

Giant Dust Particles

Claire Ryder

NERC Independent Research Fellow/Associate Professor
Department of Meteorology, University of Reading, UK

With acknowledgements to:

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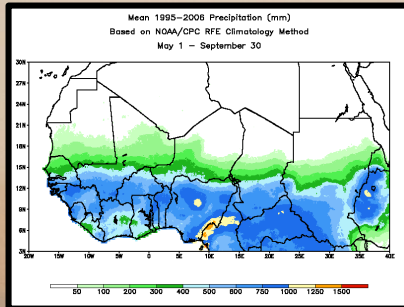
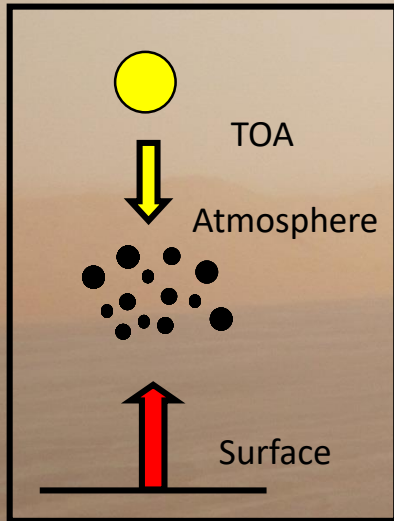
University of Manchester: Hugh Coe, Tom Choularton, Martin Gallagher, Keith Bower, Jonny Crosier,
Gary Lloyd, Dantong Liu, Jonny Taylor

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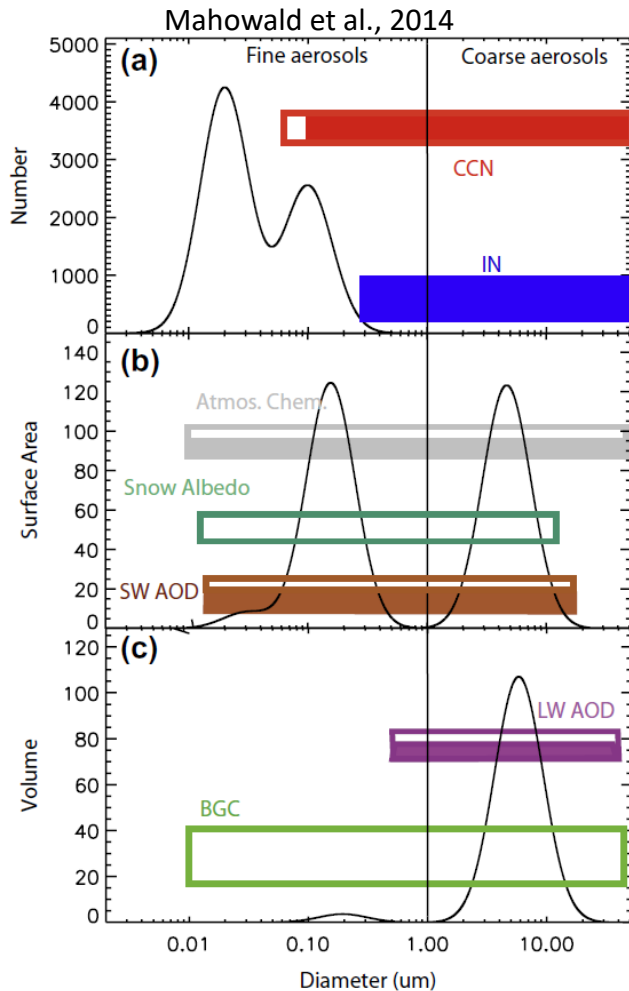
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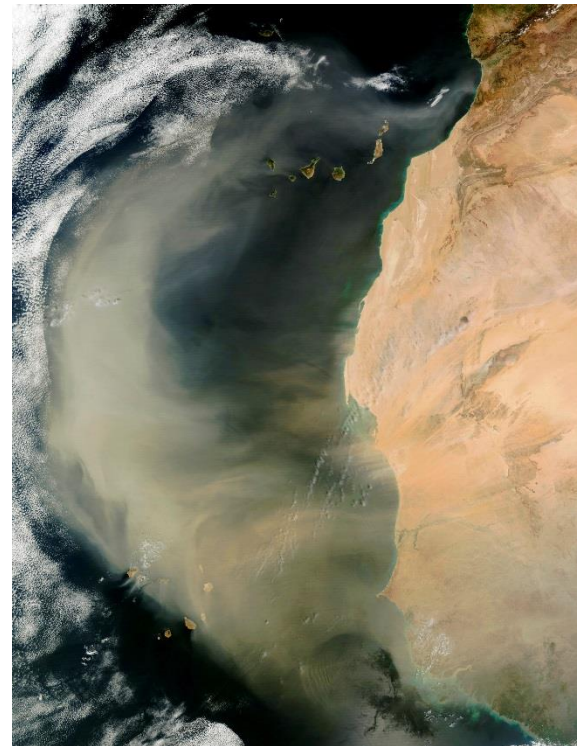
The Importance of Dust



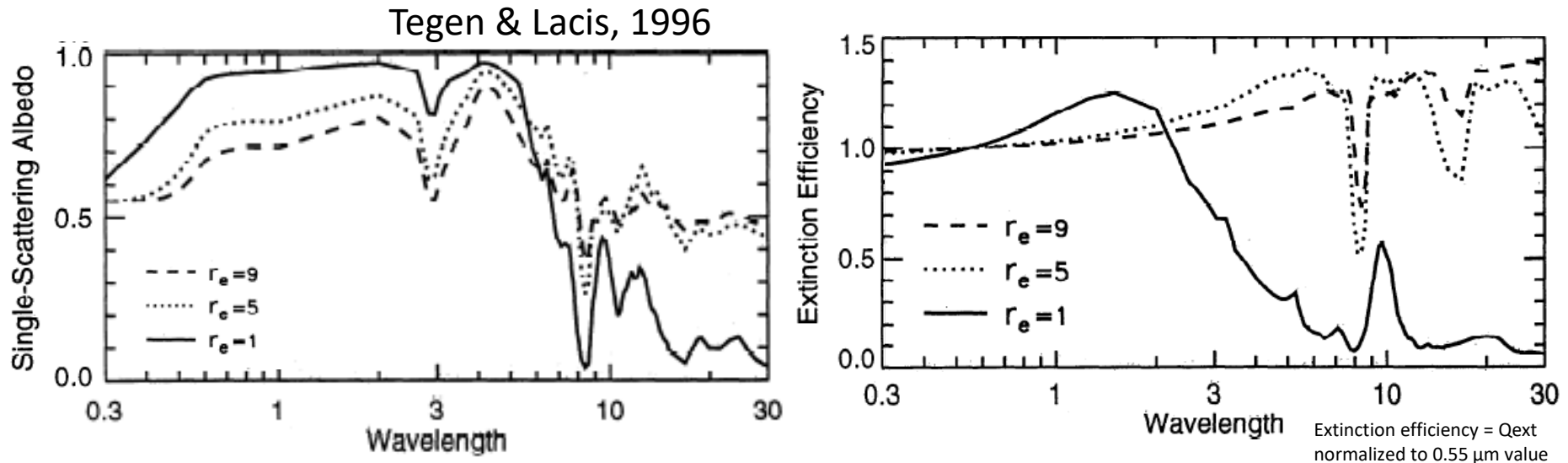
Importance of Dust Size Distribution



- Transport distances
- Dust Layer altitude



Dust size: the radiation perspective



- Solar wavelengths:
 - Larger particles reduce SSA
 - TOA forcing more positive, more atmospheric heating
- Terrestrial wavelengths:
 - Larger particles increase the extinction efficiency
 - Stronger longwave radiative effect

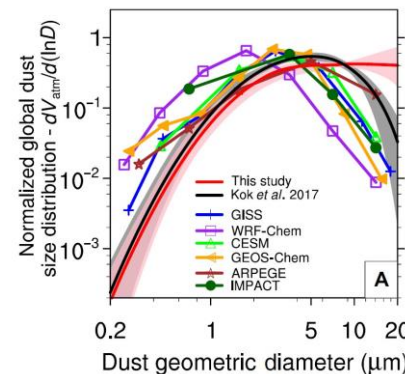
Motivation

- Historically: assumed coarse particles rapidly deposited
- Challenge for measurements, especially airborne, coarse mode frequently not measured at all
- Challenges where coarse mode was measured:
 - Inlets reducing transmission of coarse particles
 - Internal pipework causing loss of coarse particles
 - Uncertainties from optical scattering measurement techniques
- Last 10 years: airborne dust observations progressed, measuring larger particles, avoiding inlets and using non-optical techniques
- Multiple publications now report the presence of coarse and giant dust particles
- Models rarely include dust particles larger than $20\mu\text{m}$, $d > 5\mu\text{m}$: models start to underestimate dust concentration
 - Model dust is too cooling
- Aim:
 - Quantify the presence of coarse dust from aircraft observations in different regions
 - Quantify its contribution to optical properties

