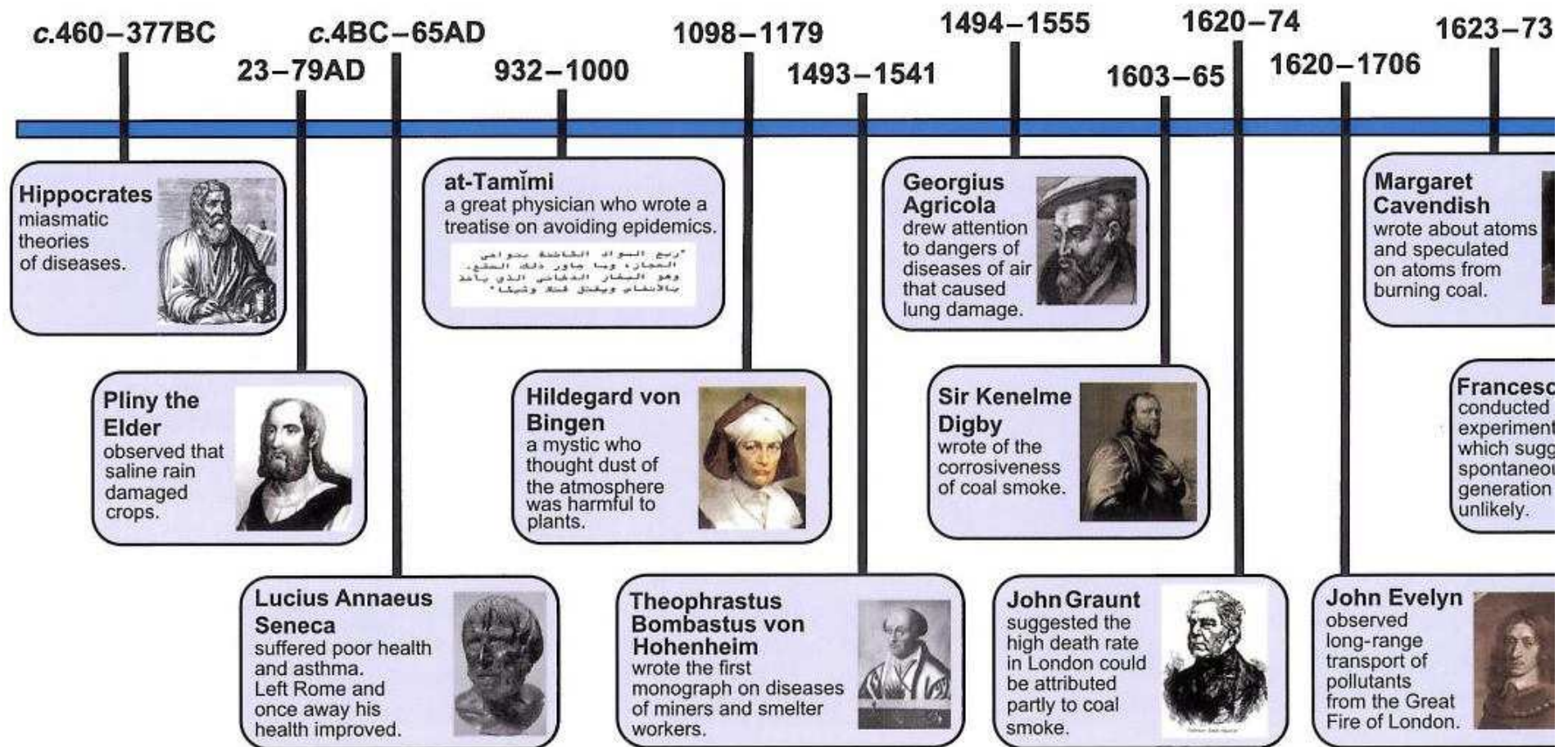
jose.prieto@eumetsat.int

Early reports on dust

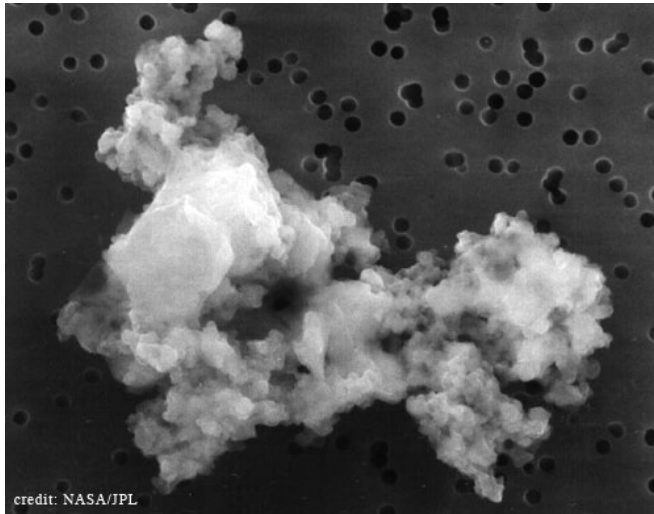


World Atlas of Atmospheric Pollution. Editor: R. S. Sokhi

... and on health



Is air dust important?



← Dust particle 12 μm →



← Earth globe 12 Mm →

Connection:

- 12 orders of magnitude difference in size
- 10^{12} kg in the atmosphere (10^{-7} of atmospheric mass)
- Climate
- Atmospheric circulation
- Life ?
- Dostojewski: “Love both, small and big!”

Optimal dust detection with the infrared?

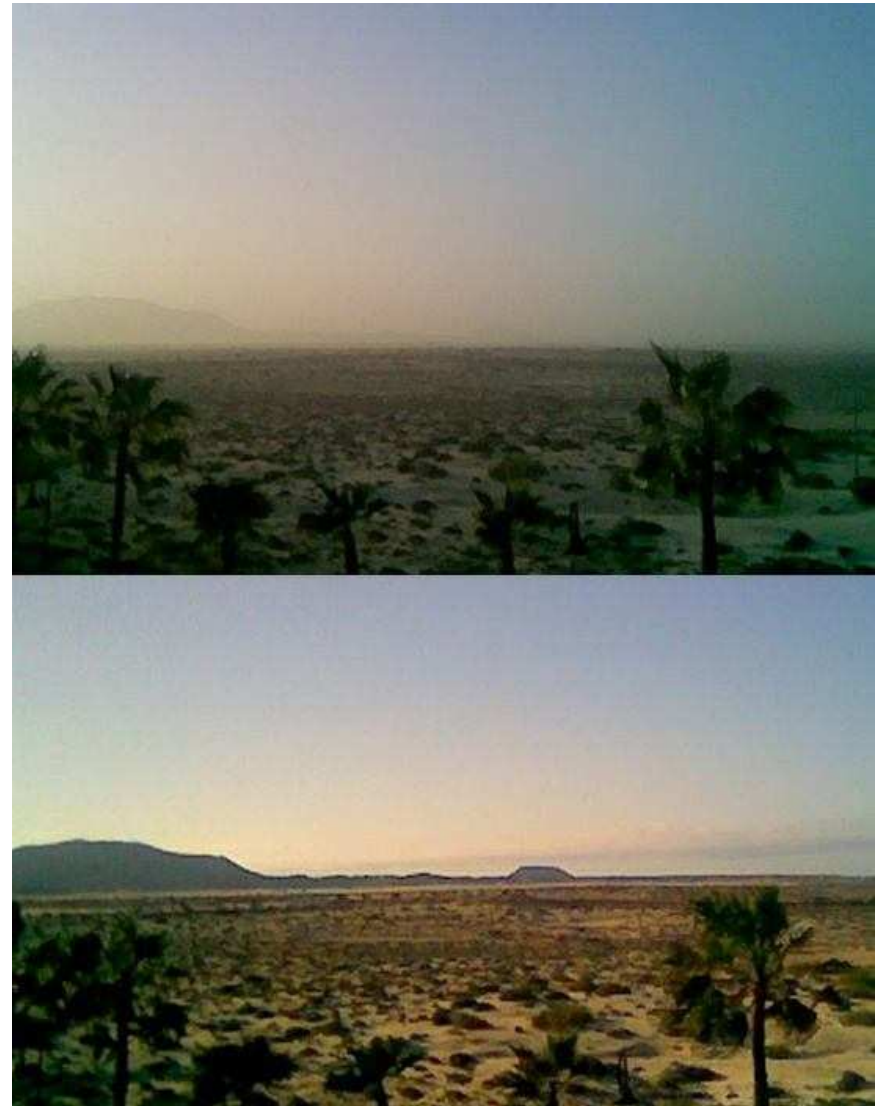
<i>Best contrast ?</i>	DAY	NIGHT
IR		
VIS		

Click one of the four fields, the one with best contrast free-surface/dust-storm

<i>Ocean</i>	DAY	NIGHT
IR	good	good
VIS	good	A/N/A

<i>Desert</i>	DAY	NIGHT
IR	good	bad
VIS	bad	A/N/A

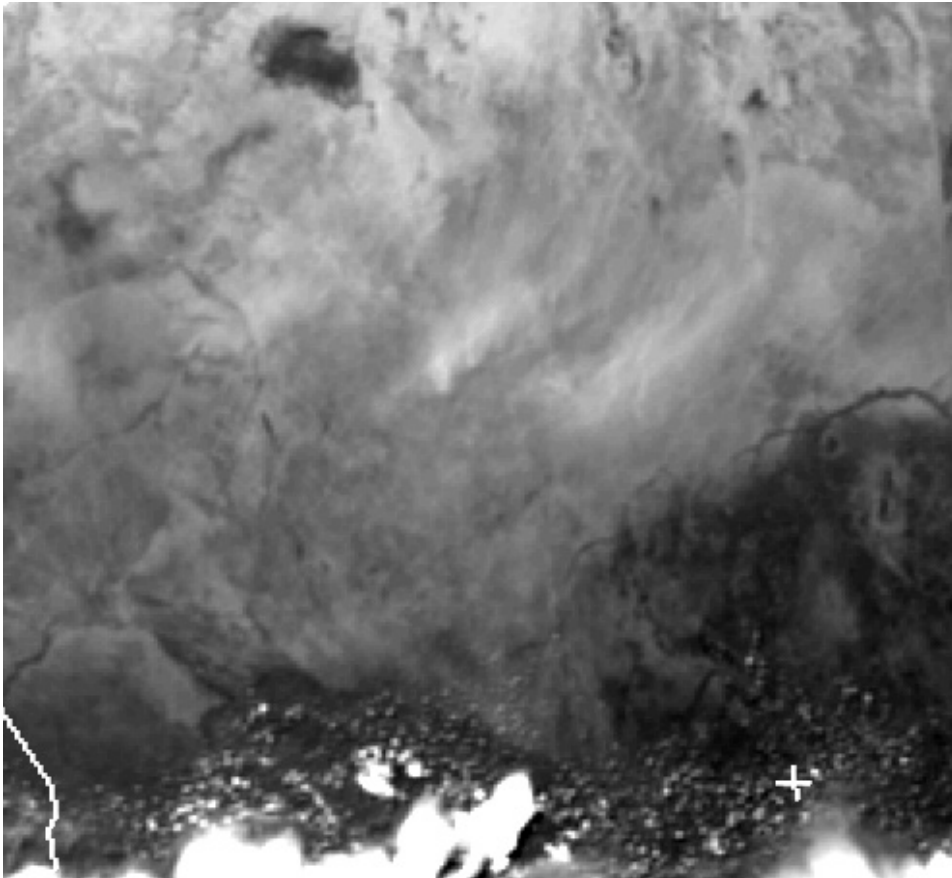
- On IR imagery, dusty air appears cool in contrast to the hot **daytime** land surface. At **night**, the thermal difference between the background and the dust lessens.
- On VIS imagery over **water**, dust is easy to note. Over **land**, however, the dust plume and the surface look similar



Consecutive days in Fuerteventura, January 2010

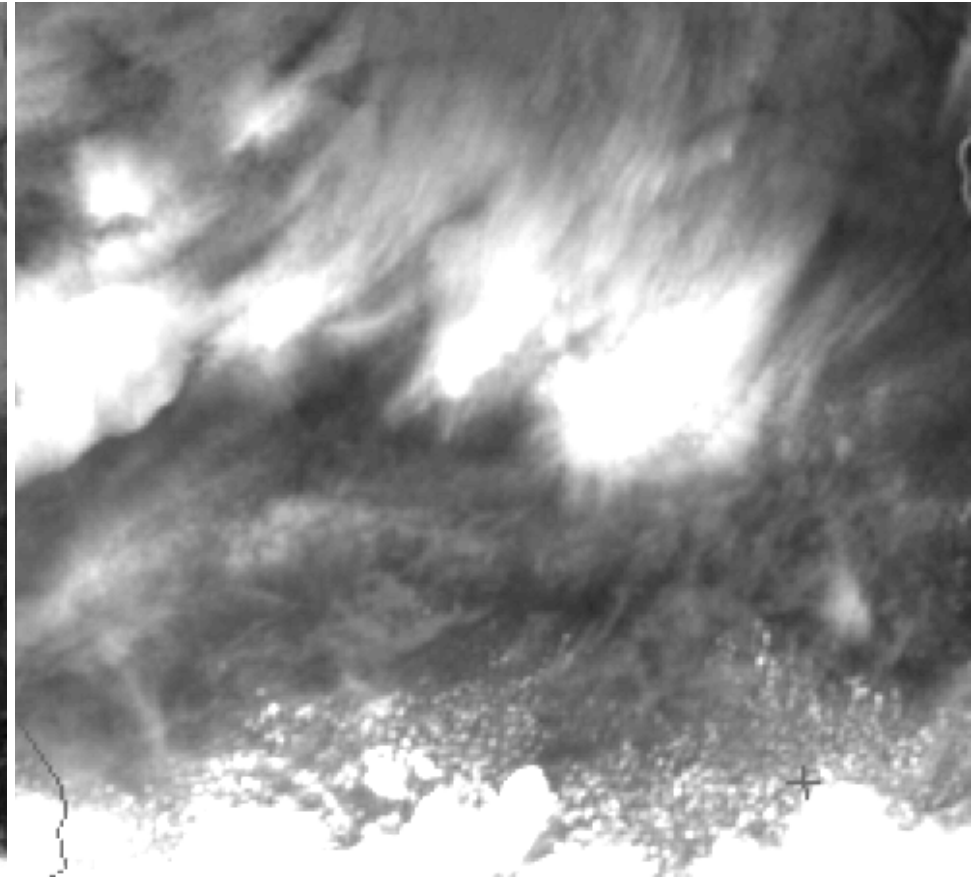
Dust on visible and infrared

- Disputed contribution to global cooling (S.K. Satheesh, 2006)



2004-05-13 13:00 UTC, 0.8 μm

- Dust **reflects** back solar energy to space



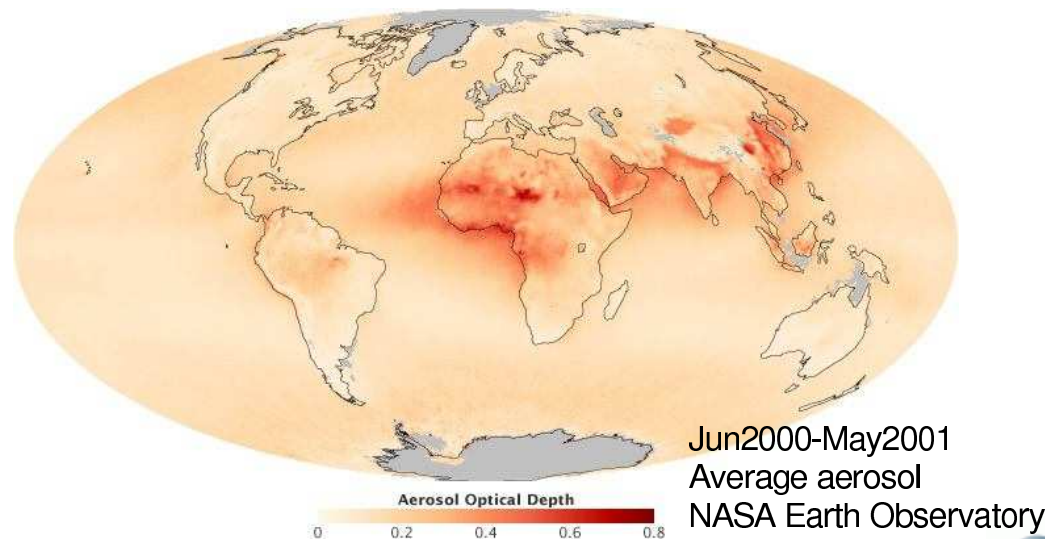
Same date-time, 10.8 μm

- Dusty air rises and **cools** down

Desert scene, Southern Sudan

Contents

- Dust characteristics
- Display techniques
- Inter-channel comparisons: plots
- Radiation transfer concepts
- Graphical retrieval
- Value added by 8.7 μm
- Conclusions



Characteristics

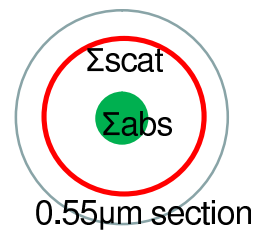
- Dust storms occasionally reach 5 km **height**, frequently thicker than 1km
- Dust optical depth around 0.5 (average on land) or 5 for storms, in the visible range. Optical **thickness** in the IR is about 40% of those values.
- TOMS Absorbing Aerosol Index (0-4) roughly equivalent of AOT (0-2)
- Aerosol **density** average in the atmosphere 10^{-7} kg/m³ (equivalent of depth 0.1)



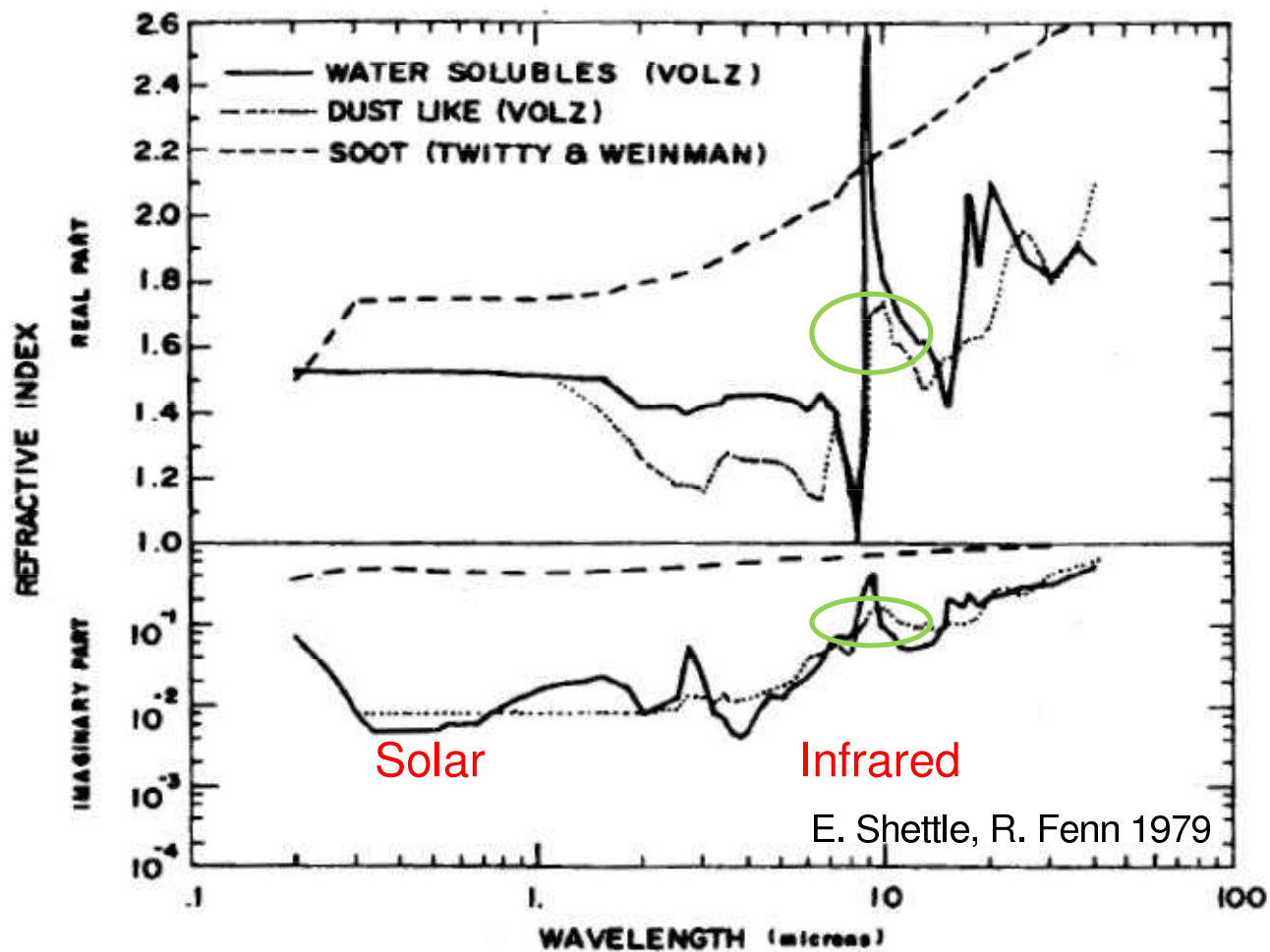
Foto: Eugenio Rodríguez

Characteristics

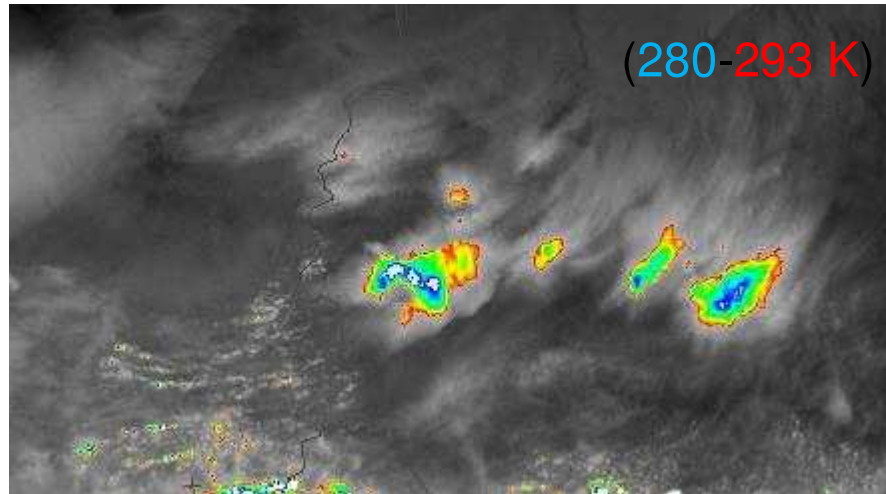
- Particle concentration for dust **storms** is about $5 \cdot 10^8$ /m³, similar to cloud
- Not understood transportation of sand particles ($>60\mu\text{m}$) beyond 1000km
- $n=1.53 + i \cdot 0.008$ in the visible range (similar to droplets), **higher absorption** in the IR
- Dust **size** is tri-modal around $0.1 \mu\text{m}$, $2 \mu\text{m}$ (marine salt), $40 \mu\text{m}$ (dust storm)
- Dust absorbs and scatters **infrared** radiation in the **Mie** range



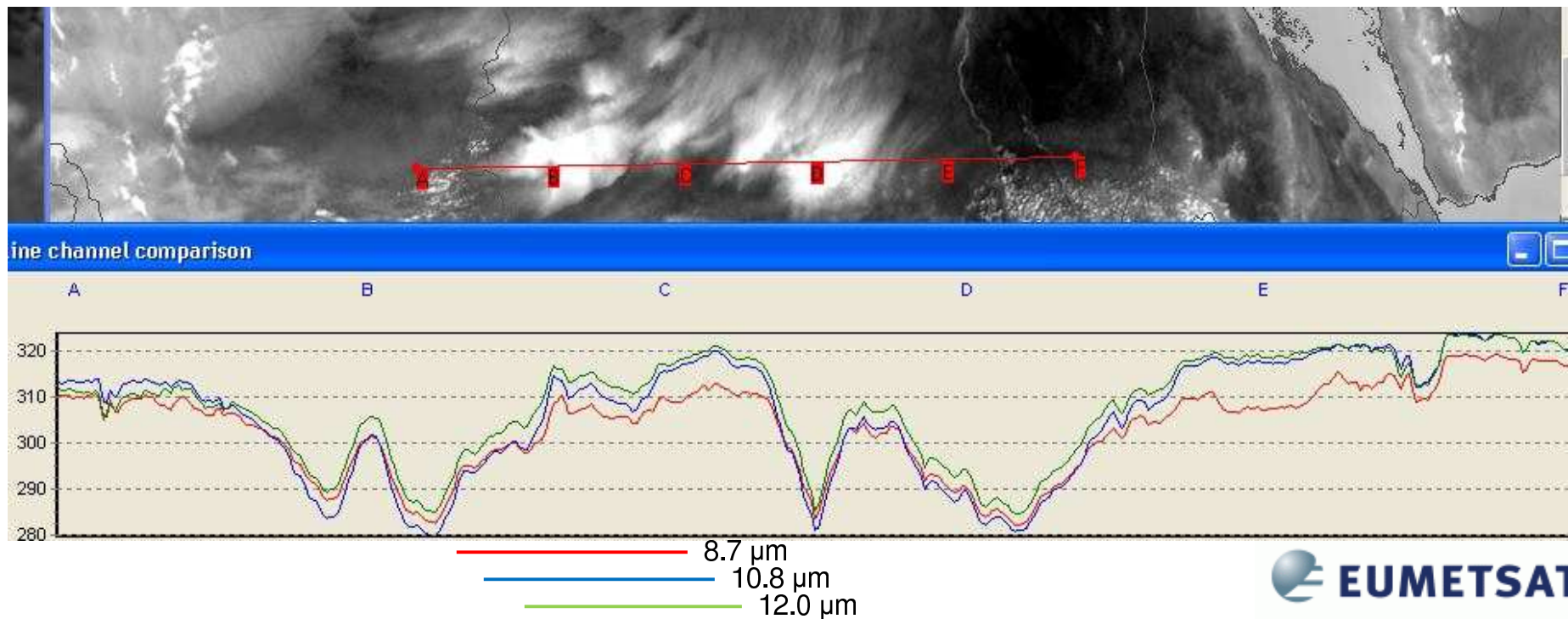
Characteristics

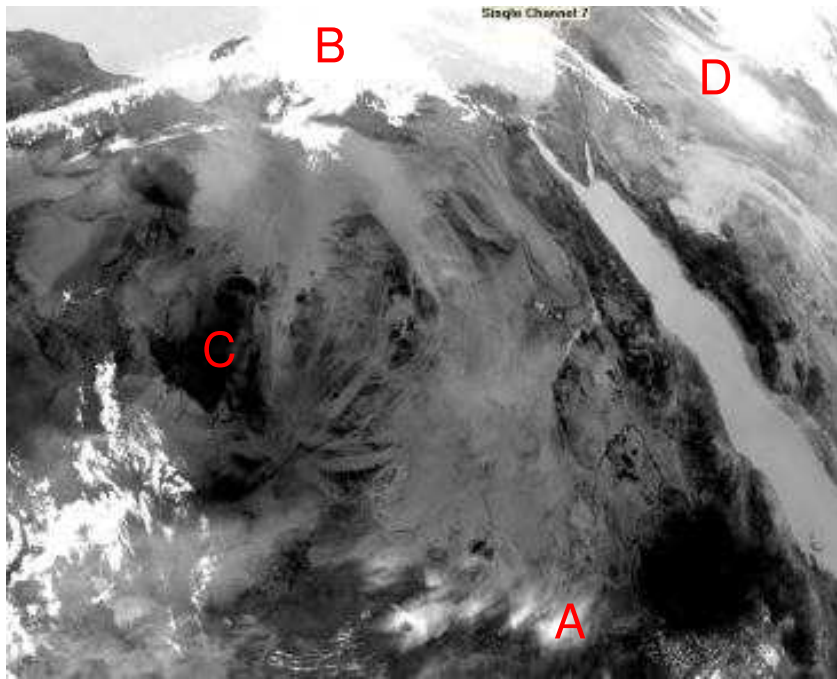


Dust enhancement in a single IR channel !?

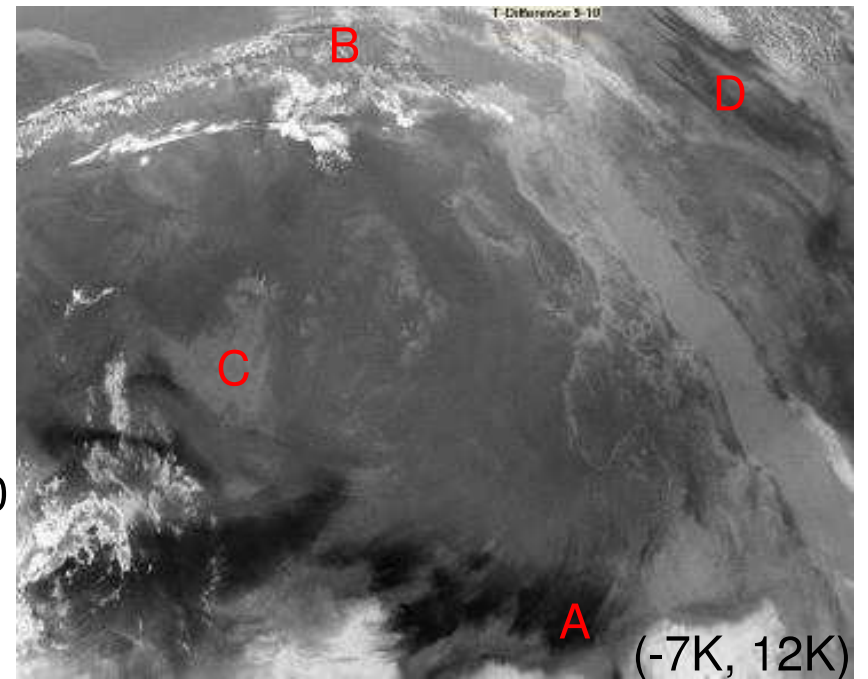


2004 May 13th 13:00 Meteosat **10.8 μ m**
colour-enhanced (left) and gray-enhanced (below)



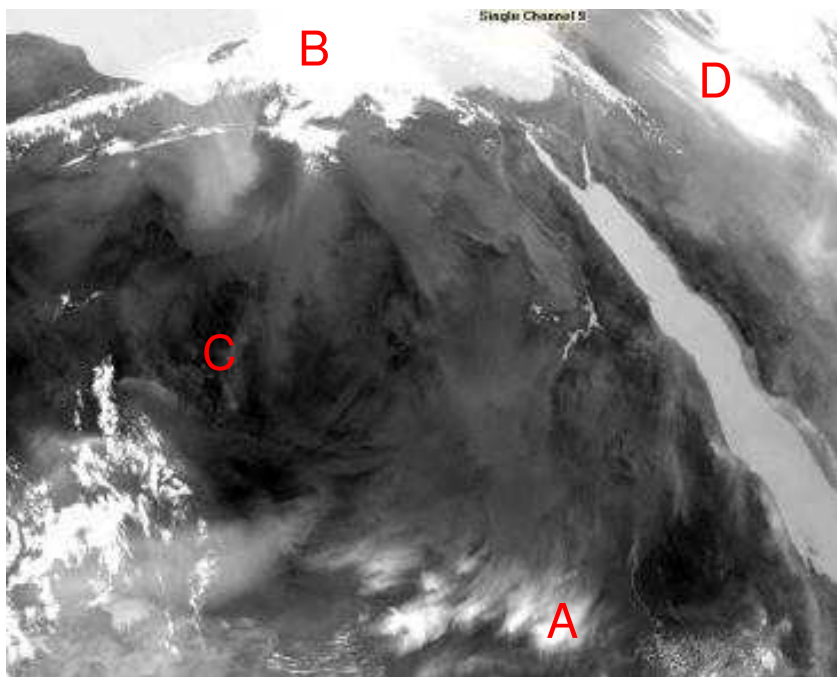


7

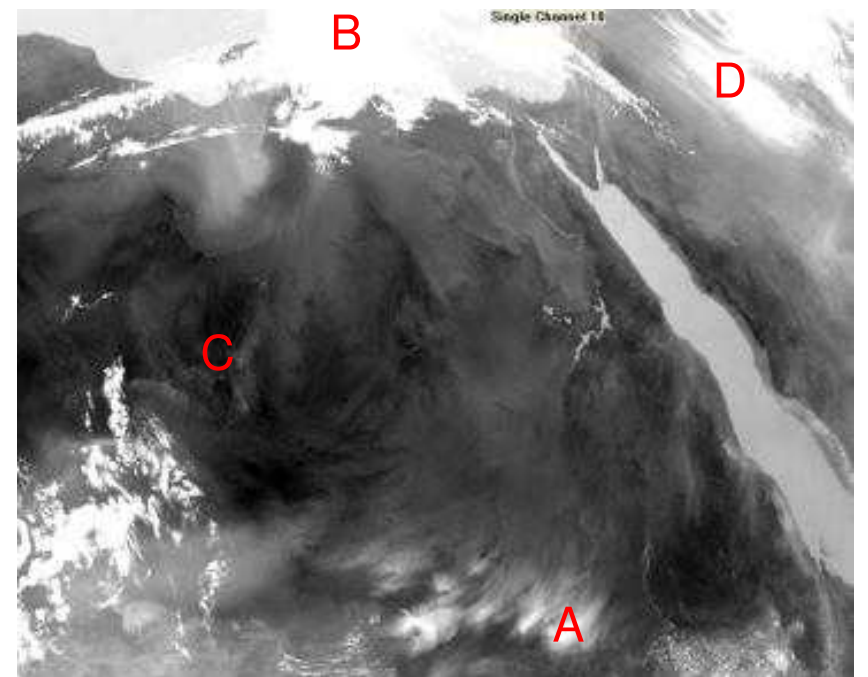


(-7K, 12K)

Single channel comparison and Ch9-Ch10 difference (*upper right*)

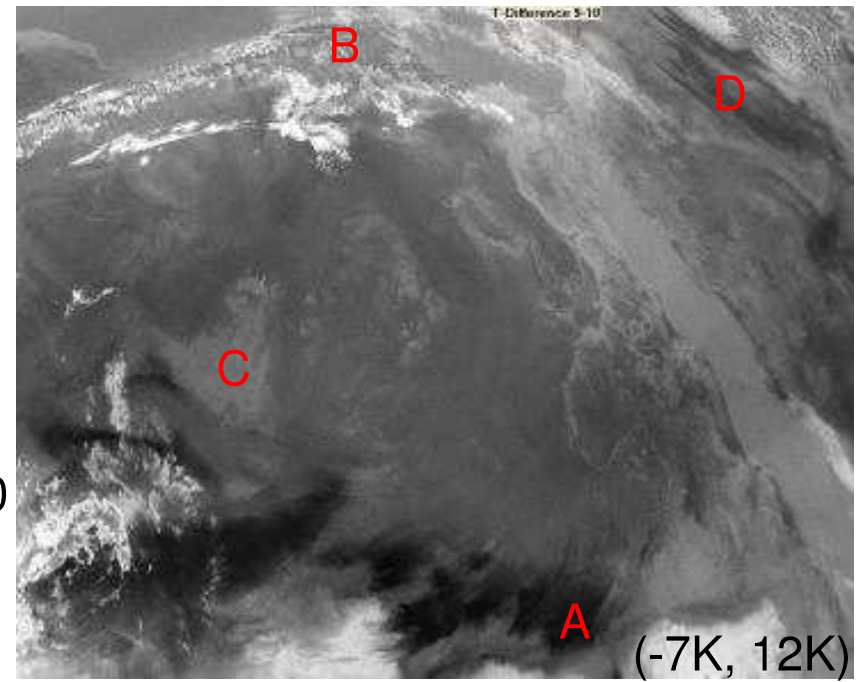


9

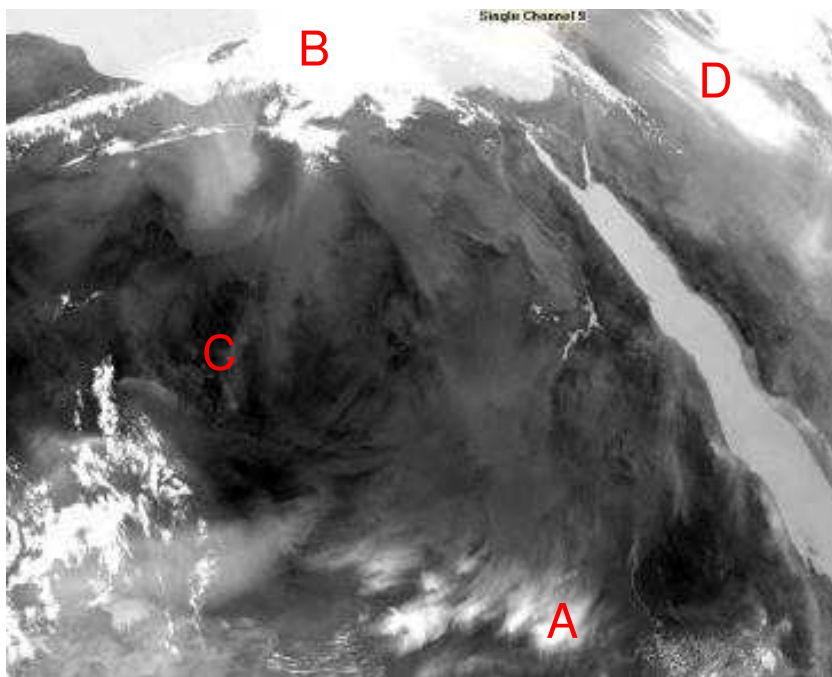


10

9-10

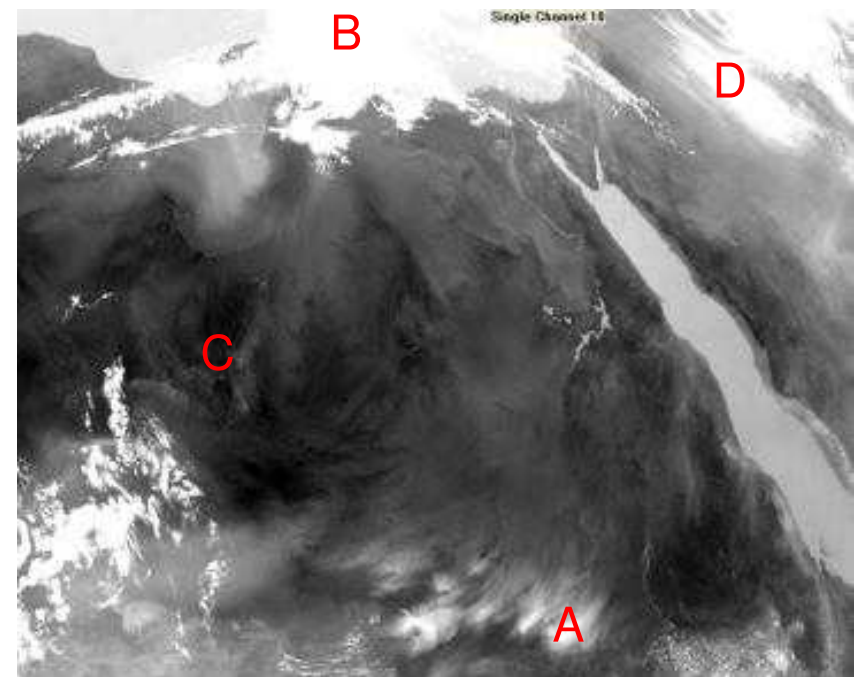


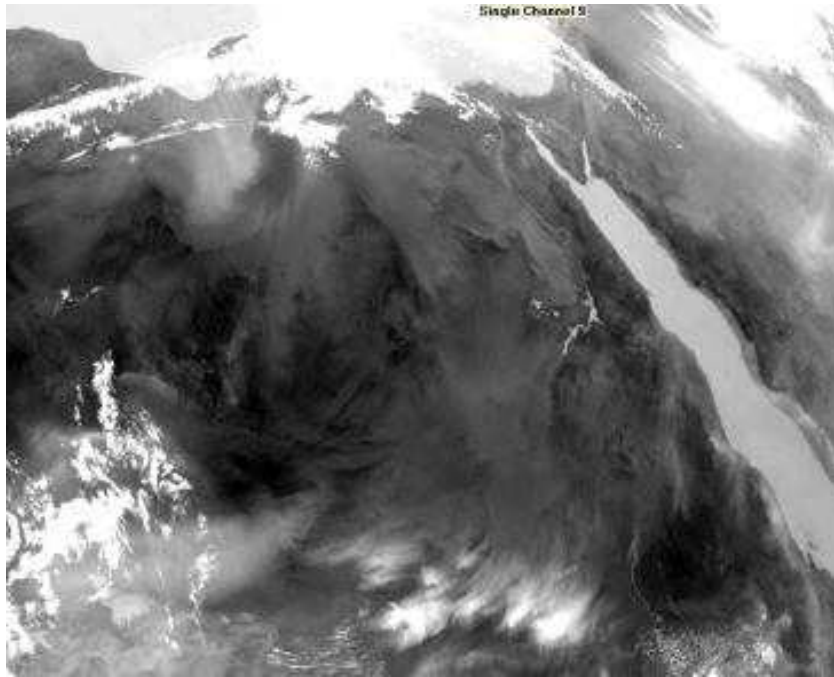
Single channel comparison and Ch9-Ch10 difference (*upper right*)