Giant dust observed in long range dust transport
van der Does et al. (2018)

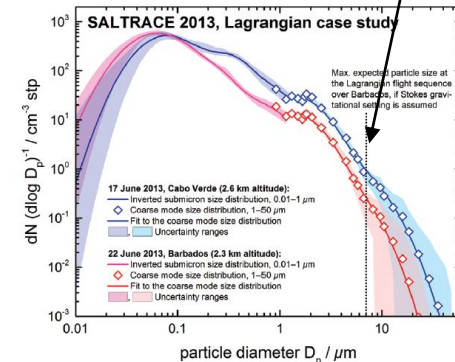


Max expected size at Barbados from gravitational settling



Adebiyi & Kok, 2020, Sci Adv:

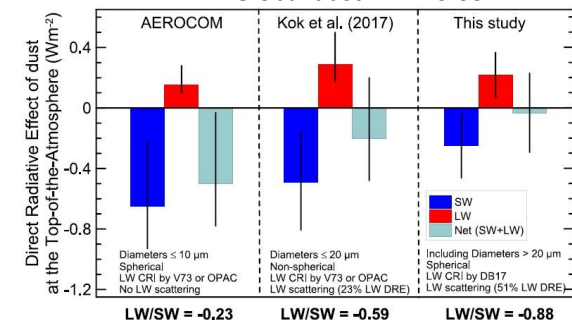
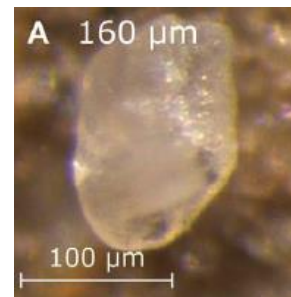
- 4 times more coarse dust than climate models simulate
- Adds 0.15Wm^{-2} warming



Weinzierl et al., 2017, BAMS

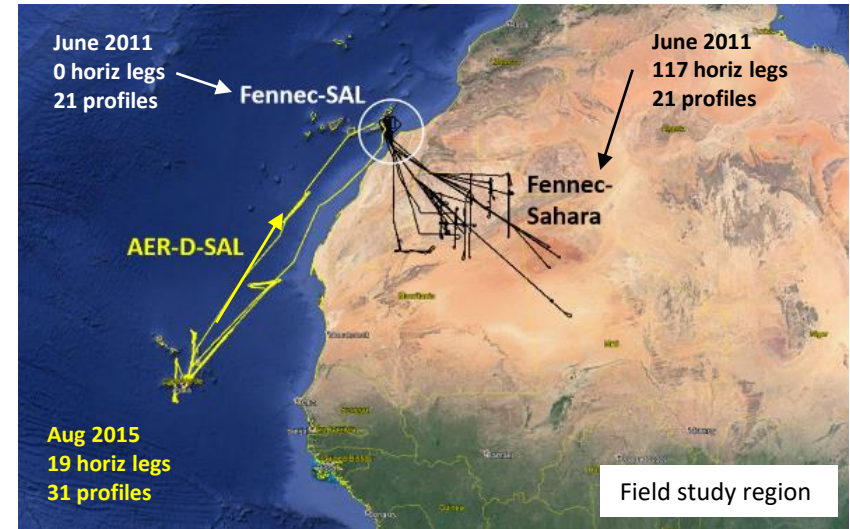
Di Biagio et al., 2020, GRL:

- Size extended beyond $20\mu\text{m}$, CRI updated
- Global dust DRE = 0.03Wm^{-1}



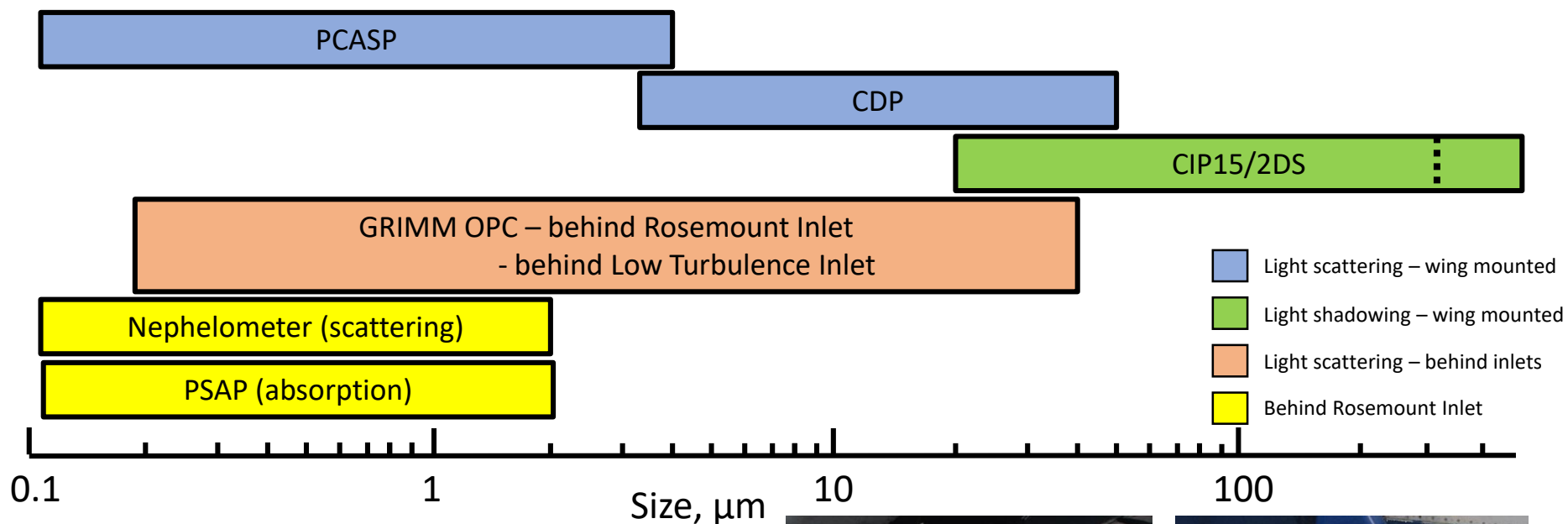
Fennec and AER-D Airborne Campaigns

- Fennec
 - June 2011/June 2012
 - Canary Islands to Mauritania & Mali
 - Far into remote desert
 - Fresh uplift, aged background dust
 - Ryder et al. (2015)
- AER-D (AERosol Properties – Dust)
 - With ICE-D
 - August 2015
 - Praia, Cape Verde Islands
 - Marengo et al. (2018)
- Fennec Movie:
 - Into the Cauldron
 - <http://fennec.ouce.ox.ac.uk/movie.html>