9

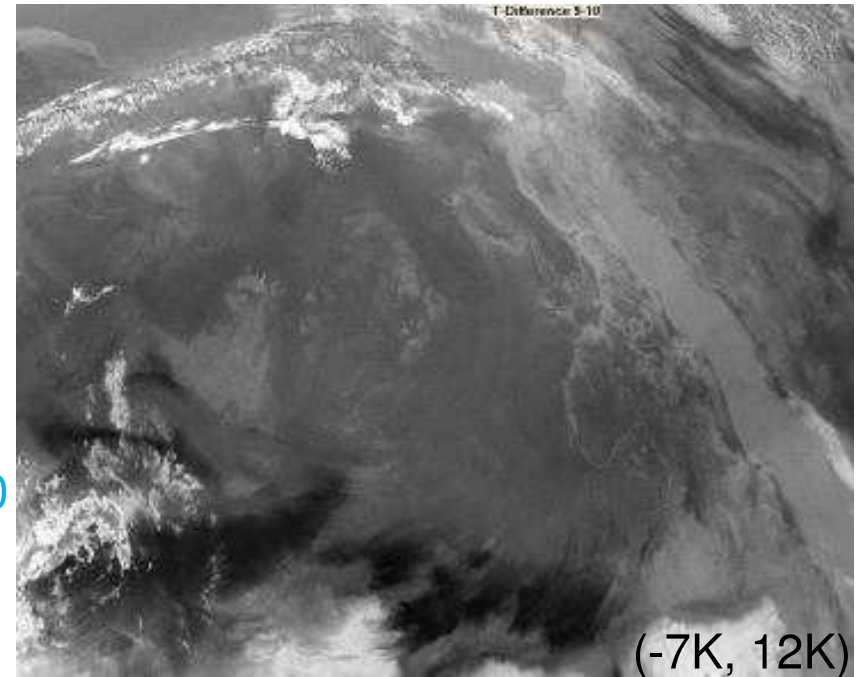
10





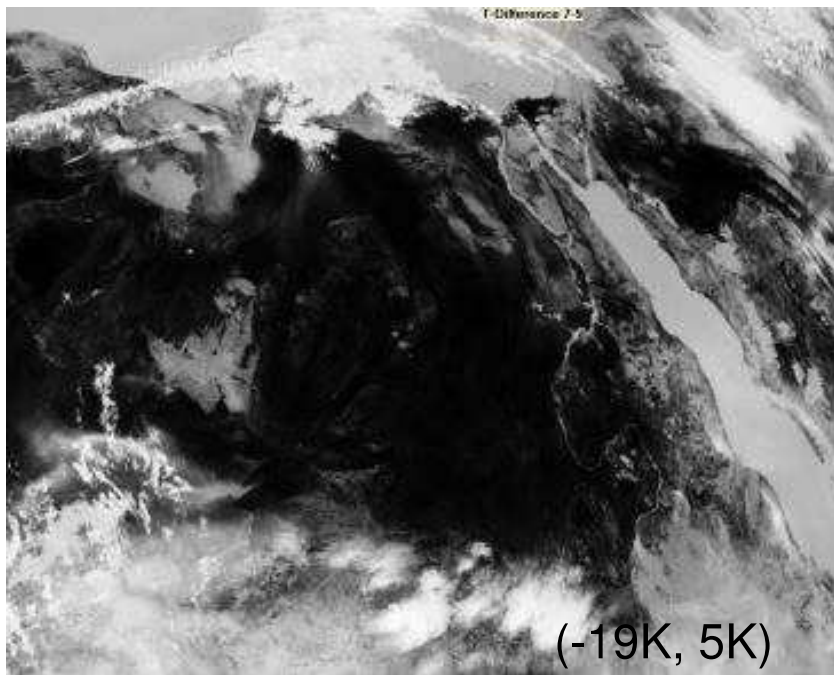
9

9-10



(-7K, 12K)

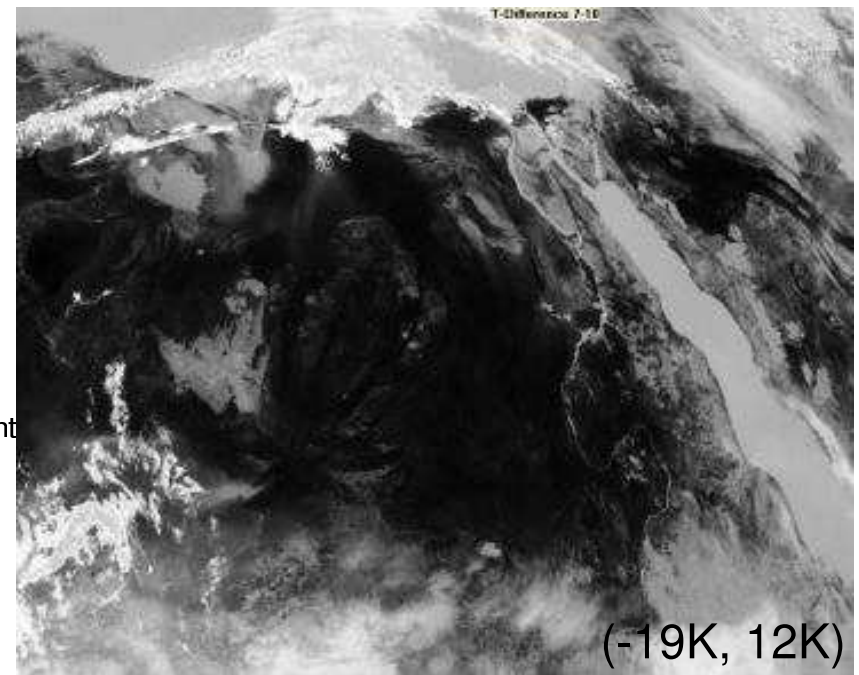
Ch9 (*upper left*) and three independent differences



7-9

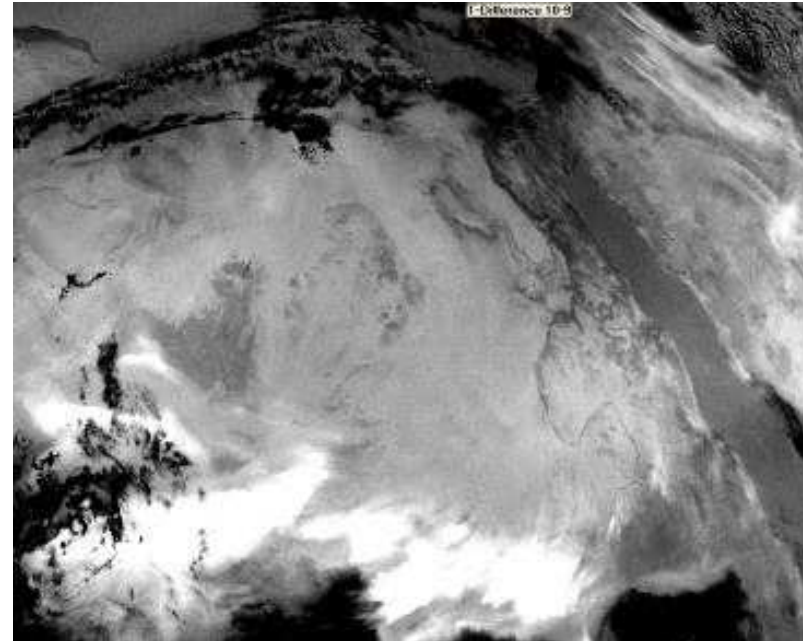
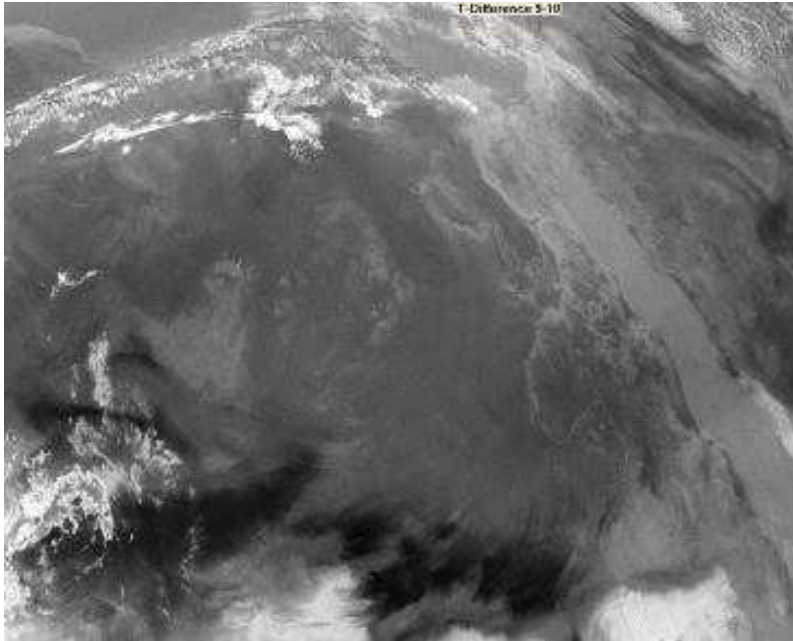
(-19K, 5K)

7-10
redundant



(-19K, 12K)

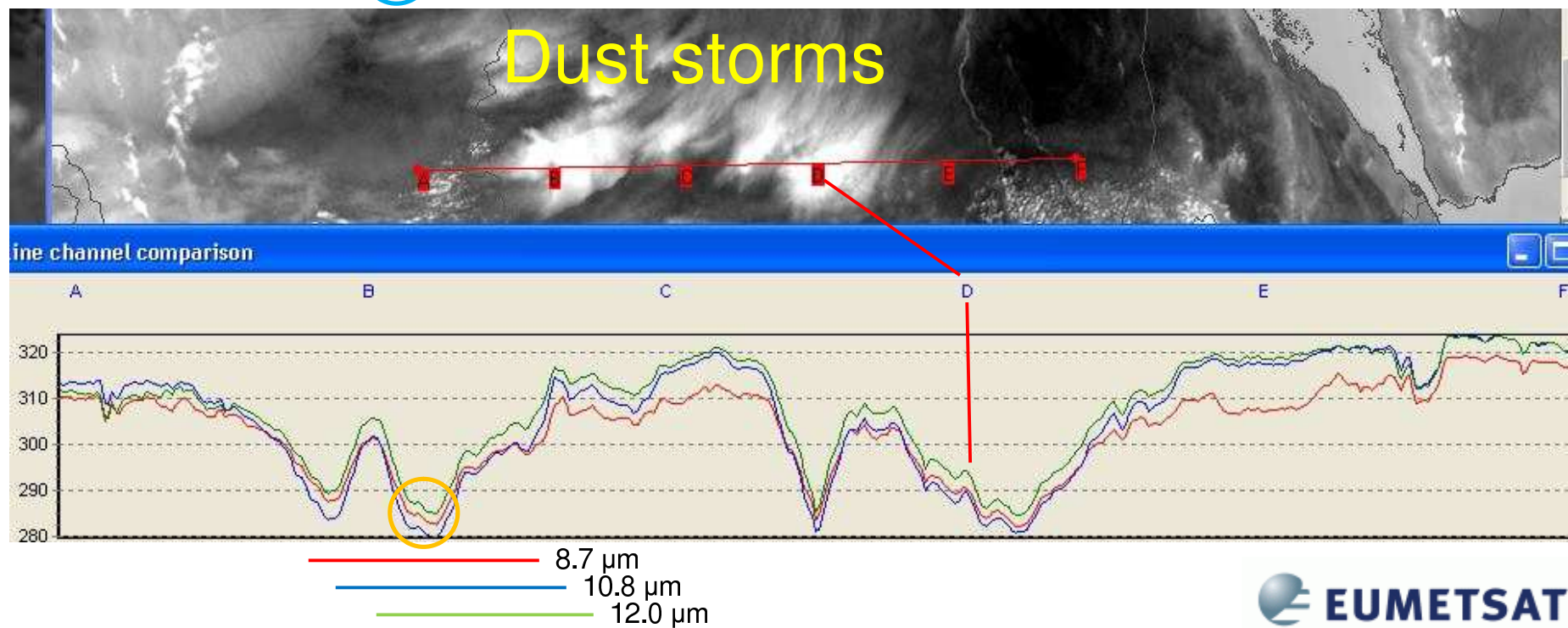
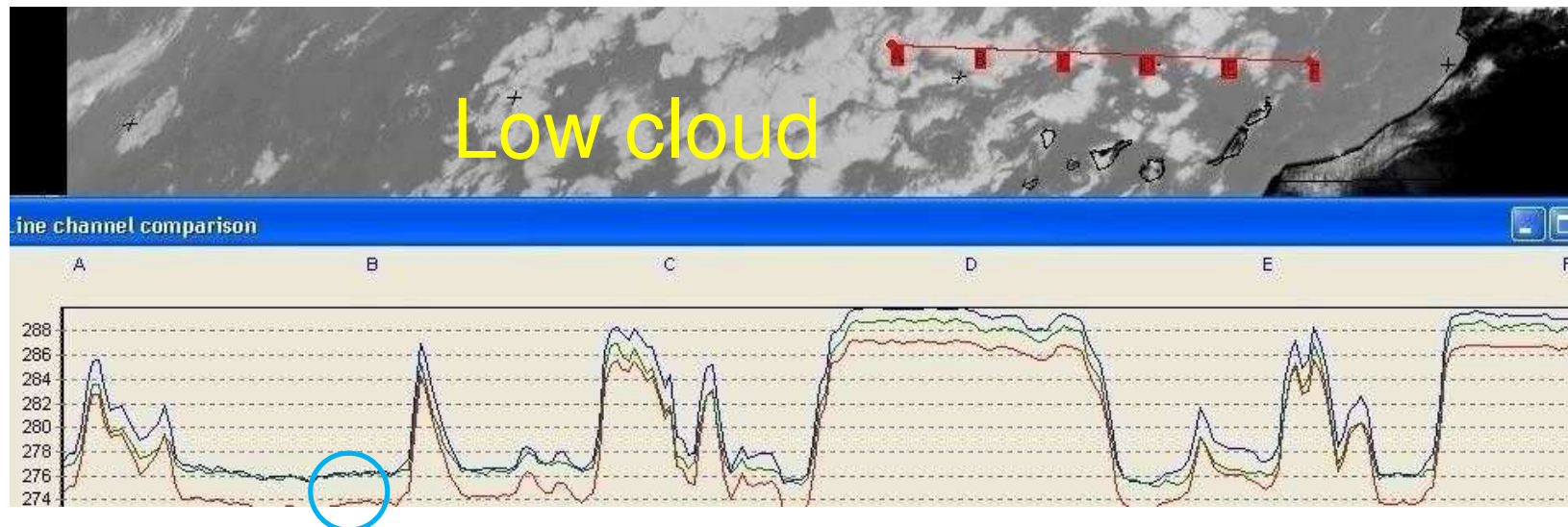
Difference image: Ch9-Ch10 or Ch10-Ch9 ?



Direct Ch9-Ch10 versus reversed difference, Ch10-Ch9

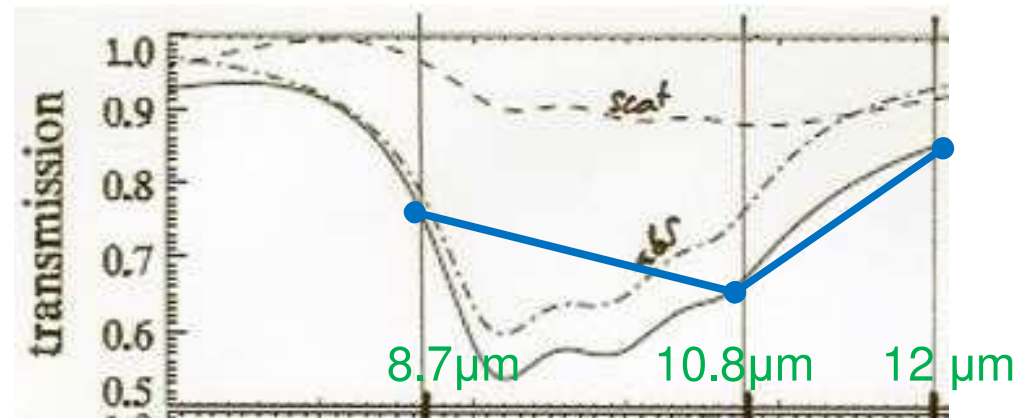
- If dust is the target, reversed difference is more of an attention grabber
- For colour composites, this difference will show better reversed as colour

Comparison of water cloud and dust in the IR window



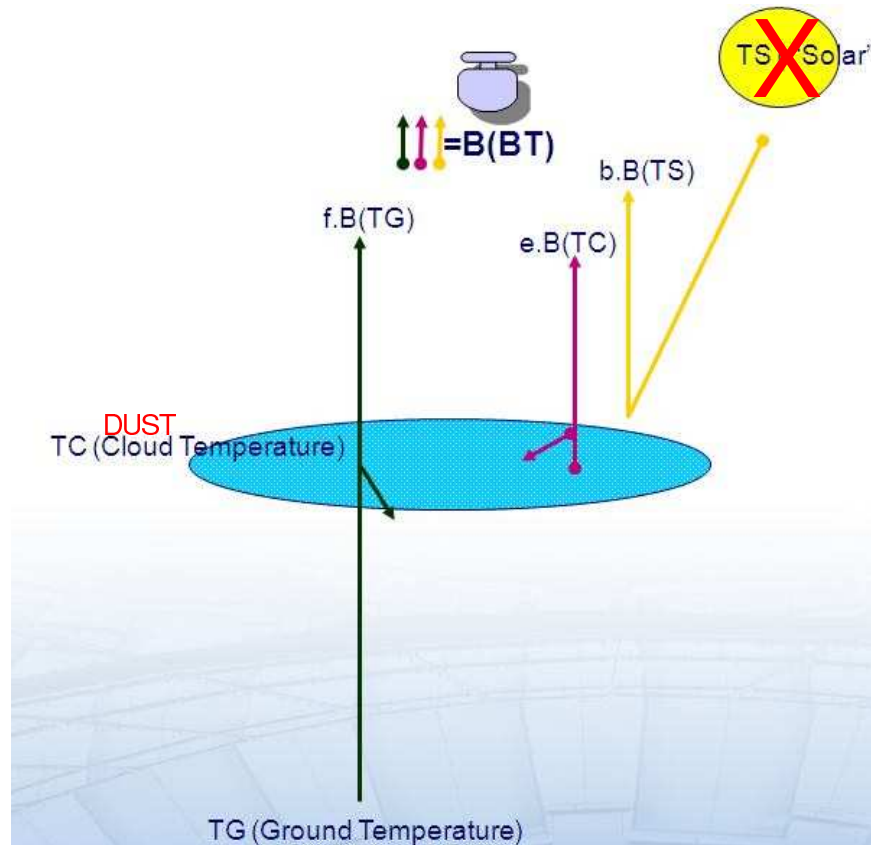
Signal reduction by dust storms (typical, dry atmosphere)

Signal reduction (K) at channel	8.7 μm	10.8 μm	12.0 μm
Water vapour absorption	-2	0	-1
Dust absorption	-5	-6	-2
Dust scattering	-2	-4	-3
Dust extinction (abs + scat)	-7	-10	-5
Desert/clay emissivity	-10	-1	0
Ocean emissivity	0	0	0



Transmission through 3 μm -ash
for wavelengths in the atmospheric window

Scattering and absorption by a dust layer



- No solar contribution in the spectral range above $8\mu\text{m}$

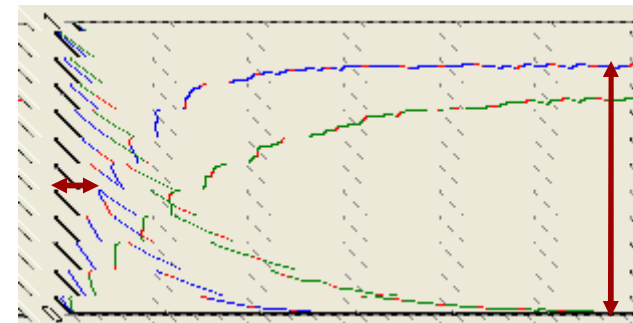
- But **scattering** !

- $f + e < 1$

- For thick dust layers $e < 1$

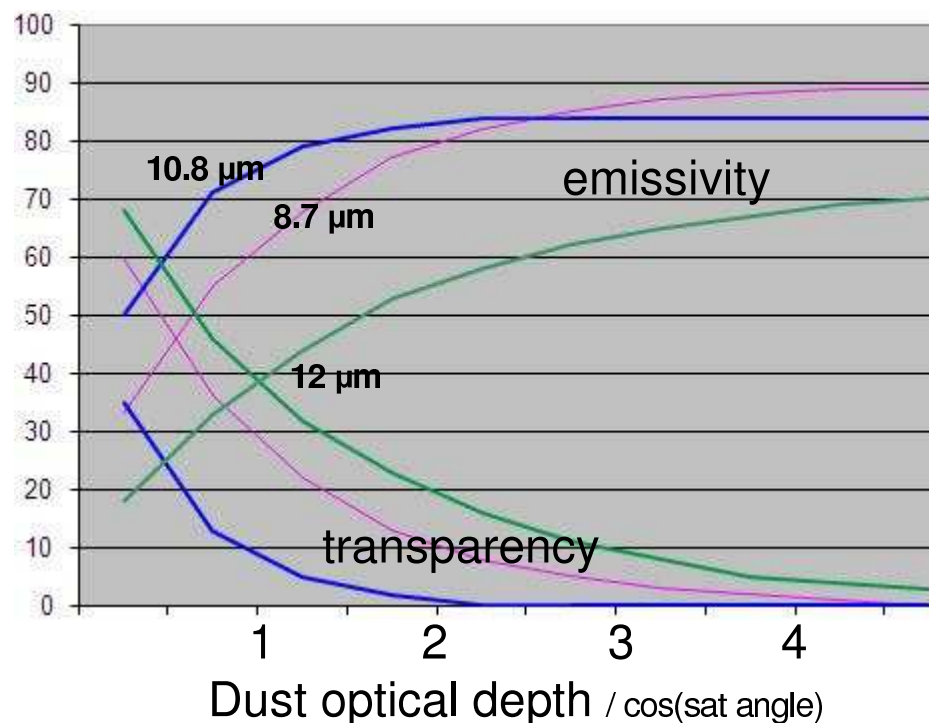
- **Thick layer emissivity** dependent on absorption/scattering efficiencies **ratio**

- **Contributing layer** inversely to absorption+scattering (extinction) **sum**



optical depth →

Emissivity and transparency



➤ **EMISSIONITY**: The emissivity of the dust layer grows with the dust optical depth up to a saturation value. It is higher for 10.8μm than for 12.0μm.

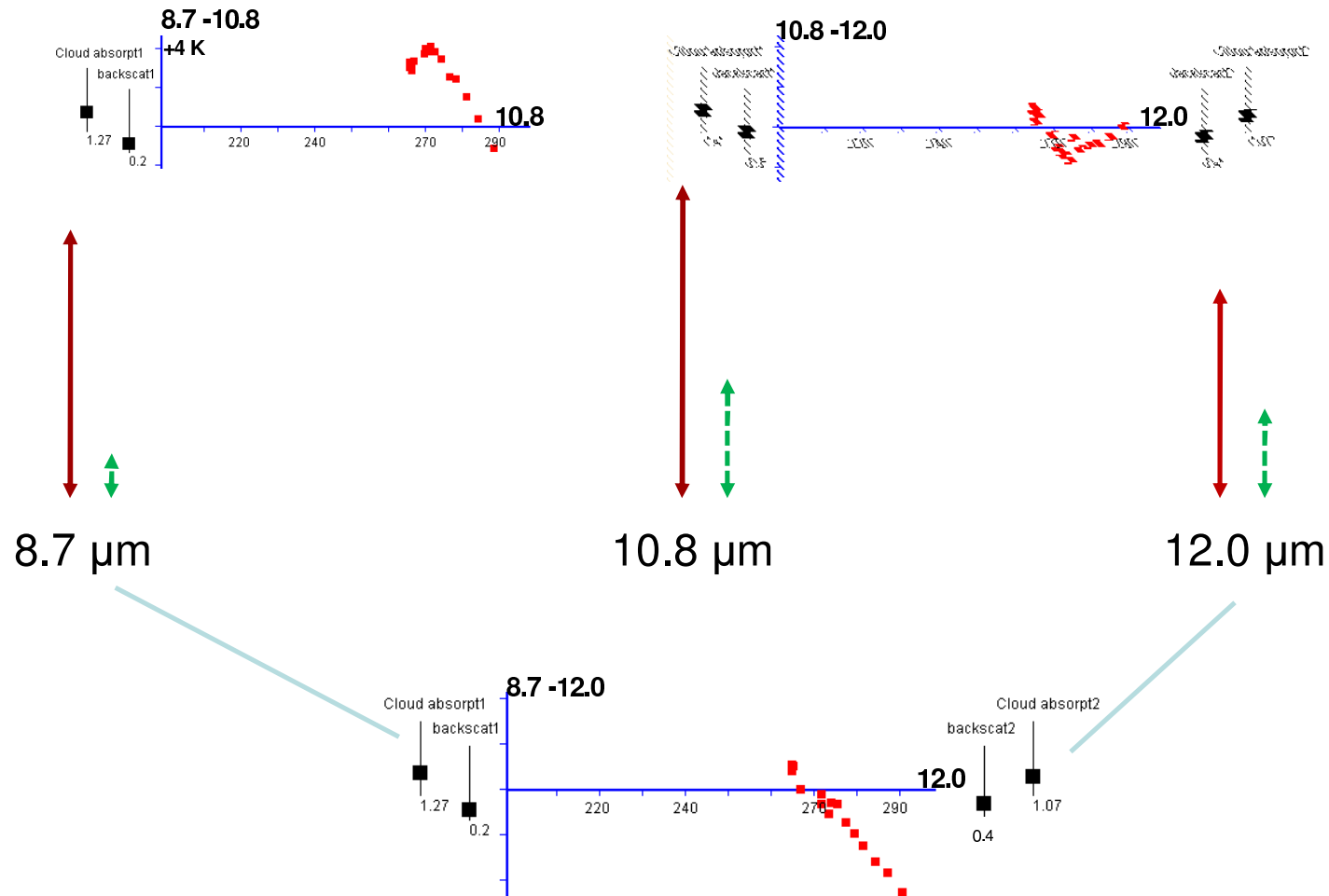
➤ **TRANSPARENCY**: The transparency supplies warm signal from the ground, making the signal stronger at 12.0μm than at 10.8μm (opposite to warm bias)

➤ **CONTRIBUTION LAYER**: For deep dust (no ground contribution) the upper layer (supplying the signal) is thicker (and a few K warmer) at 12.0μm

thin

thick

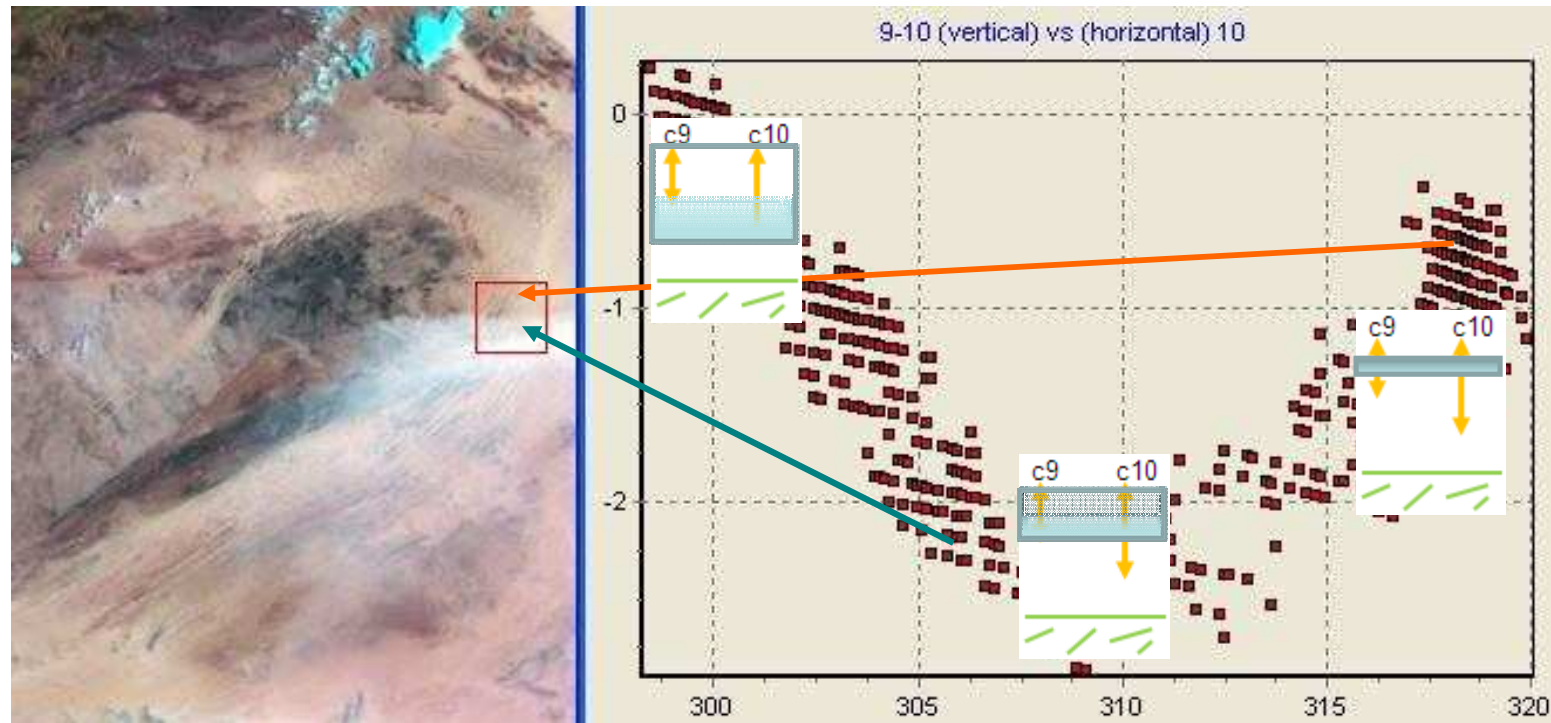
Absorption and scattering efficiencies



Red dots indicate dust optical depth in the range from 0 to 3 units

8.7 μm is second in absorption, 12.0 μm is second in scattering

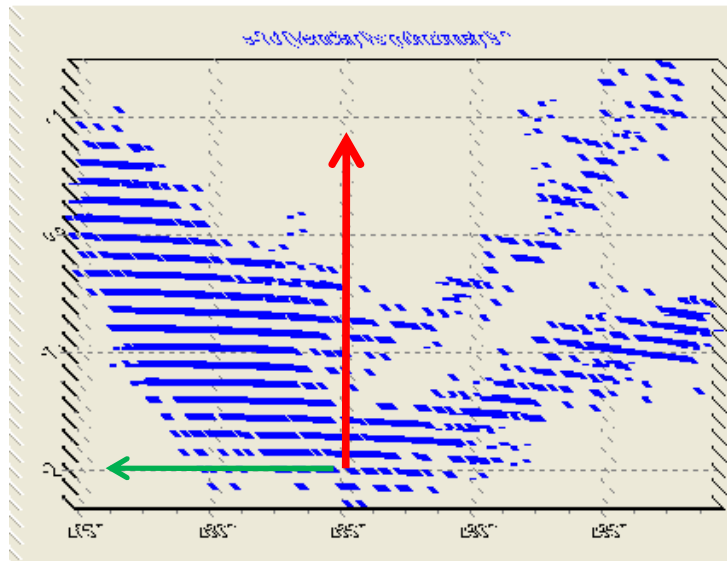
“Reversed arc” for dust scenes: Ch9-Ch10 versus Ch10



MSG Natural (solar) RGB composite 4-July-2003 10:00 UTC

- ❑ $10.8\mu\text{m}$ radiation is more absorbed and more backscattered by dust than $12.0\mu\text{m}$
- ❑ For dust or ash, arc is inverted due to the thinner contribution layer at $10.8\mu\text{m}$
- ❑ $10.8\mu\text{m}$ channel shows higher BT for thick dust due to higher emissivity

Positive Ch9-Ch10 for thick dust cloud

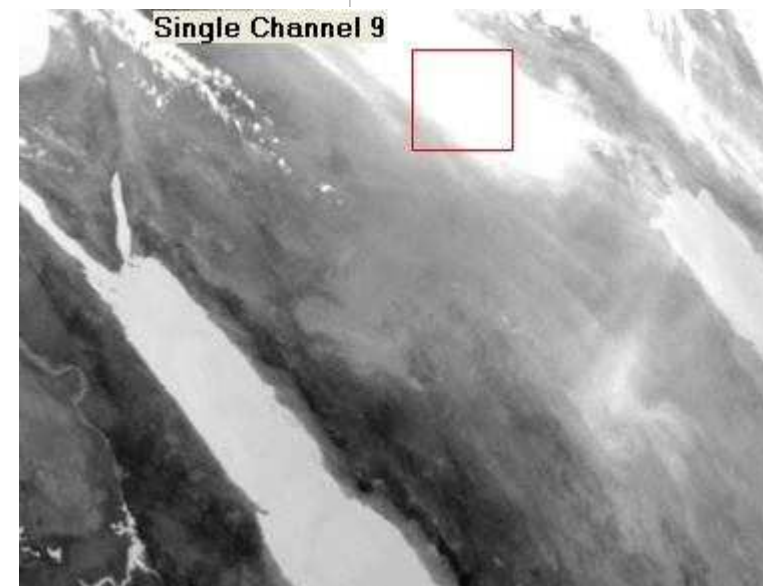
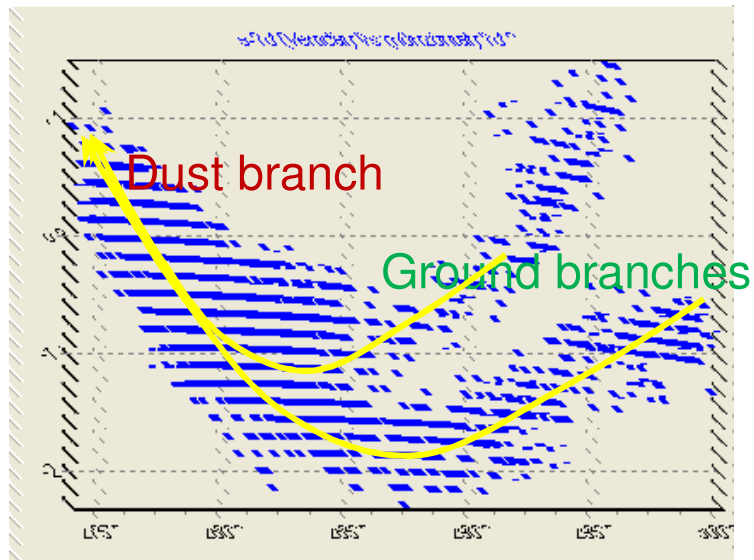
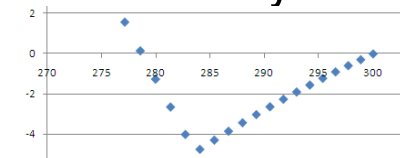


❑ Dust tends to higher levels far from the source, decreasing in **particle size**

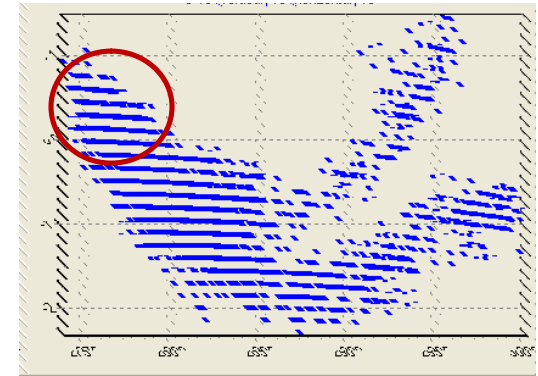
❑ **Higher** dust directly correlates with **thicker** dust until the ground signal vanishes on Ch10

❑ 10K height ~ 3K 'thick' :classify dust storms

❑ Not really an arc:



Exercise: plots 9-10 versus 10



Why is the brightness temperature difference Ch9-Ch10 turning positive for very thick dust layers?

- S E ☐ Dust shows a higher emissivity at Ch9 than at Ch10
- S E ☐ Water vapour condenses on dust and favours Ch9 emissivity over Ch10 emissivity
- S E ☐ Dust has a scattering component, higher at Ch9 than at Ch10
- S E ☐ Dust Ch10 signal comes from a thicker (and therefore warmer) layer
- S E ☐ Thicker layers normally reach higher in the atmosphere, and the signal decreases more on Ch10 for lower temperatures

Retrieval solutions: numerical and graphical

Pixel-based / numerical:

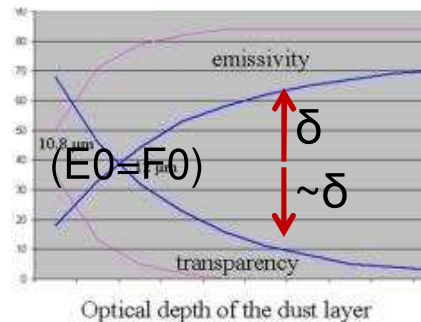
- Big errors in apriori information for variational assimilation
- Doubtful assumptions, gaussian-normal distribution for dust!

Context-based / graphical:

- Even more assumptions (uniform dust type in area, height independence for size, uniform ground, ..)
- Intuitive
- Precise results

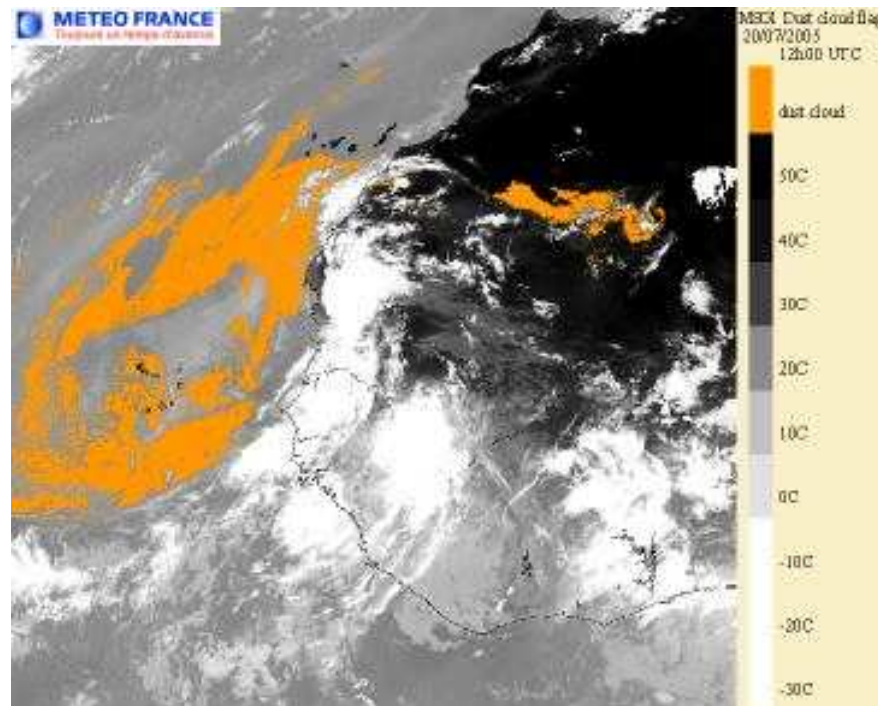
Numerical retrieval solutions

- Signal = Dust signal (T) + Ground signal (T0)
- $B(BT) = (E0 + \delta) B(T) + (F0 - \delta) B(T0)$
- δ can be translated to optical depth via E-F graph