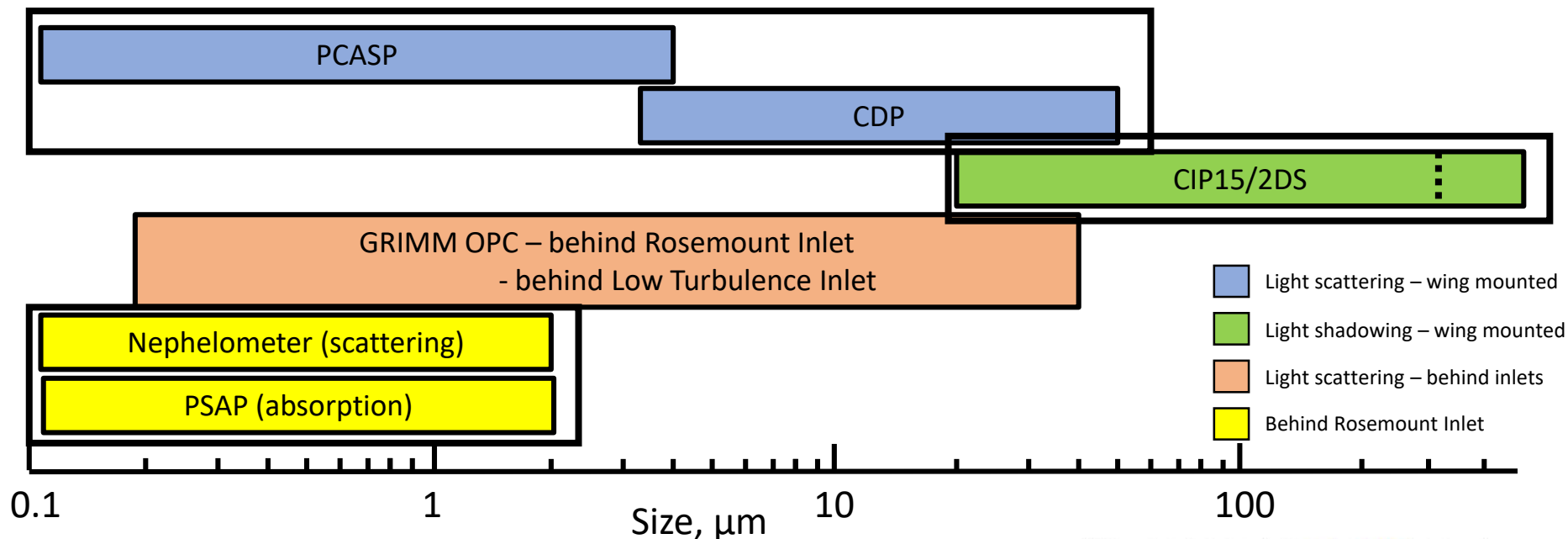


FAAM BAe146
Research Aircraft

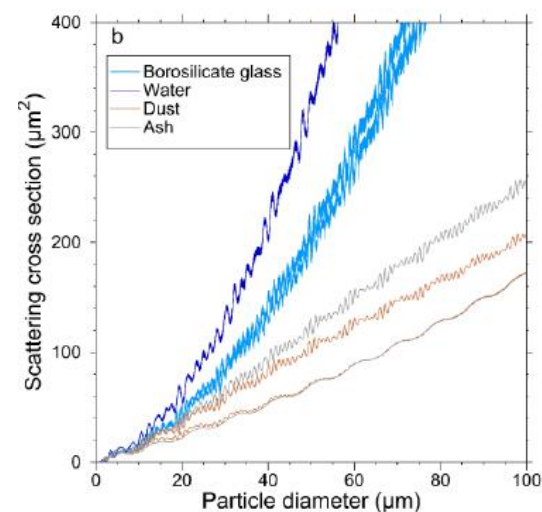
Airborne Measurements of Aerosol Size



Airborne Measurements of Aerosol Size



- Light scattering
 - Scattering cross-section converted to particle size
 - Depends on refractive index (composition) of particle
 - Not a unique solution – uncertainties can be large
 - Wing Probes (no inlets)
- Light shadowing
 - OAPs (optical array probes) – geometric
 - Sensitive to sizing metric/shape
 - Wing Probes (no inlets)
- In-cabin measurements (behind inlets)
 - Restrict measurement to a portion of the size range
 - Can bias optical properties



Rosenberg et al., 2012

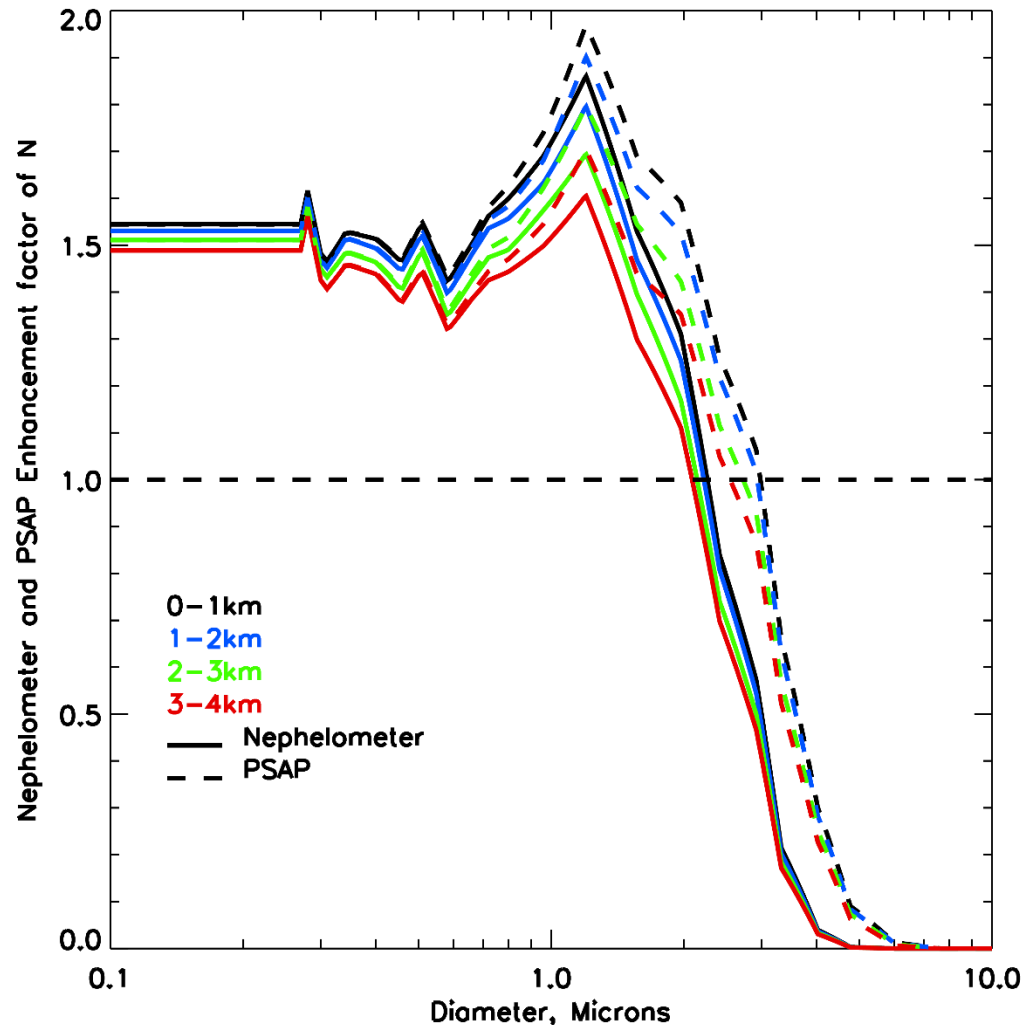
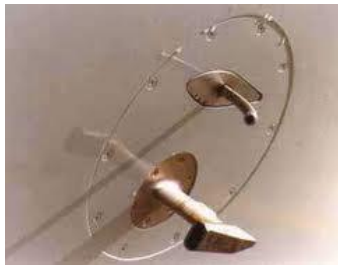
Fennec: Inlet Characterization

- Size distribution (PSD) measured behind
 - Rosemount inlets
 - Suffer loss of coarse particles
 - New Low Turbulence Inlet (LTI)
 - Size losses well-characterized
- Evaluate losses by Rosemount inlets for first time
- Modelled pipeline losses
- Nephelometer & PSAP (scattering, absorption: SSA) do not measure $d > 2.5 \mu\text{m}$
- Biased towards smaller sizes