Nowcasting SAF dust flag

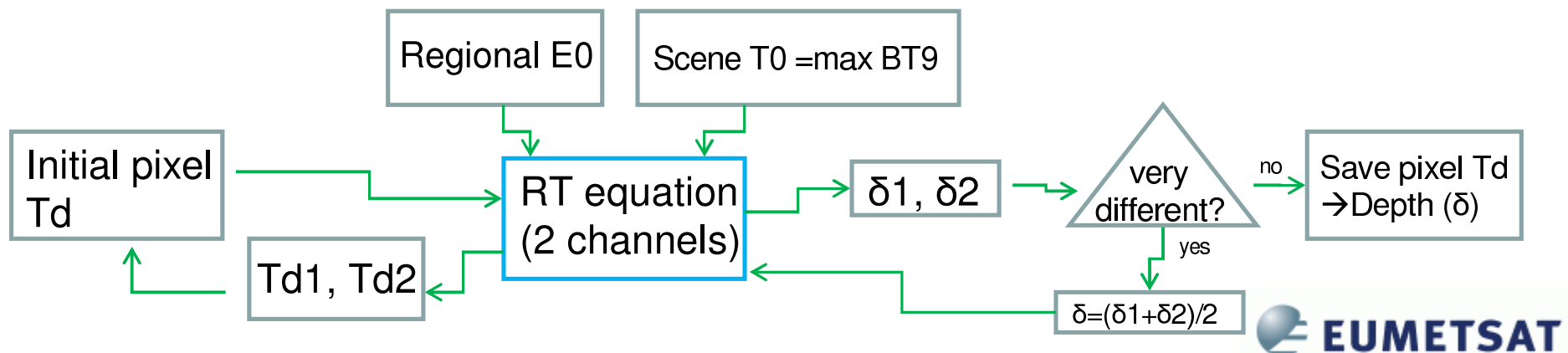
- *For the ocean, day time:* R1.6/R0.6 high, T12.0-T10.8 high, SD(T10.8-T3.9) smooth
- *For the ocean, night time:* same IR, T8.7-T10.8 high
- *For continental surfaces, day time:* not cold T10.8, smooth T10.8, filters for cloud



Numerical retrievals with two inputs

– For two channels (e.g., 10.8 and 12.0 μm):

- E0 and F0 are **adjusted** in advance for a **region** (same as efficiencies)
- **T0** is estimated for the whole **scene** as maximum value for BT9
- Initial dust T = BT9 in the pixel
- Calculate δ in the two equations, compare δ_1 and δ_2
- If δ_1 and δ_2 are very different, calculate T's for average δ in the 2 equations
- Iterate till differences get **smaller than** (2%, 2K)
- Save the pixel solutions

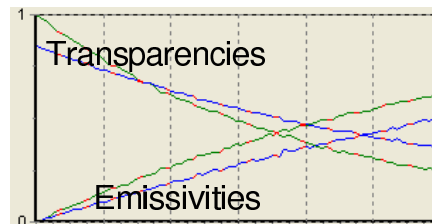


Numerical retrievals with three inputs

– For three channels (8.7, 10.8, 12.0 μm):

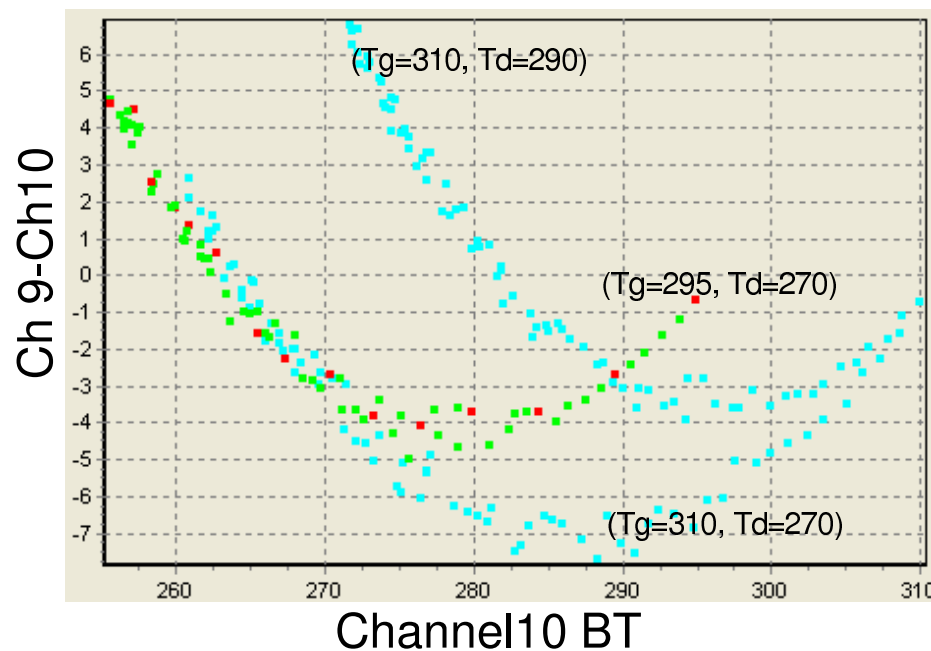
- Channel **8.7 adds** information and
- Provides scope for the retrieval of **three parameters**: (T_o , T_d , optical depth)
- However, it introduces the **uncertainty** about **ground emissivity**
- And does **not** help with the depth retrieval for **thin dust** layers
- Still, improves retrieval for **thick dust**

Emissivities and transparencies
at 8.7 μm and 10.8 μm

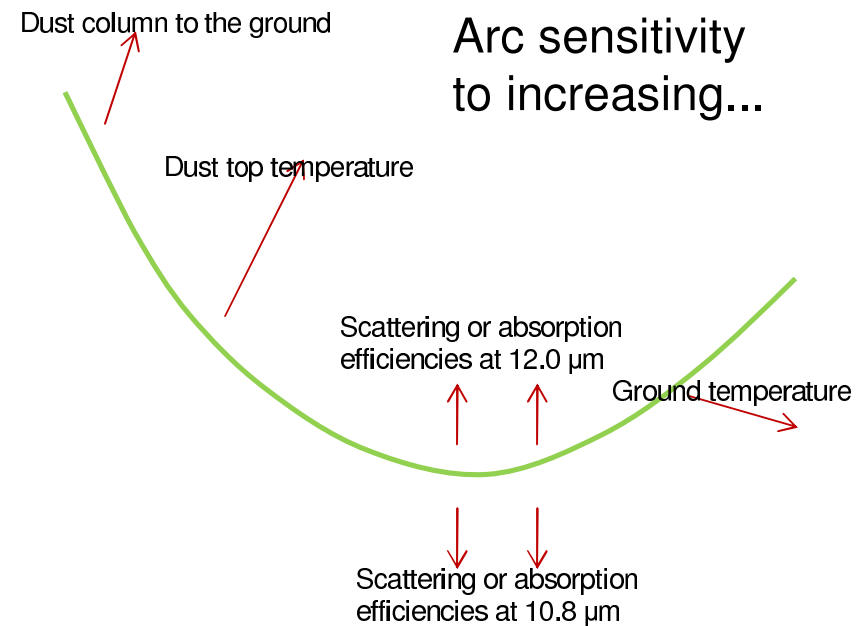


Graphical retrieval

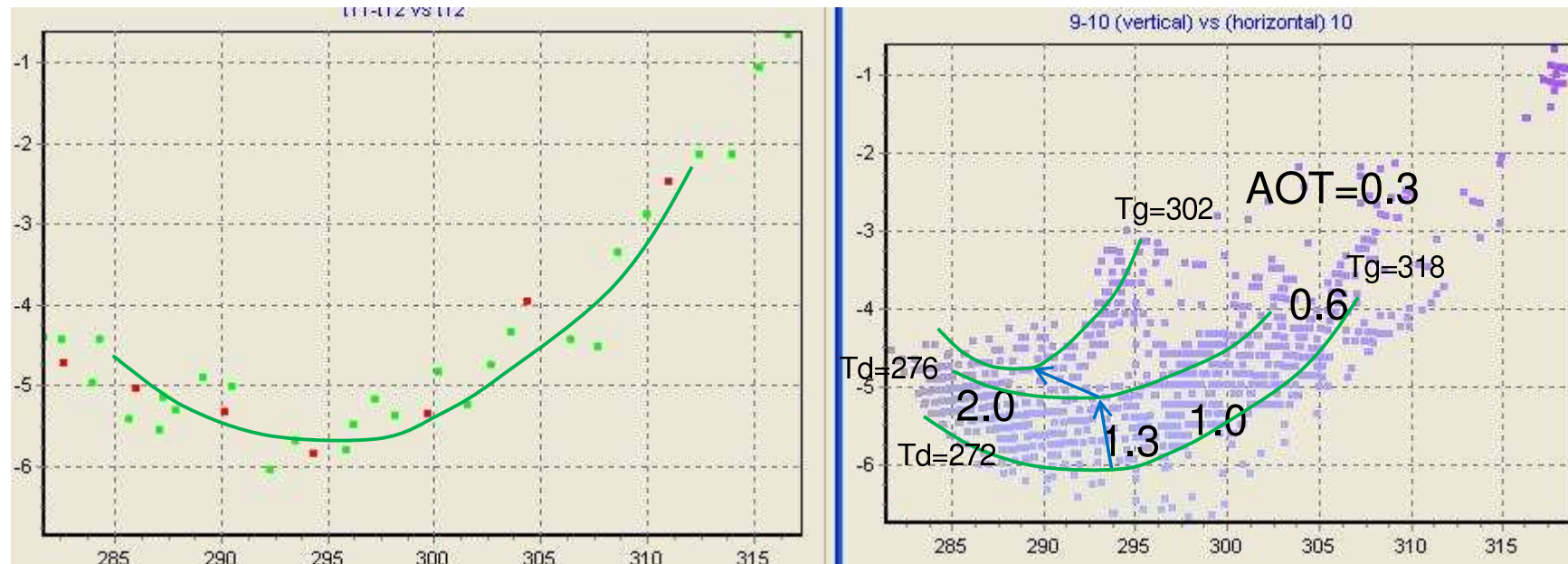
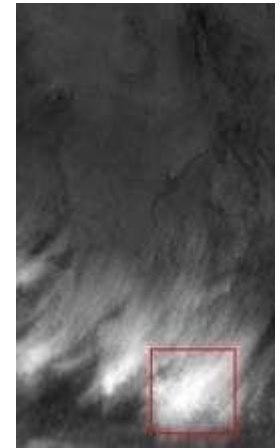
- Thick dust cloud at low level can be confused with a thin layer high above
- Uncertain reduction of the ground temperature by dust ('thermal deficit')



Green-red dotted curve for $(T_{\text{ground}}=295, T_{\text{dust}}=270)$
Cyan curves for $T_{\text{ground}}=310$, and two values of $T_{\text{dust}}=270$ and 290



Dust (Td) and ground (Tg) temperatures estimates



Real (blue dots, right h.s.) compared with simulated (left h.s. and lines) scatterograms based on $T_g=318$ $T_d=272$

$\Sigma_{11}=0.6, 0.3$ $\Sigma_{12}=0.2, 0.25$

The simulation (left h.s.) allows the thickness calibration (red dots for one-third depth units. Numbers are AOT units)

Ground thermal deficit: 30% of ground-dust difference, dust column down to 50% of that temperature difference

The upper pixels are the result of lower ground temperatures, acting as the blue arrows on the minimum.

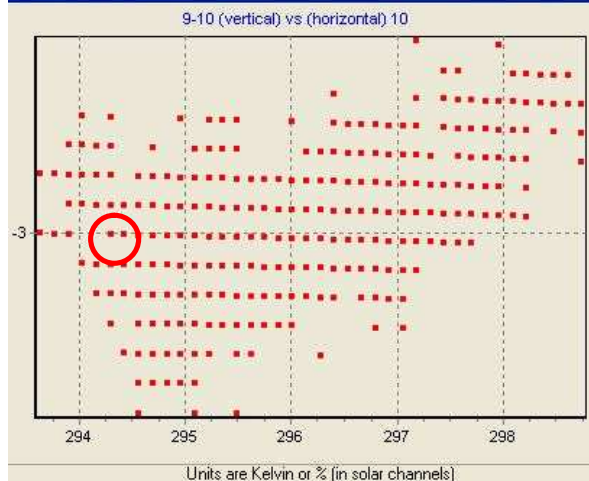
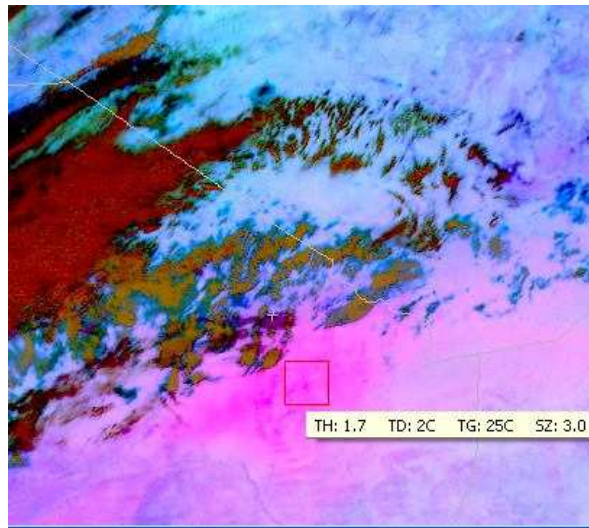


Low level dust forming a dust wall in Niamey (courtesy of E. Kploguede)

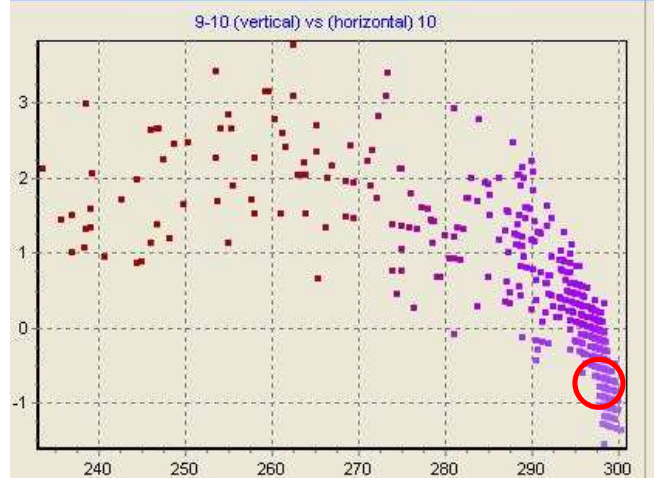
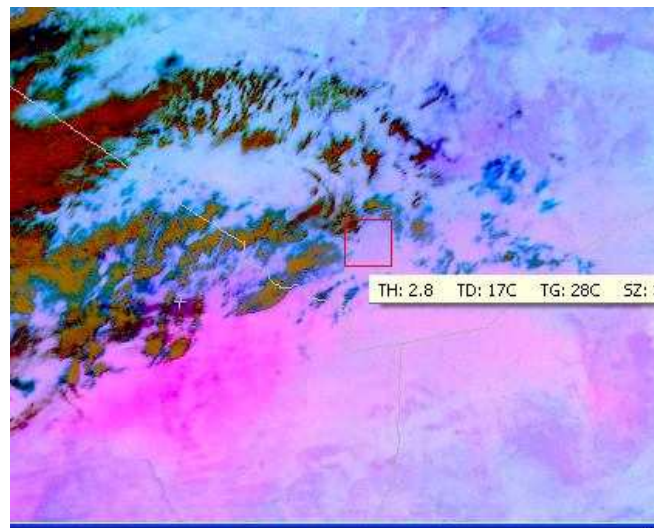
Graphical-numerical retrieval: Td, Tg, AOT, Size

- ❖ Threshold in the box (32x32 surroundings): $T_{10.8}-T_{12.0} < -1.3K$
- ❖ AOT estimate from location in the diagram: $0.14---1.3---3.5---4.8$
- ❖ Td estimate, $\max(T_{12.0}) + 5 * \min(T_{10.8}-T_{12.0})$
- ❖ Forward model to simulate 10.8 and 12.0 μ m brightness temperatures
- ❖ Corrections to retrieved parameters based on sensitivity
- ❖ Distribute corrections in up to 32 steps
- ❖ Convergence criteria (too small changes in Td and AOT estimates)
- ❖ Four result categories:
 - ❖ Dust
 - ❖ Mixed with cloud
 - ❖ Only dust traces
 - ❖ Dust-free (or low-level only, or night-time, or dry ground)

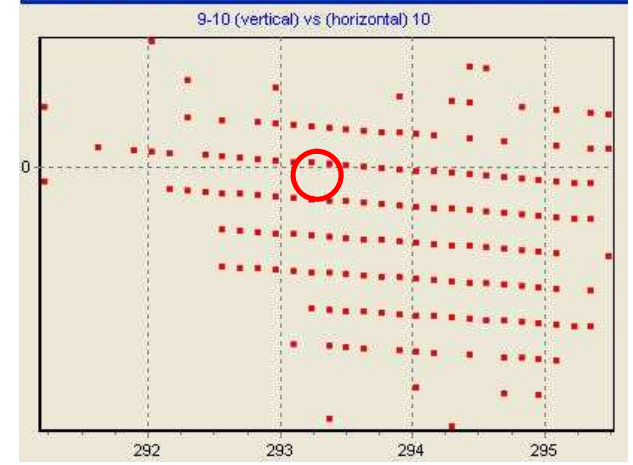
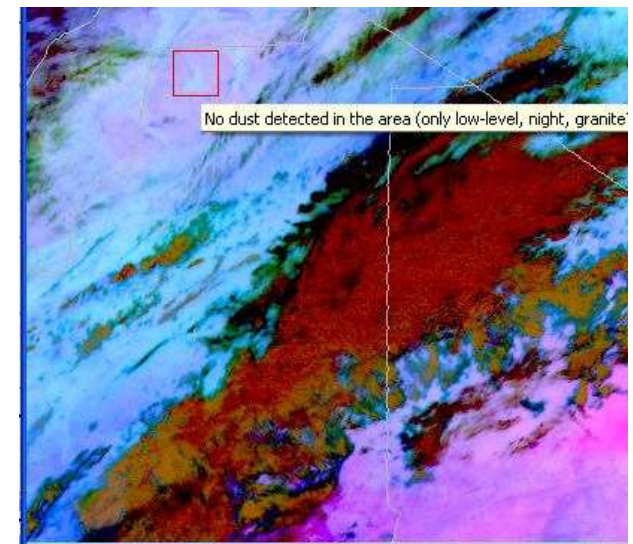
Graphical-numerical retrieval: scene determination



threshold $\text{ch9-ch10} < -1.3\text{K}$
AOT = 1.7, strong depth

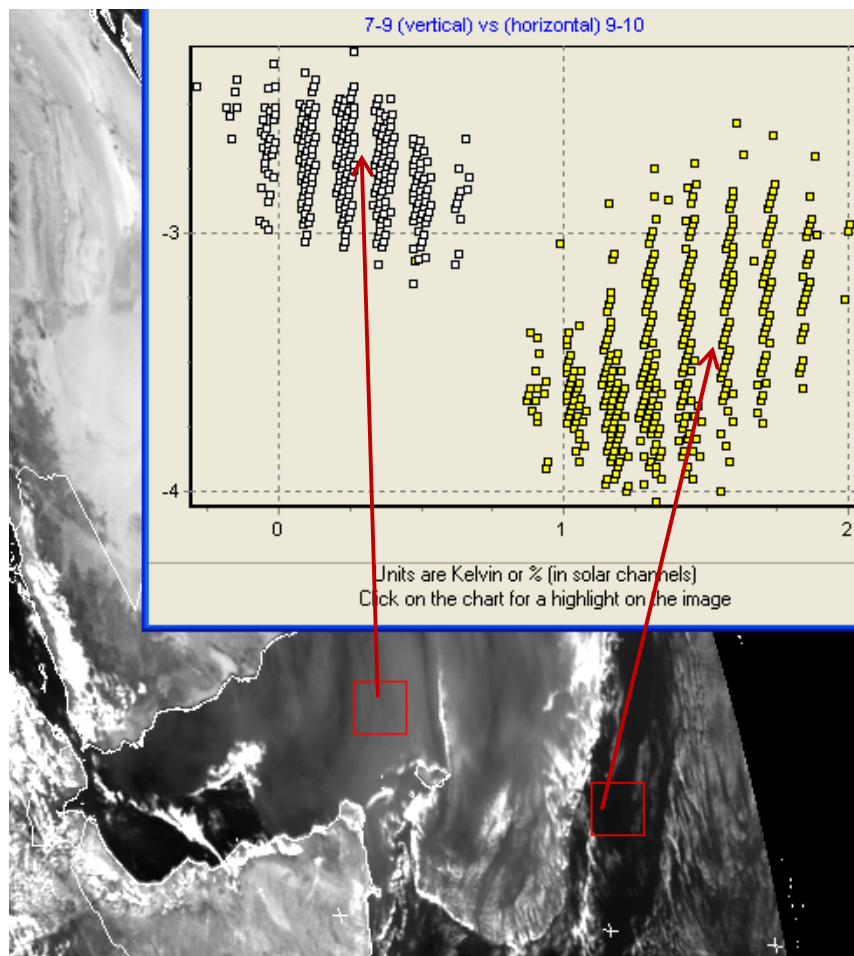


threshold $\text{ch9-ch10} < -1.3\text{K}$
AOT = 2.8, very strong depth
But suspicious of cloud



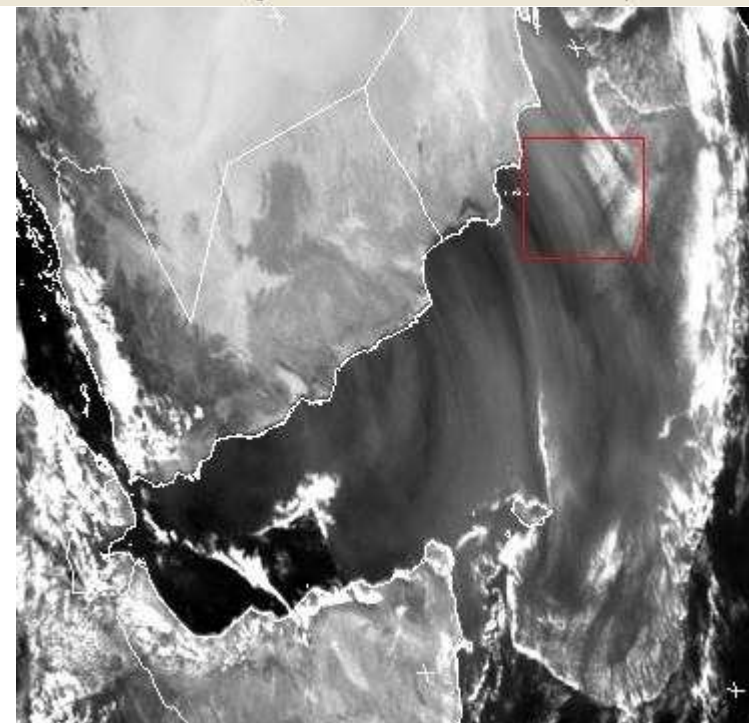
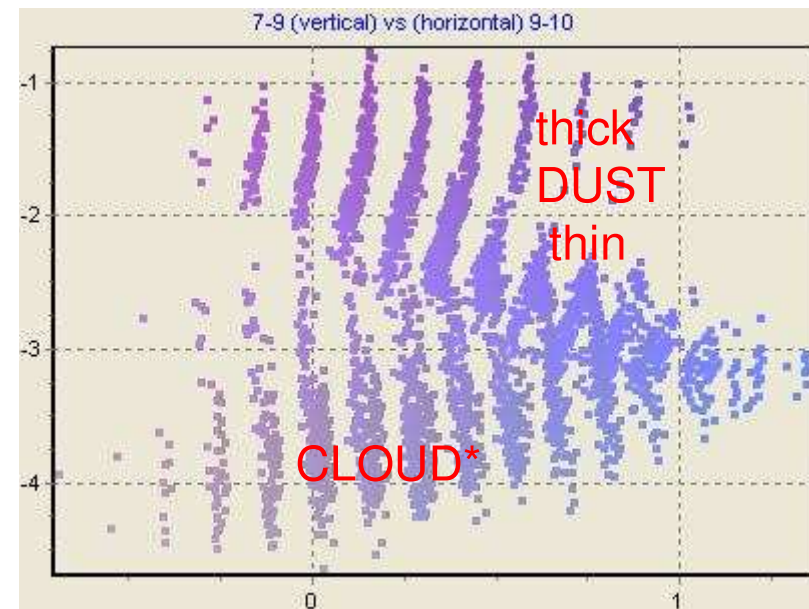
threshold NOT $< -1.3\text{K}$
AOT not calculated
Possible causes hinted

The oceans: double difference plot

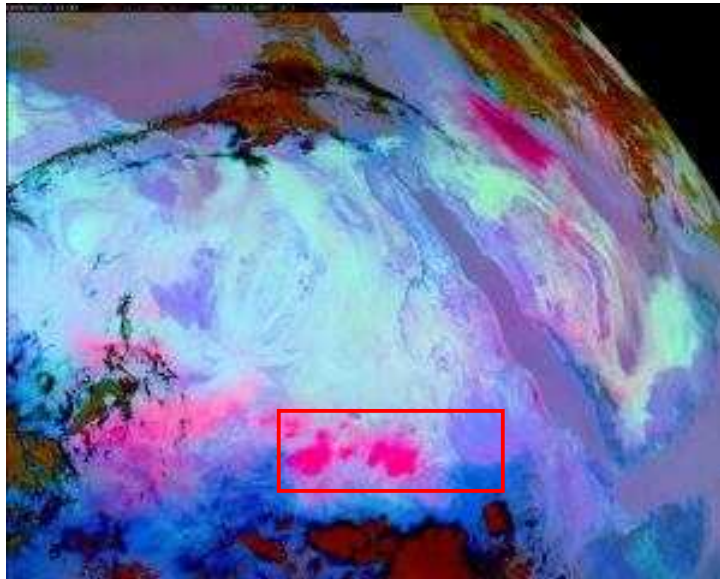


Meteosat-9 **0.6 μ m** 2008-Feb-02 10:15 UTC
Hardly shows in IR channels

*Oblique path in the area enhances the negative 7-9 difference for cloud

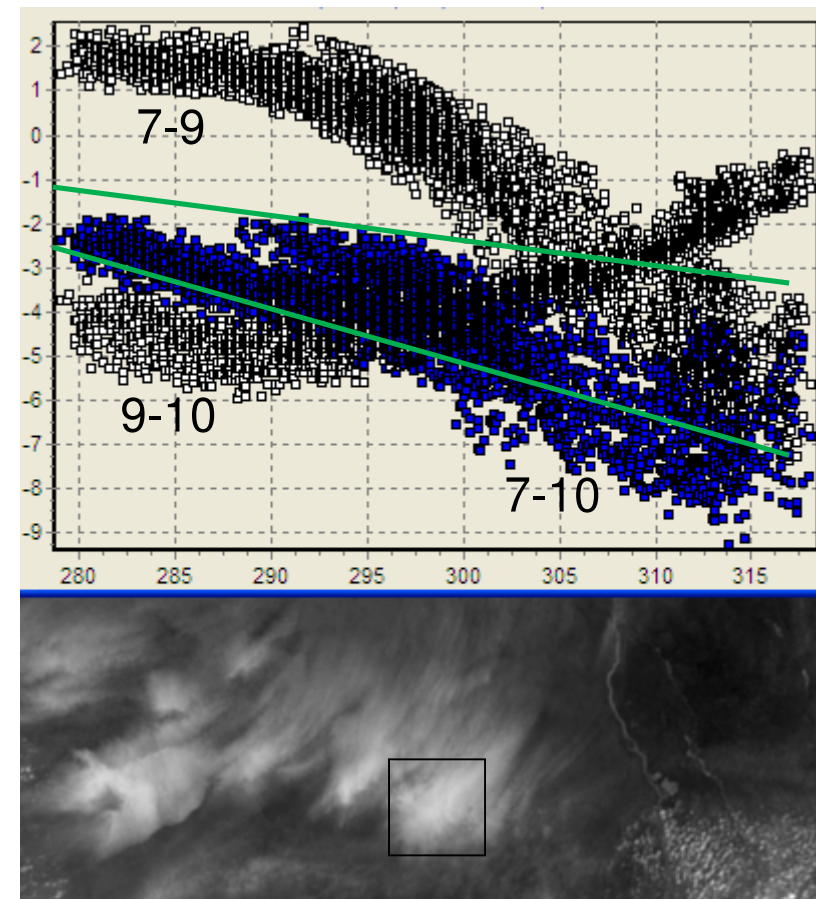


Single and double plot of differences

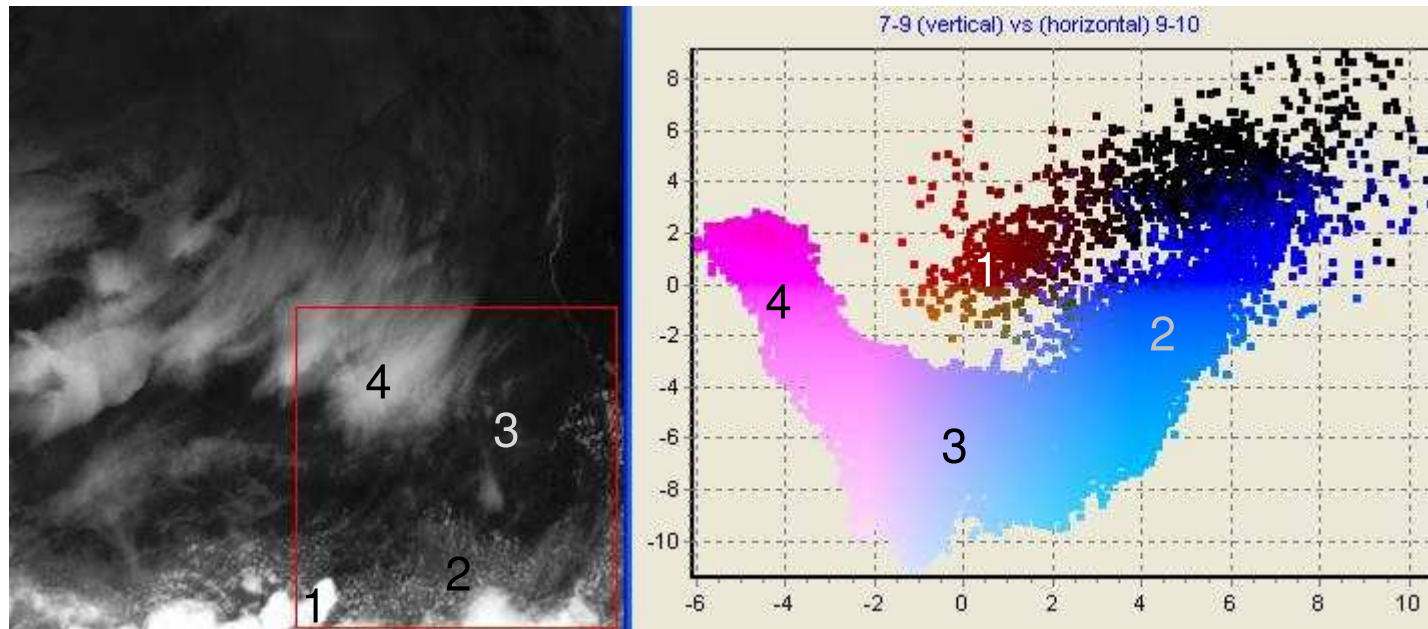


2004-05-13 13:00 UTC, 10.8 μm

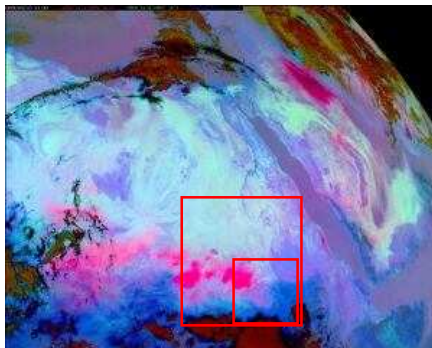
- The difference 7-10 is twice the average of the others, and rather straight in single plots for dust
- The other two differences, 7-9 and 9-10, are usually preferred



The **cloud-to-dust spiral** in the differences graph

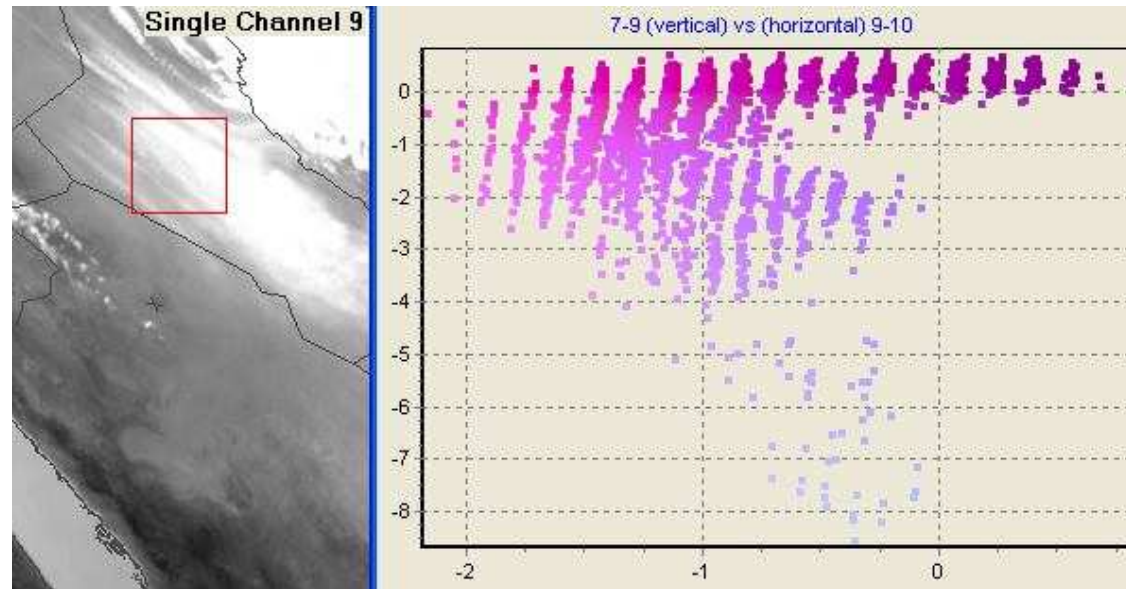


2004-05-13 13:00 UTC, 10.8 μm

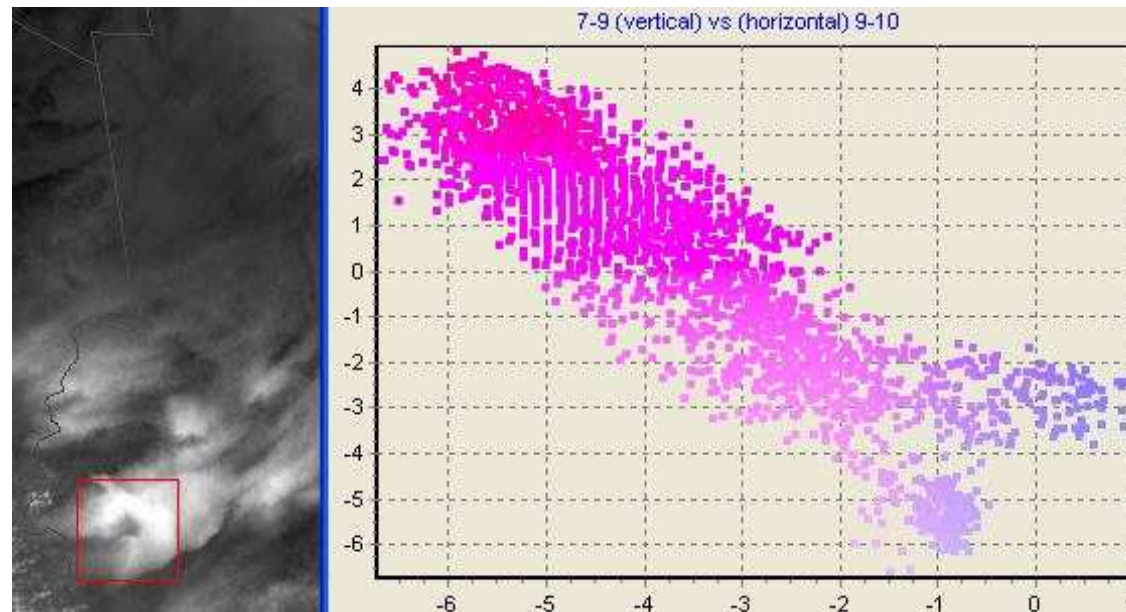


- 1: Thick high cloud
- 2: Broken low cloud
- 3: Ground, drier air towards 4
- 4: Dust cloud

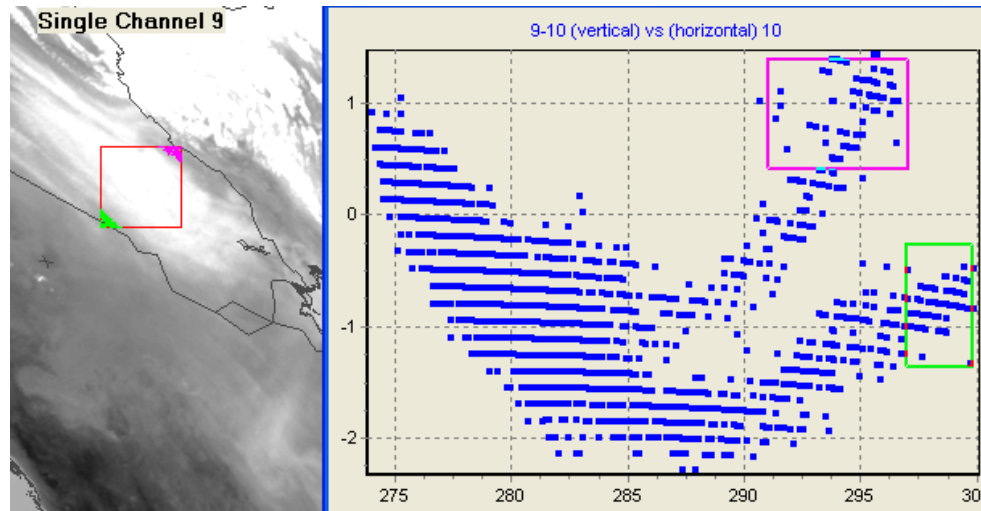
Dust branch often due to slant view of the dust storm (double diff)



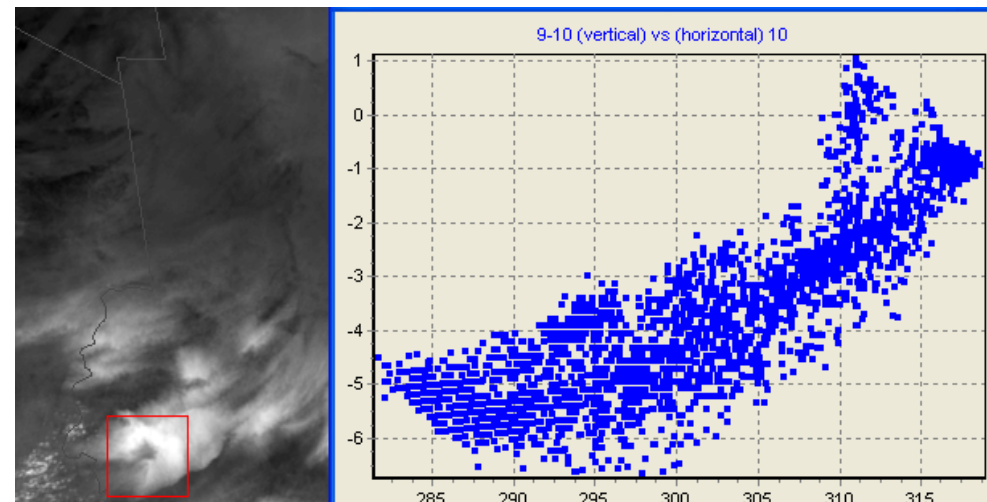
Meteosat-8 10.8 μ m 2004 May 13th 13:00



Dust branch only for intense dust storms or oblique views

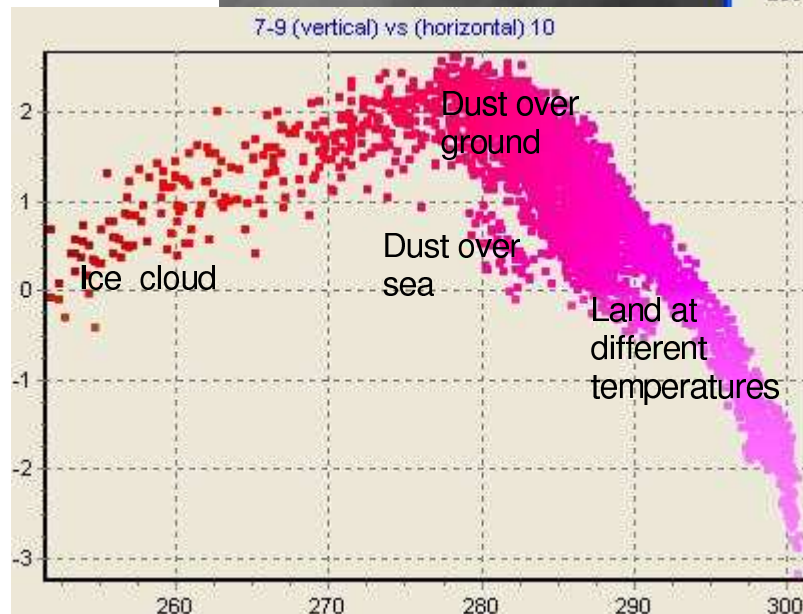
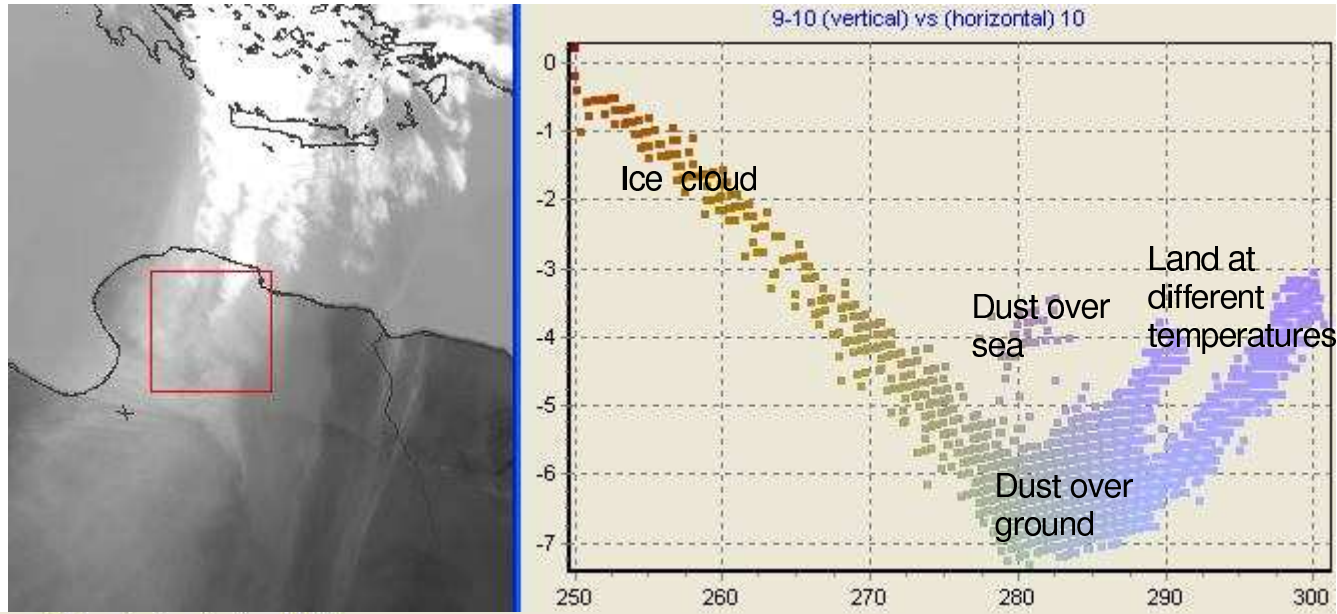