Low Turbulence Inlet

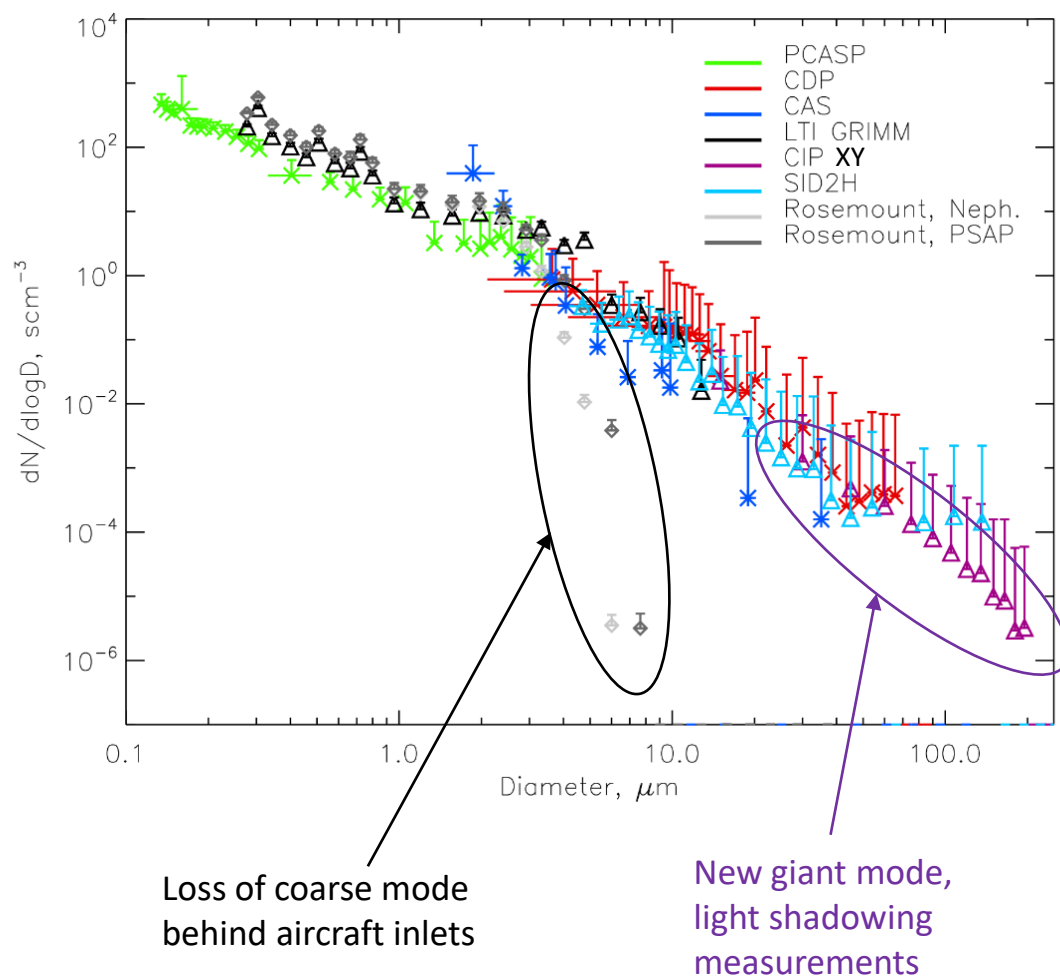


Rosemount Inlet

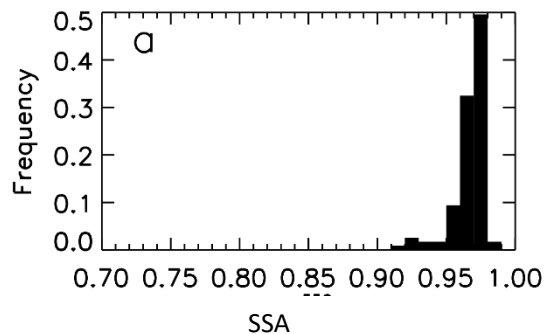


Fennec: Light Shadowing Measurements

- CIP15: Cloud Imaging Probe, wing-mounted
- Light shadowing Measurement technique
- First airborne CIP15 measurements in dust
- Confirmed presence of coarse and giant particles
- Validated optical probes measurements of coarse particles

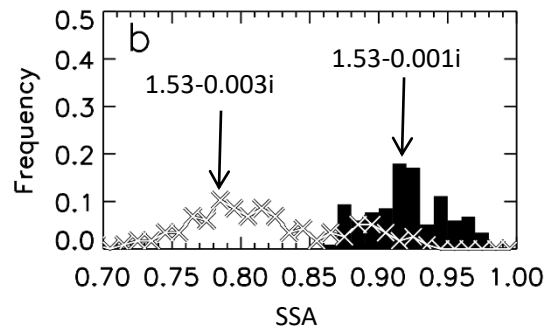


Impact of new Measurements on Optical Properties



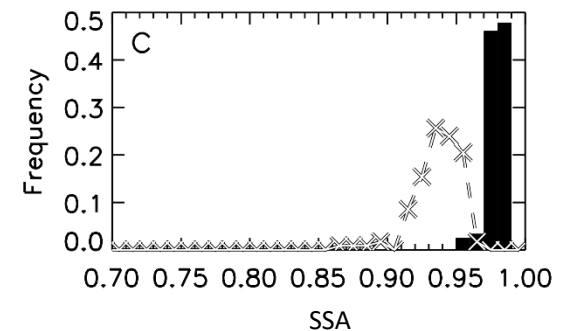
Observations, behind inlet

SSA=0.91-0.99



Mie Calculations,
Full Size Distribution

SSA=0.86-0.97

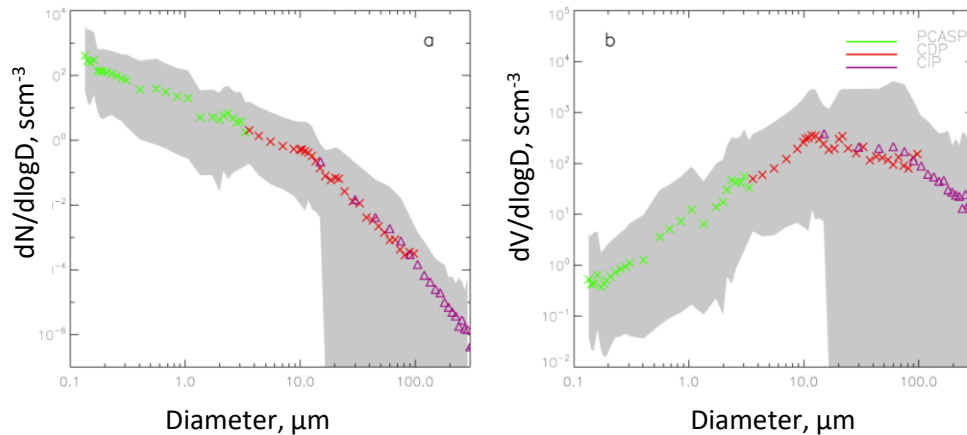


Mie Calculations,
Size Distribution behind inlet

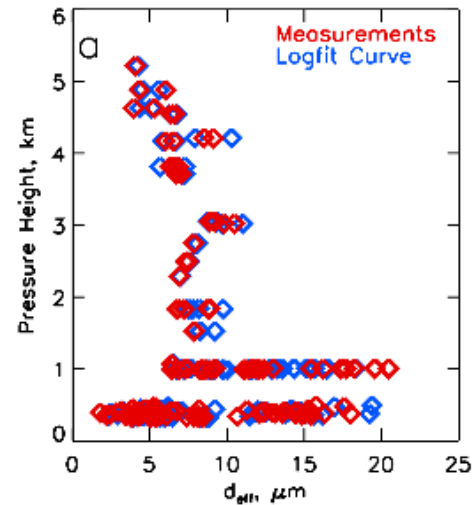
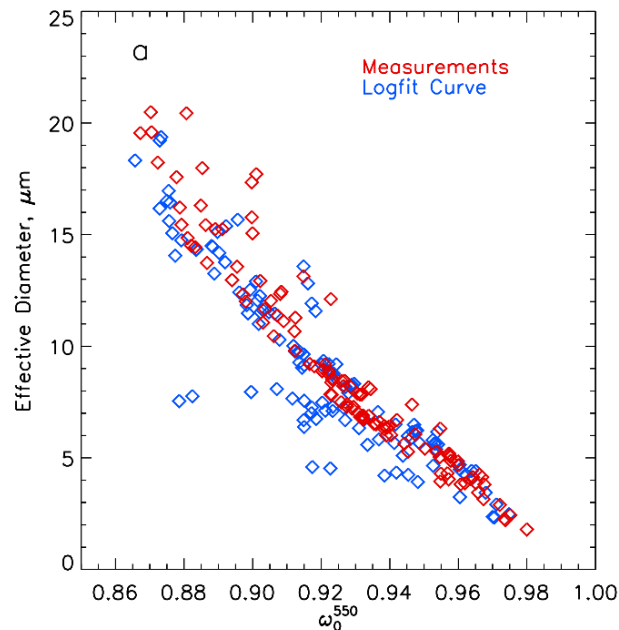
SSA=0.95-0.99