- 2004 May 13 13:00
- 9-10 turns positive for thick dust (upper graph, Saudi Arabia) due to higher dust emissivity at channel 9
- Not thick enough (less slant view) in Sudan (graph on the right), far from reaching the zero difference on the dust branch



Single plot: Ch9-Ch10 vertical, Ch10 horizontal

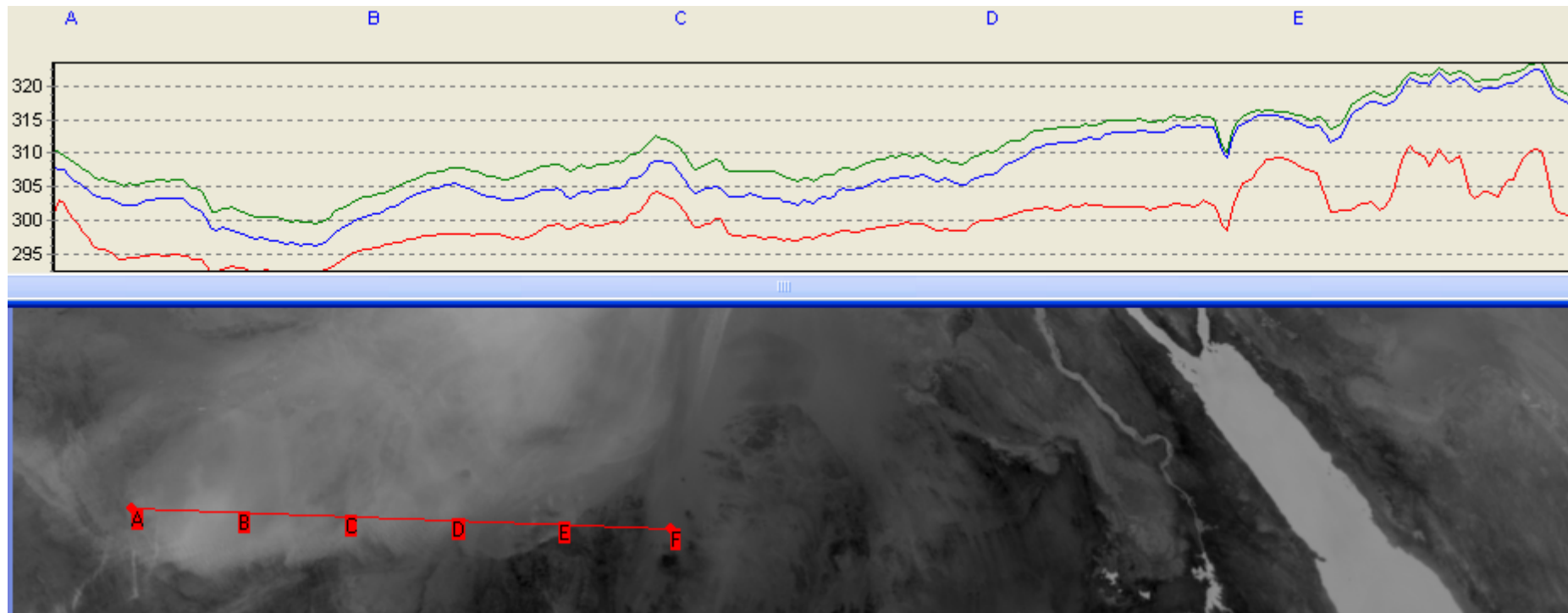
2008-03-23 11:30 UTC Meteosat Ch9



At which temperature is the cloud boundary?

- 265 K
- 275 K
- 285 K

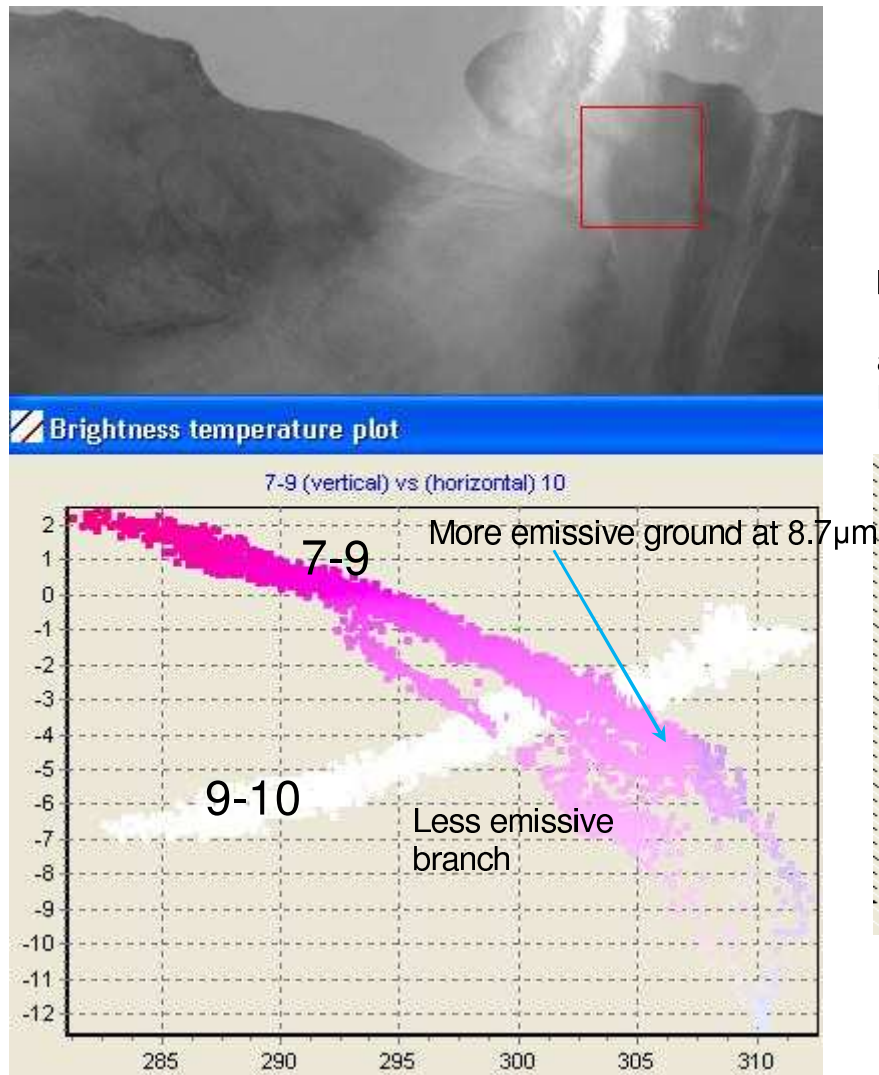
The value added by the channel 7 (at $8.7\mu\text{m}$)



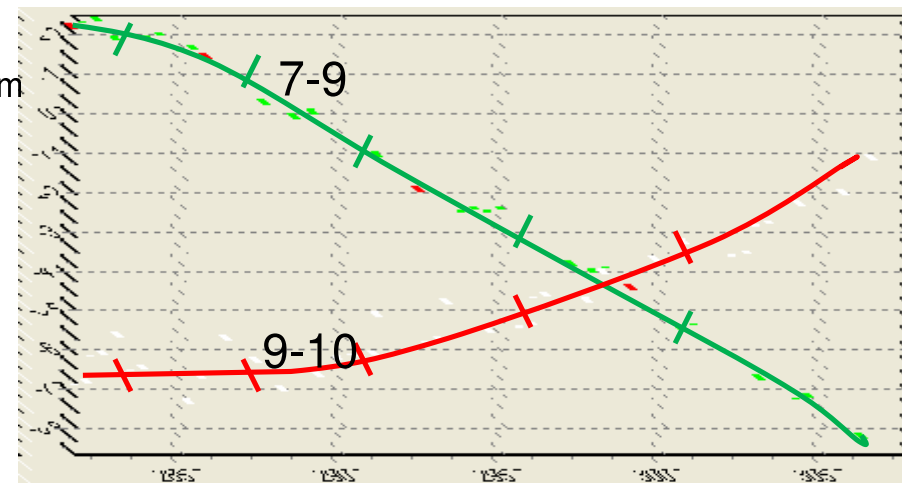
2008-March-23 11:30 UTC, Meteosat-9

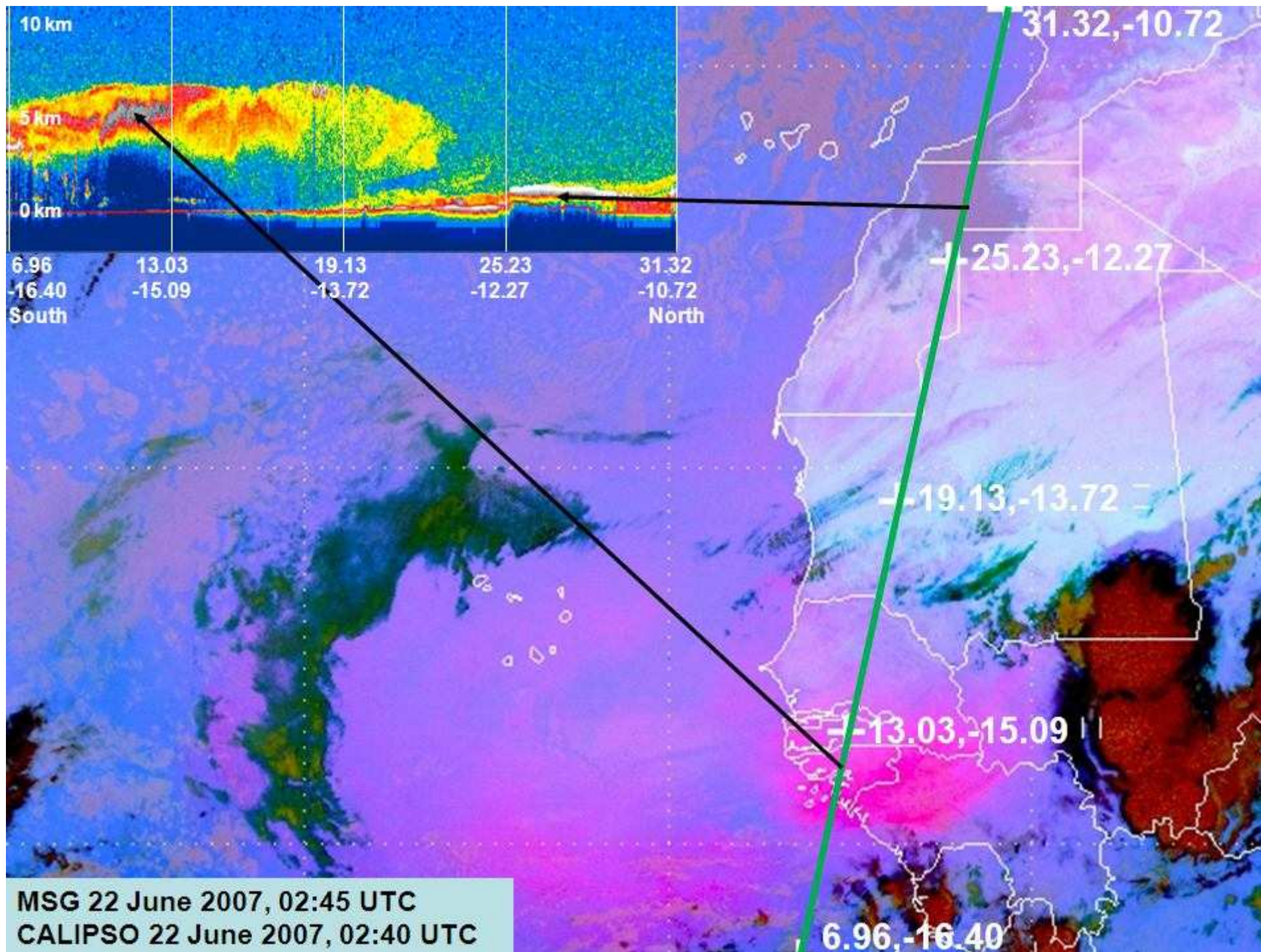
- Difference 7-9 depends on dust concentration, humidity and ground emissivity
- Difference 9-10 depends on dust concentration and humidity

The value added by the channel 7 (at $8.7\mu\text{m}$)



Real (left h.s.) compared with simulated (right h.s.) scatterograms
based on $T_g=308$ $T_d=266$
 $\Sigma_{8.7}=0.35, 0.2$ $\Sigma_{11}=0.6, 0.3$ $\Sigma_{12}=0.2, 0.25$
and ground emissivity 85% at $8.7\mu\text{m}$
Marks at optical-thickness third-units from the right ends

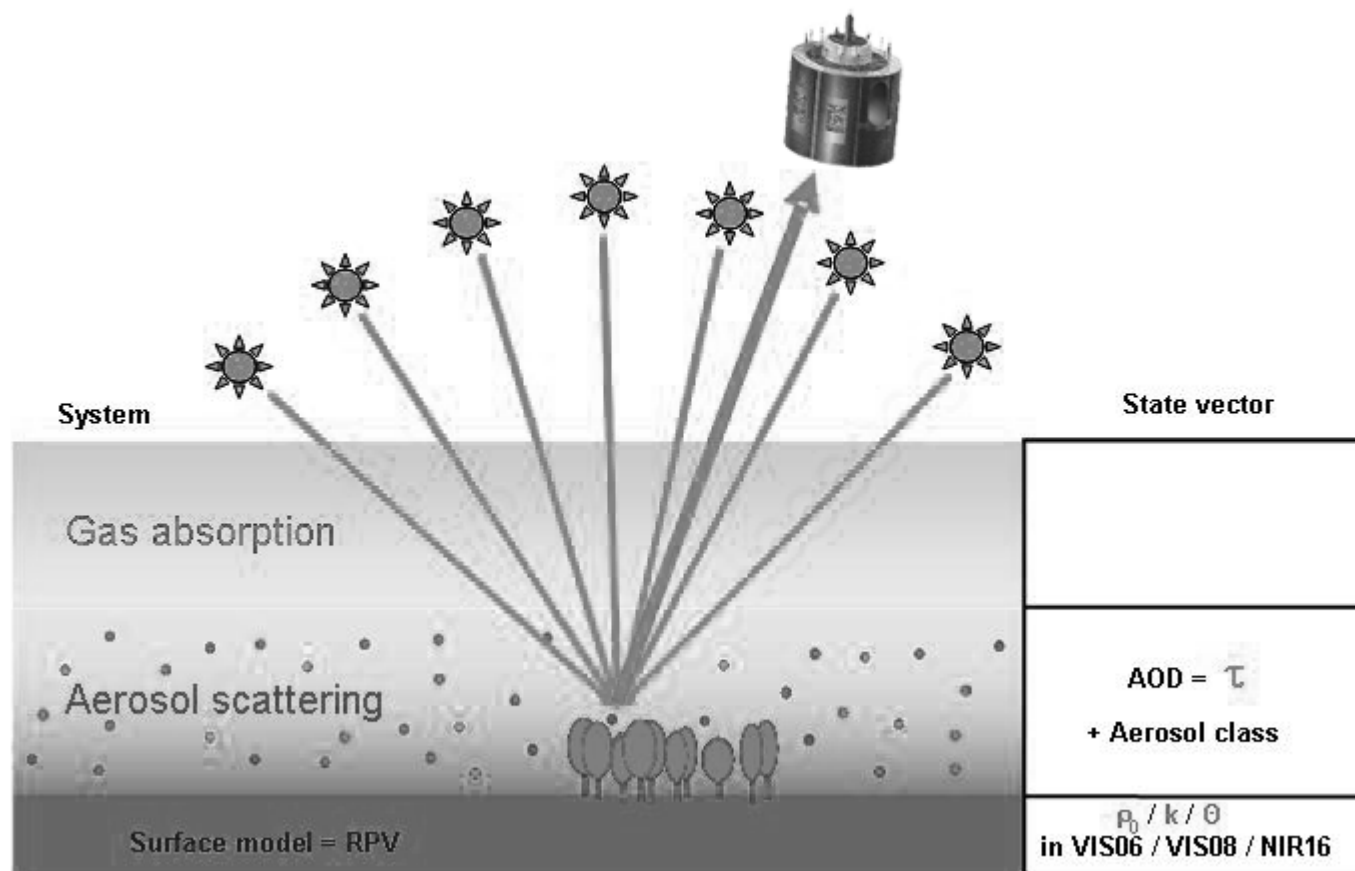




Conclusions

- Graph adjustment is more **subjective** than numerical retrieval, but it facilitates **understanding** of the radiative transfer and provides the parameters for dust storm **classification**
- Channel at **8.7 μm** provides **refinement** for *T_{dust}* and *Depth* values at the thick end of the curves
- An operational detection algorithm can be based alternatively on **thresholds** of the type: $\text{Ch9-Ch10} < -1.3\text{K}$ for pixels, plus texture indices in the area. However those thresholds are not necessary for dust

Land daily aerosol algorithm



Outlook

- A model for **dust cooling** and **height-thickness** modelling will improve the simulation of the observed radiances
- A study of the **particle size** effect on channel emissivities will improve the match between expected and real radiances
- Looking into the BT's for **dust mixed with water** will clarify the role of aerosols in cooling the atmosphere
- Coupling IR technique with existing methods for solar channel simultaneous retrieval of surface **albedo** and **aerosol optical depth**
- A comparison with results from the **visible techniques** for dust estimation will provide calibration and skill for the IR estimates, obtainable during night too

DUST RGB composite: the strength of infrared for dust detection



IR RGB composite based on
channels at 8.7, 10.8 and 12.0 μm

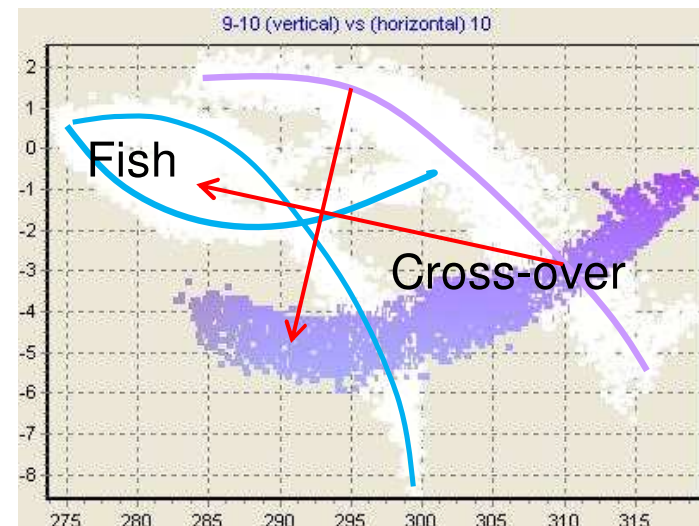


Solar RGB composite based on
channels at 1.6, 0.8 and 0.6 μm

THANKS FOR YOUR ATTENTION !

- List of referred events:

- 2004-05-13 12:00, Sudan and Saudi Arabia
- 2008-02-02 06:00, Saudi Arabia
- 2008-03-23 12:00, Libya
- 2009-03-28 18:00, Argentina



Size behaviour

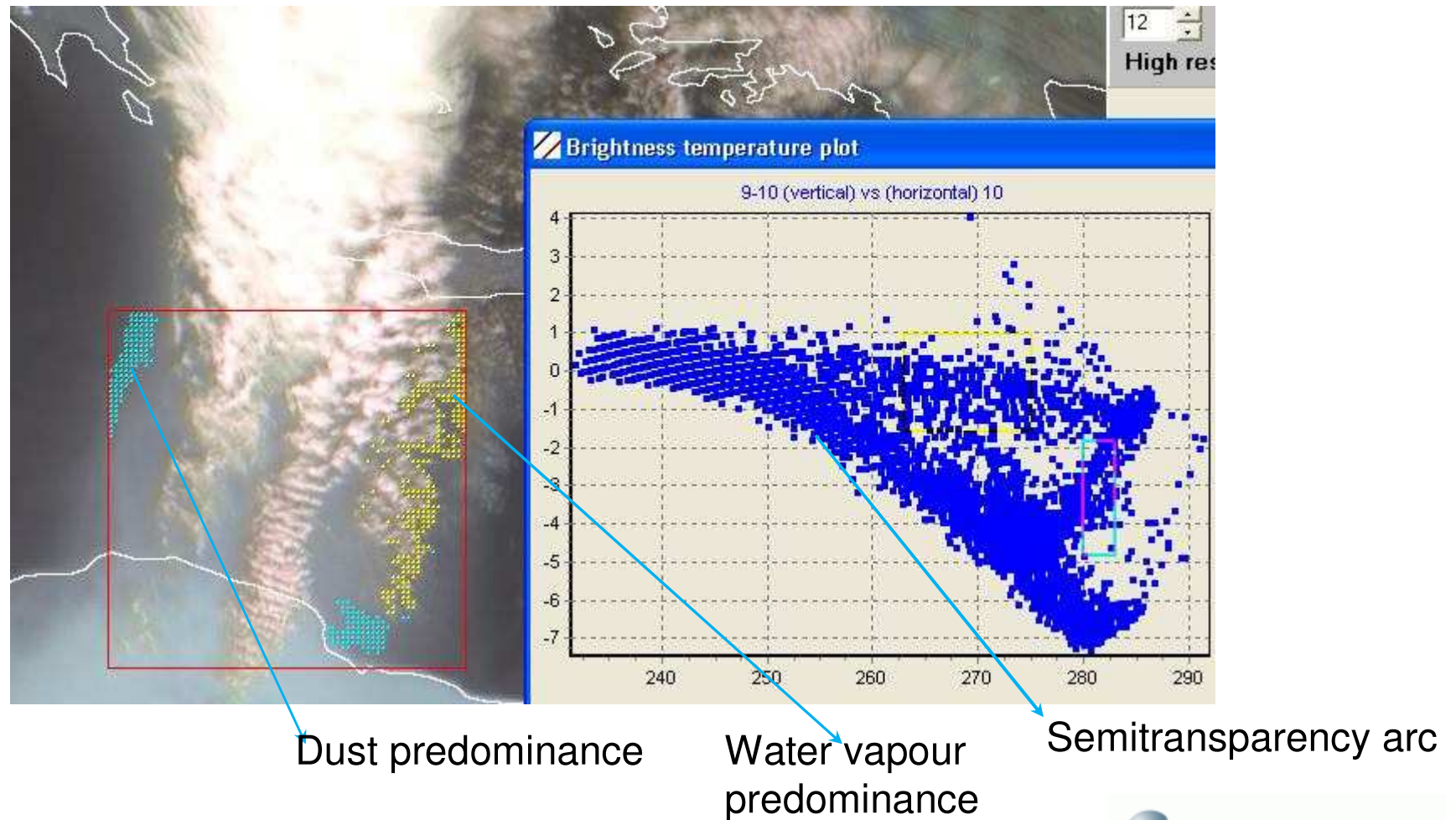
Dust Ø 2 µm	$\Sigma a = 0.1$	$\Sigma s = 0.05$
Dust Ø 20 µm	$\Sigma a = 1.2$	$\Sigma s = 0.06$
Cross-section	x1200	x120
Mass (20/2)	x1000	
Section (20/2)	/100	

Absorption is proportional to particle mass, not dependent on particulation

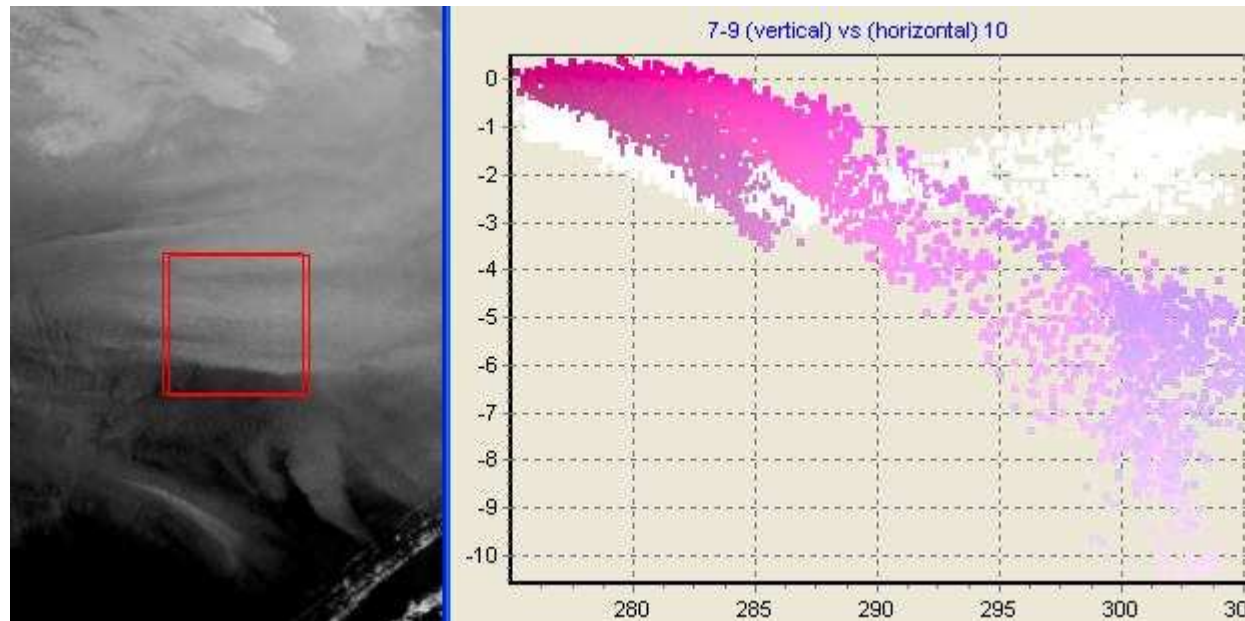
Scattering is under-proportional to mass

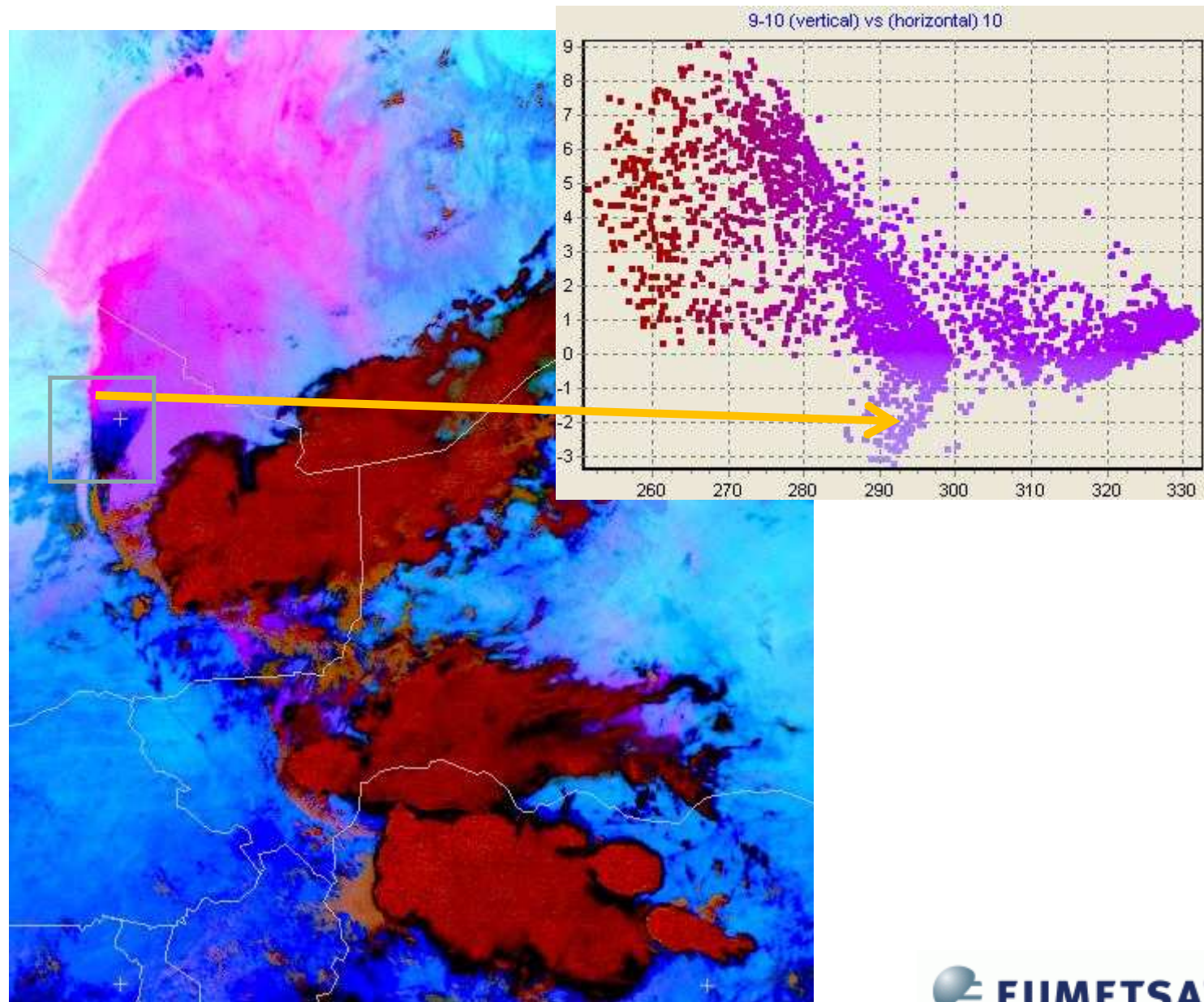
	Σa	Σs
Dust Ø 2 µm	0.1	0.05
Dust Ø 20 µm	1.2	0.06
CrossSect 20/2	1200	120
Sect/mass		10

Dust and water vapour in competition

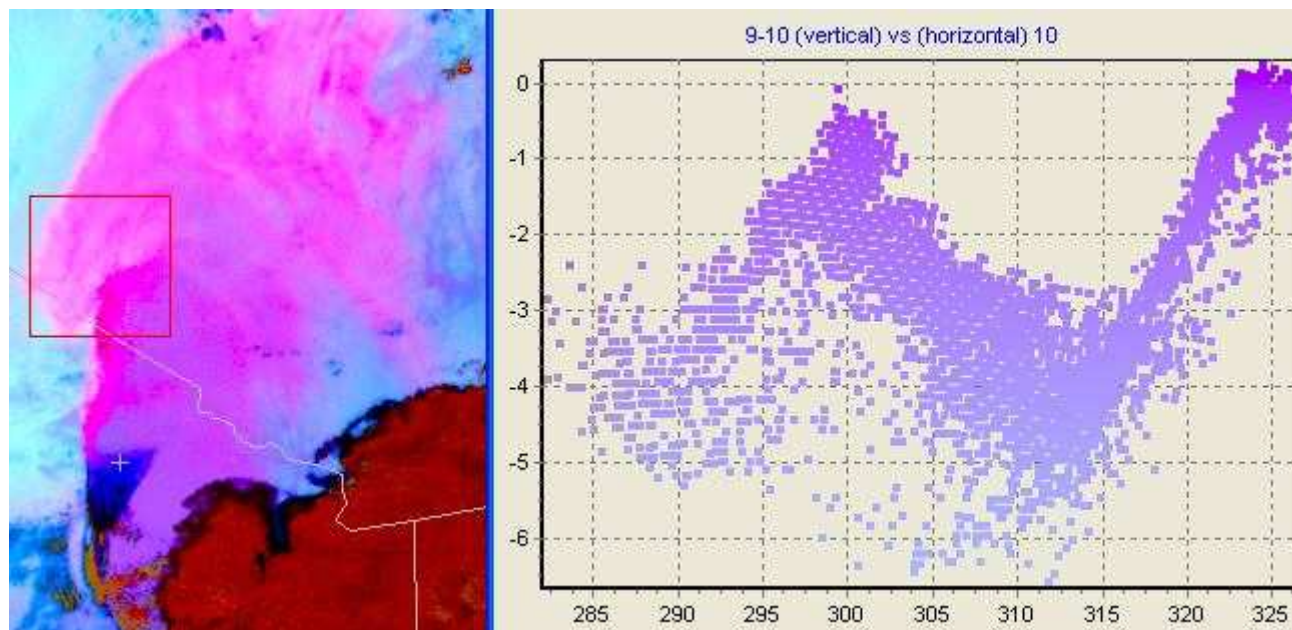


2006-Feb-23 12:00 Sahara desert, 10E, 27N



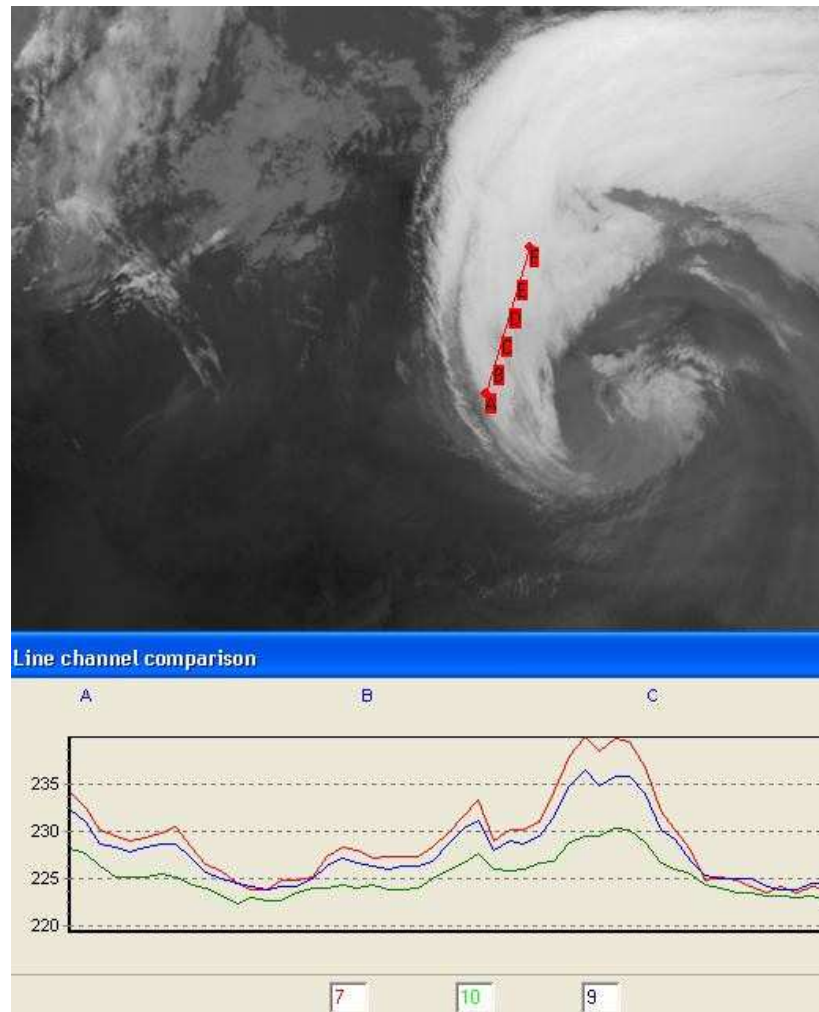


2010 Jun 09 12:30 UTC
Dust (and convection) over
the Sahara desert
DUST RGB



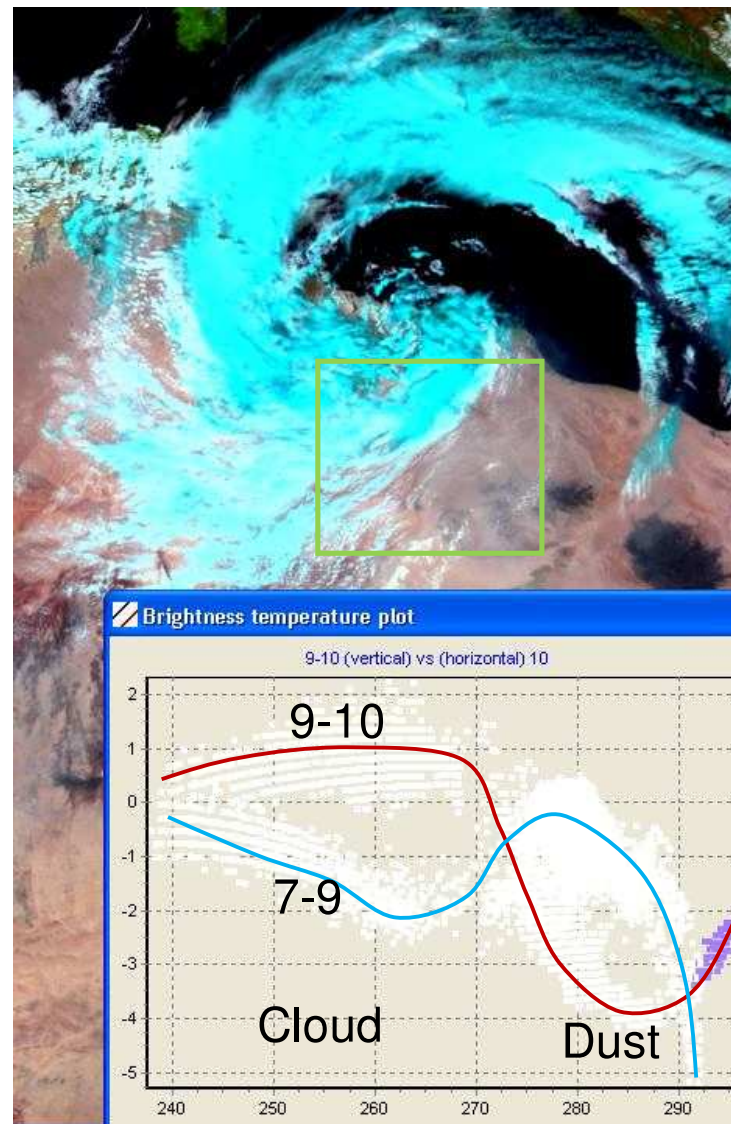
Cloud with dust shows 7-9-10 big differences

2007-Feb-22 12:00 Sahara desert, 7E, 32N



Cloud with dust shows 7-9-10 big differences

2007-March-09 12:00 Sahara desert, 14E, 29N



References:

<http://earthobservatory.nasa.gov/Features/Aerosols/page5.php>

- Large dust absorption of infrared radiation over Afro-Asian regions: evidence for anthropogenic impact Satheesh, S.K.; Deepshikha, S.; Srinivasan, J.; Kaufman, Y.J. Geoscience and Remote Sensing Letters, IEEE Volume 3, Issue 3, July 2006 Page(s):307 – 311
- Models of the Aerosols of the Lower Atmosphere and the Effects of Humidity variations on their optical properties. P. Shettle, R W Fenn. Sept.1979
- Absorption of sunlight by dust as inferred from satellite and ground-based remote sensing. Y J Kaufman et al.
- Geophysical Research Letters, vol.28, no.8 1497-82, April 2001
- Desert Dust in the global system. Springer. A S Goodie, N J Middleton