- Observations behind Rosemount inlets omit coarse particles
- 550nm SSA observations behind Rosemount inlets are an overestimate. I.e. absorption, (& warming effect) is underestimated
- Radiative transfer calculations: SW heating rates are 2-3 times greater with measured coarse particles

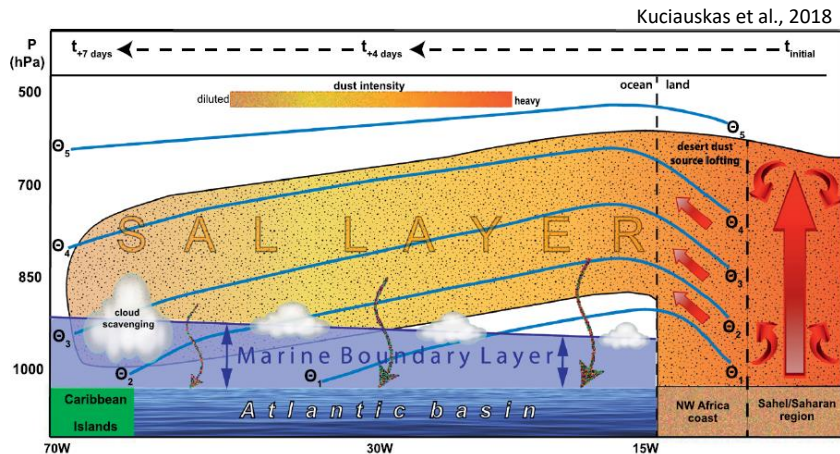
SSA vs Size in Fennec



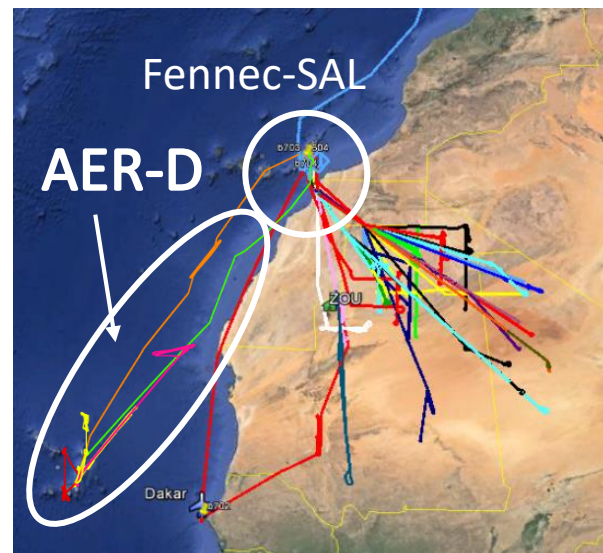
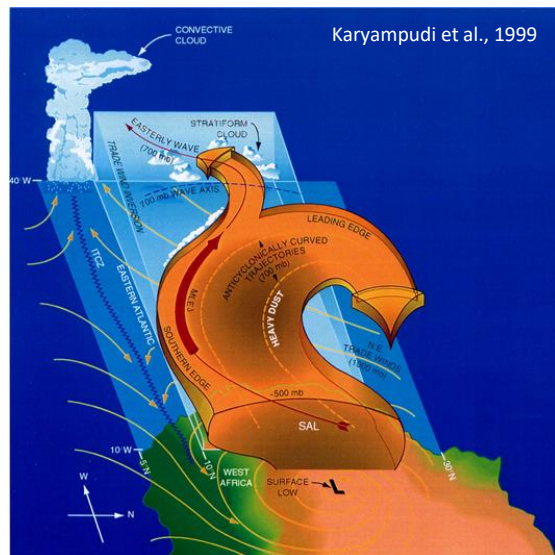
- During Fennec, size distribution was highly variable
- Volume PSD peak $\sim 20\mu\text{m}$
- SSA variability was strongly controlled by particle size distribution
- Largest sizes dominate at low altitudes, $z < 1\text{km}$
- D_{eff} 4-10 μm found up to 5km throughout Saharan Residual Layer
- Potential for long transport times and radiative interactions



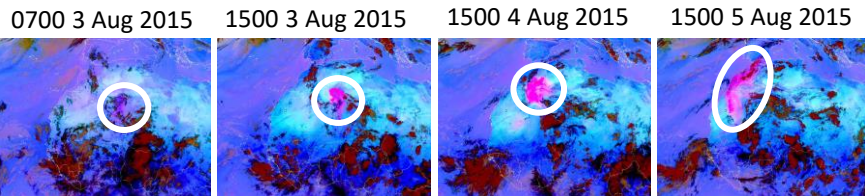
Coarse particles in long-range dust transport



- Coarse particles are prevalent over the remote Sahara
 - Are they present, and to what extent, over the tropical Atlantic?
1. Observations from Fennec over Canary Islands
 2. Observations from AER-D in the Saharan Air Layer



Impact of Transport on size: Fennec



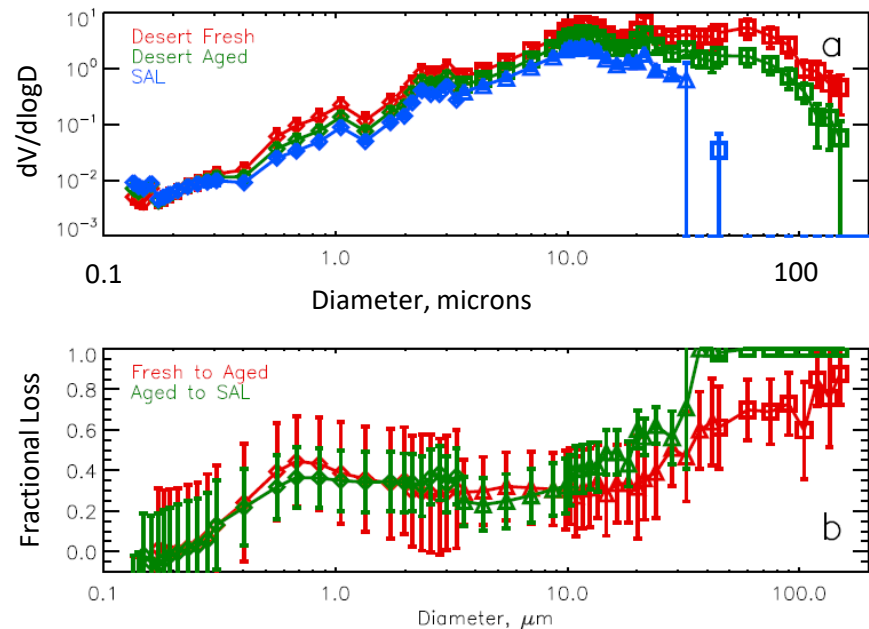
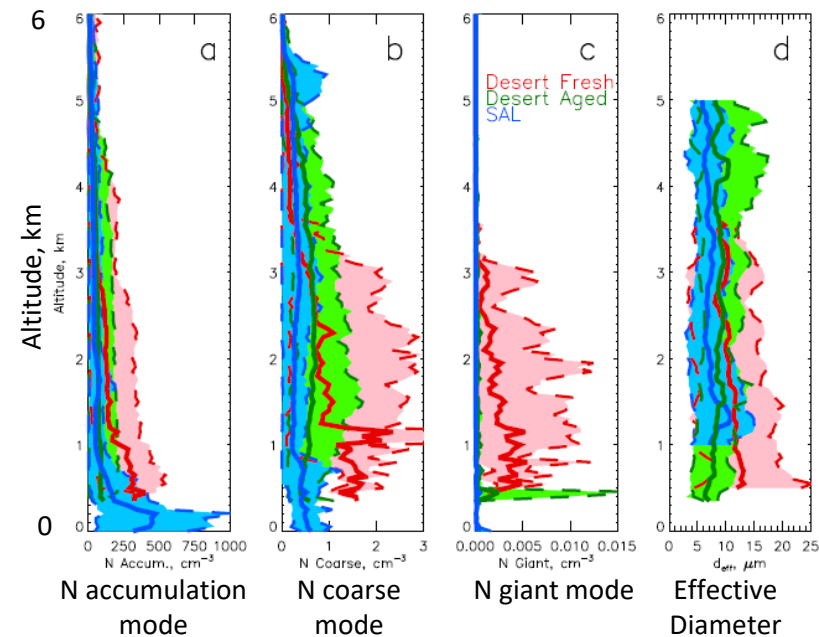
Estimating dust age:

Fresh: <12h since uplift

Aged: >12h since uplift

SAL: at Canary Islands, 18-120h since uplift

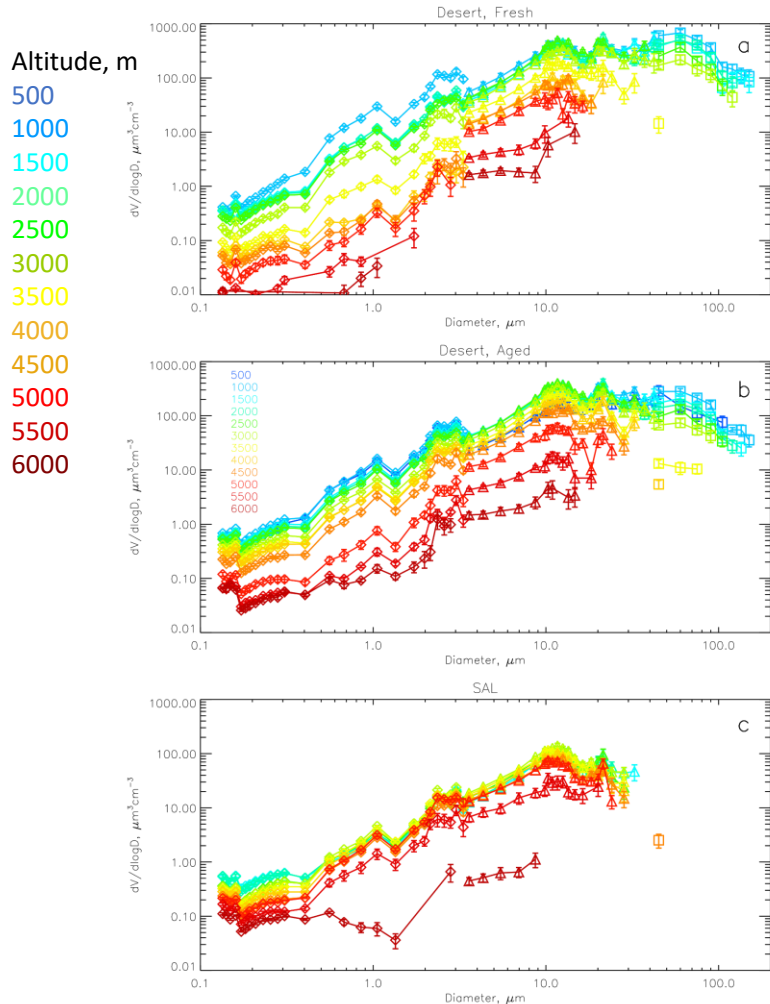
- Giant mode ($d > 37.5 \mu\text{m}$) expected to fall out within $\sim 12\text{h}$ (5km SALB) – supported by observations
- Coarse mode fall 5km in 20h ($30 \mu\text{m}$) or several days ($4 \mu\text{m}$) – not observed



Ryder et al.,
2013a, GRL,
after Maring
et al., 2003

Coarse/giant particles transported further
than explained by settling velocities

Size Distribution: Transport & Altitude

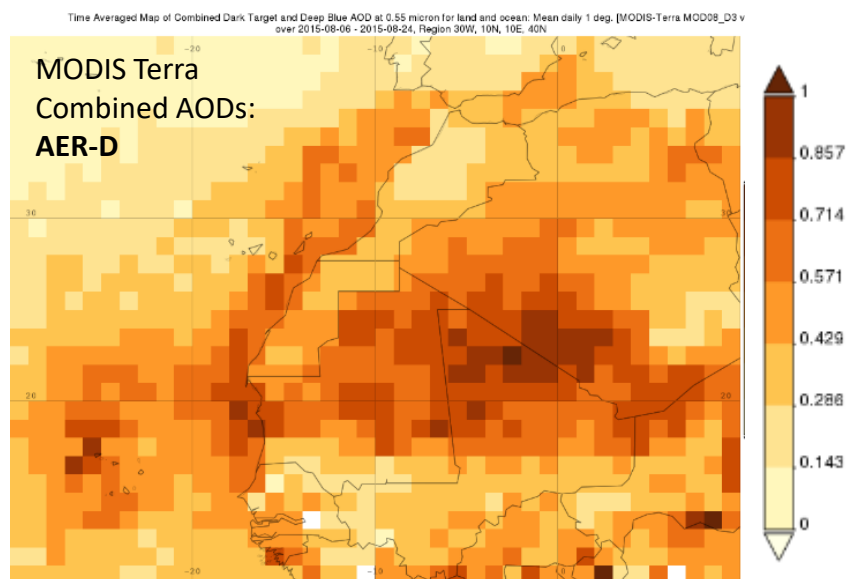
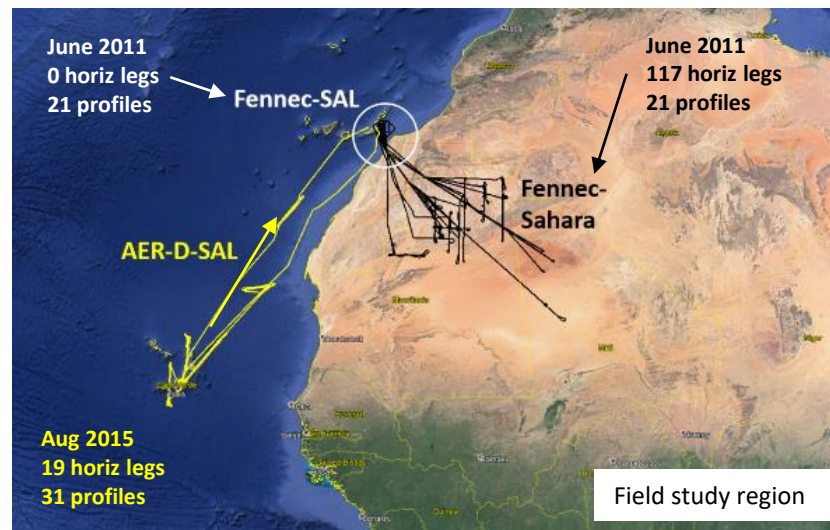


	SSA ₅₅₀	ξ_{550}	MEC ₅₅₀ /m ² g ⁻¹
Fresh (z<1.5km)	0.92	0.76	0.15
Aged (z<1.5km)	0.94	0.74	0.23
Aged (full column mean)	0.94	0.75	0.23
SAL (full column mean)	0.95	0.74	0.39

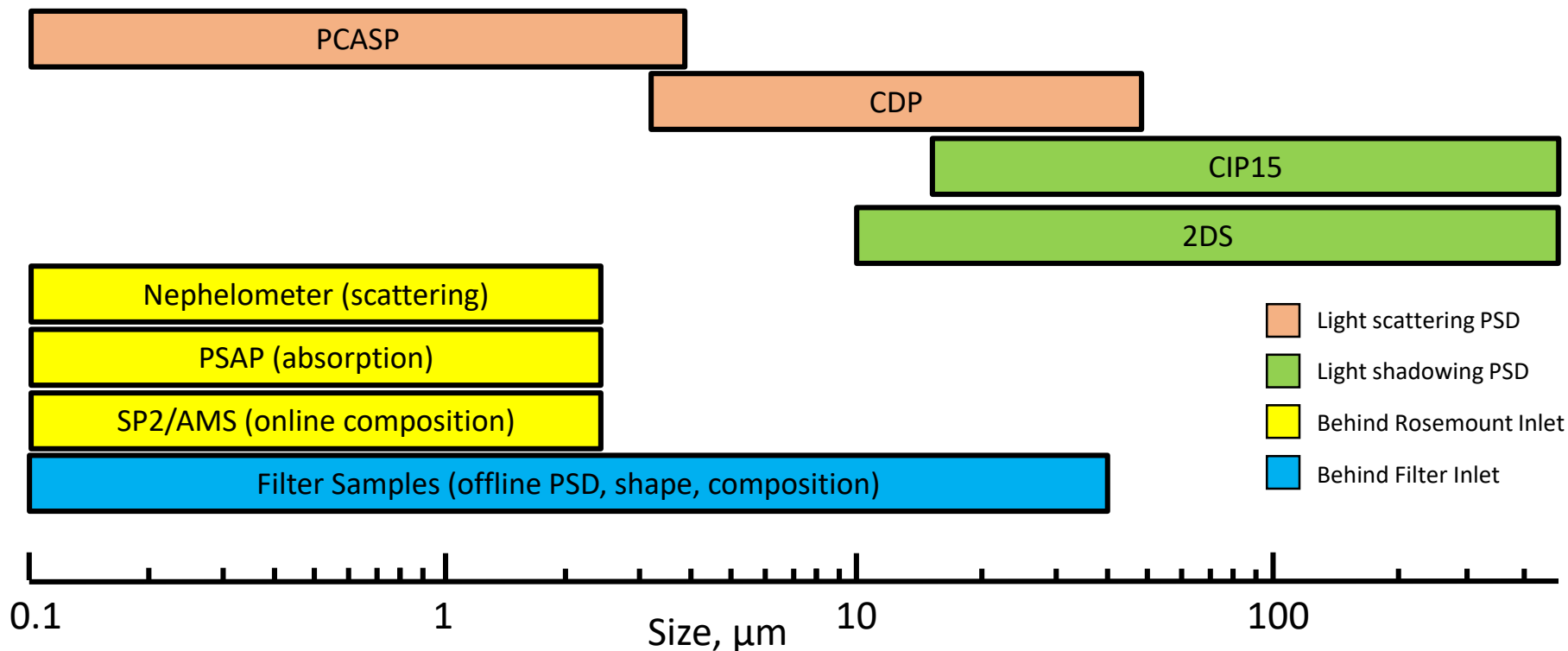
- Fresh
 - Coarse & giant consistently mixed up to 3km (lime green)
 - Giant particles absent above 3.5km
- Aged
 - Main change in concentration occurs at 5km (red) – top of SABL
- SAL
 - Well-mixed vertically, even more so than aged-desert. Few particles $d > 40\mu m$.

AER-D Airborne Campaign

- With ICE-D, August 2015
- Praia, Cape Verde Islands
- AODs up to 2.0, mostly 0.3-0.6
- AER-D publications:
 - Marenco et al. (2018, ACP): AER-D summary, extreme dust front event & giant particles
 - D. Liu et al. (2018, ACP: dust & hematite content)
 - Price et al. (2018, JGR, dust and ice nuclei)



AER-D: FAAM In-situ Measurements of Aerosol

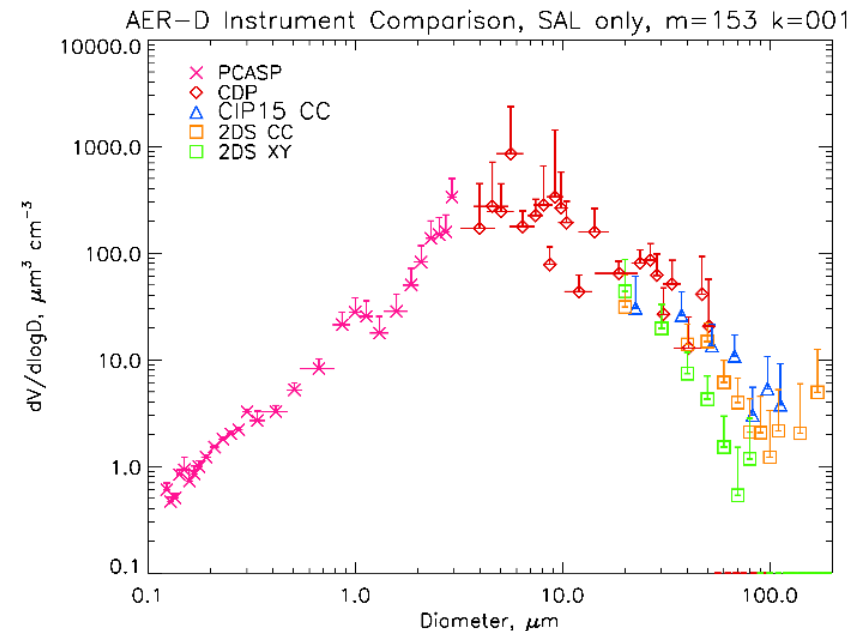
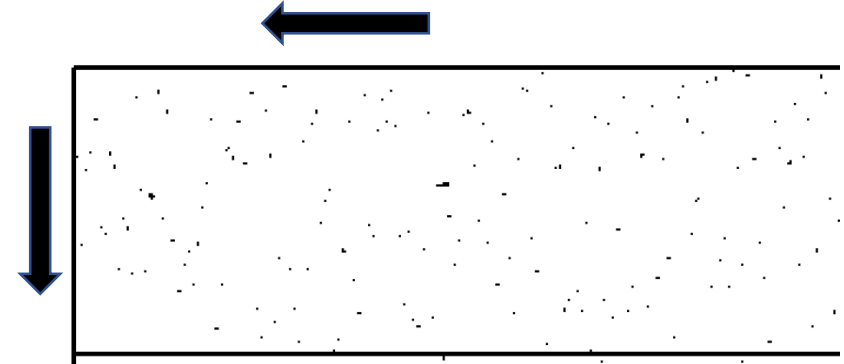


- 2DS in addition to CIP15
- Online accumulation mode composition
- Filter Samples

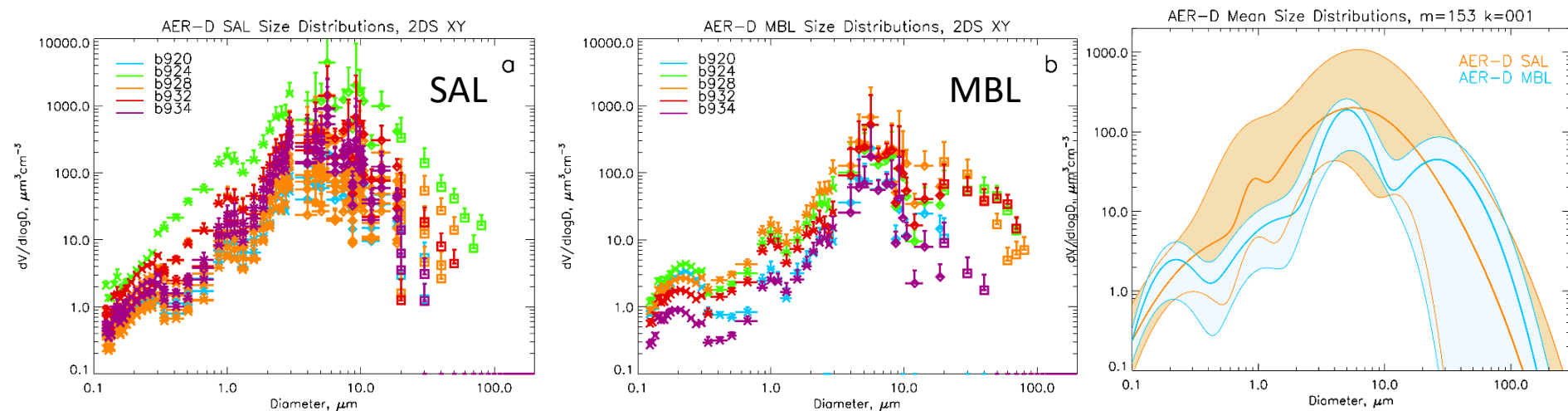


Optical Array Probe Sizing

- CIP15/CIP100/2DS
- Light shadowing techniques
- Size measured across:
 - Direction of aircraft motion
 - Aligned with photodiode array
- Sizing by:
 - Mean x & y dimensions (XY)
 - Smallest circumscribing circle (CC)
 - Other methods (McFarquahr et al., 2017)
- Size distribution sensitive to size metric
- No requirement for refractive index knowledge/assumptions or Mie scattering conversions
- **2DS XY** selected: consistent with Fennec (CIP15 XY)
 - Coarse & giant mode represent lower bounds



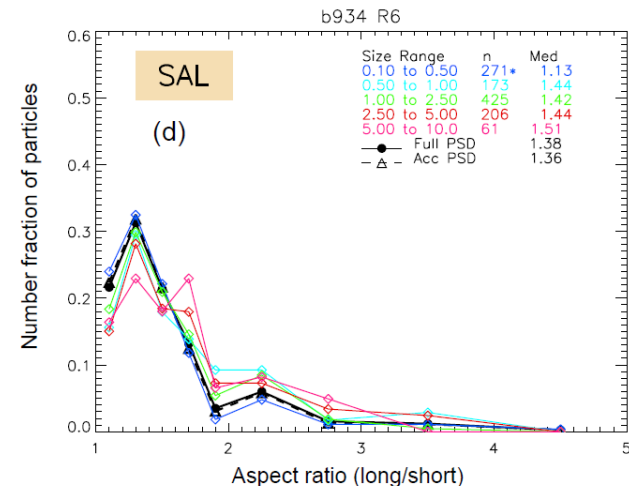
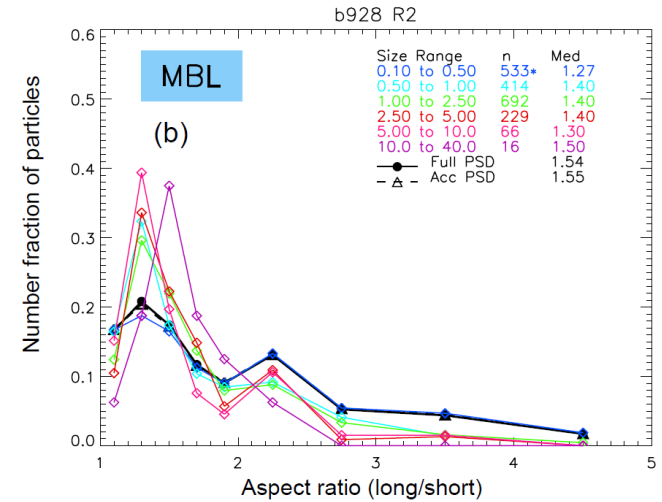
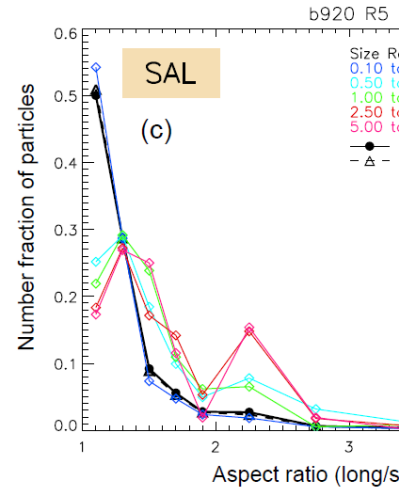
AER-D Size Distributions



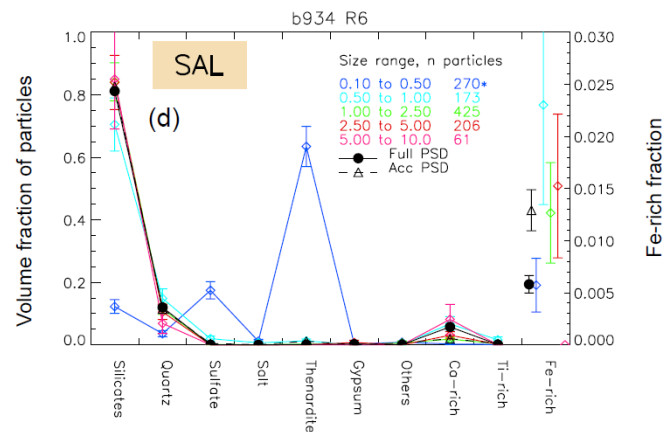
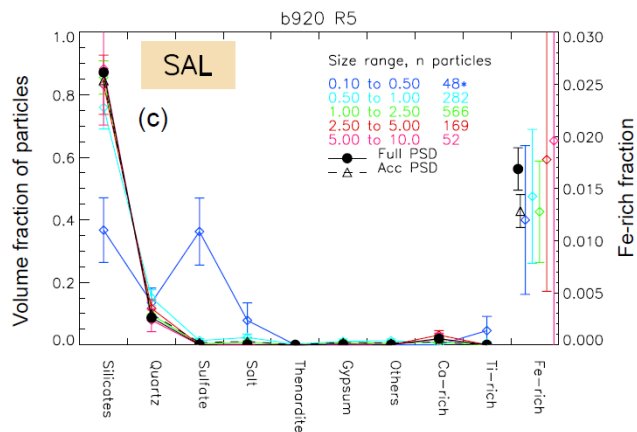
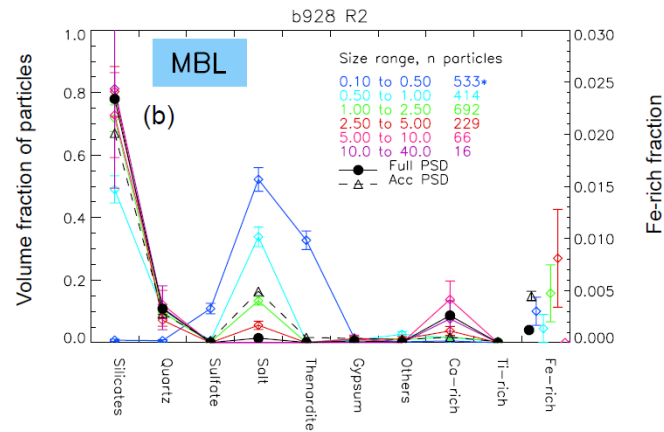
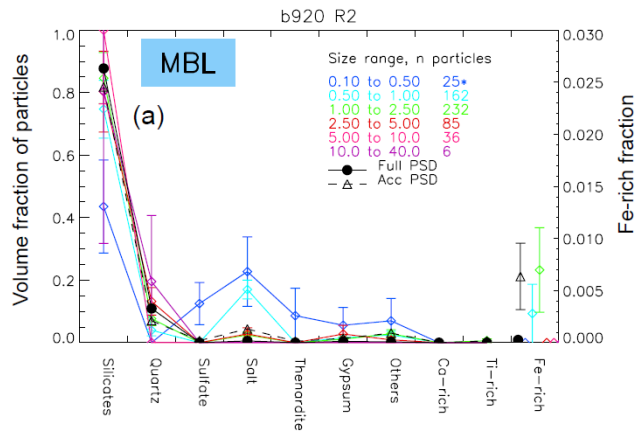
- SAL PSD shape consistent; concentrations rise and fall
- MBL PSD – enhanced fine mode, giant mode present d \sim 10-80 μm when loadings above in SAL are high
- Fit lognormals to smooth out instrumental noise, reproducibility

Filter Sample Analysis: Aspect Ratios

- Ellipse fitted to area projecti particle
- Diameter = circular area equivalent diameter
- Aspect Ratio (AR) measured
- 2D projected size of 3D non-spherical particle
- Plate-like flat particles will fall flat on substrate and particles will be oversized
- Chou et al. (2008)
height=length/3
- D=5-40 μm median AR 1.3-1.5
- D=0.5-45 μm median AR 1.3-1.4
- MBL Giant mode more non-spherical

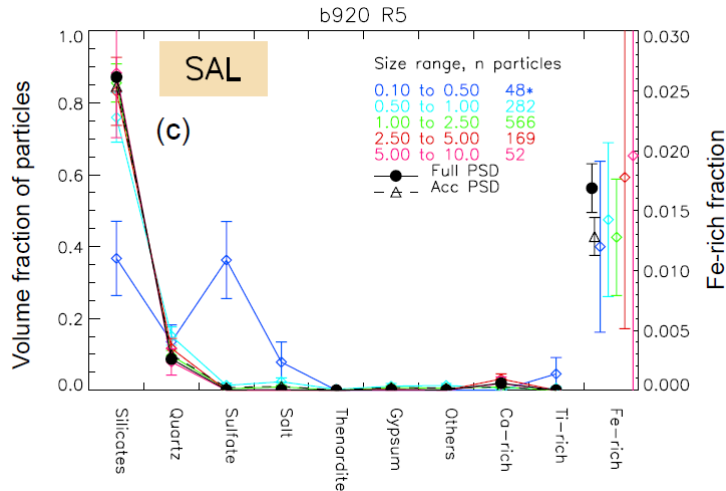


Composition from Filter Samples

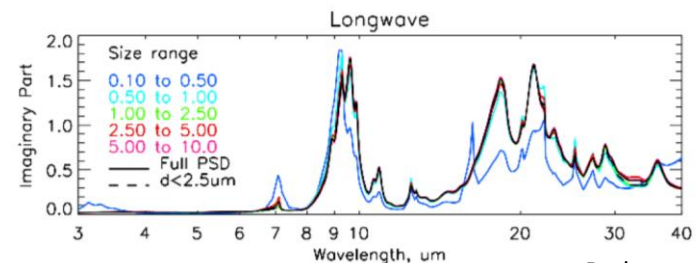
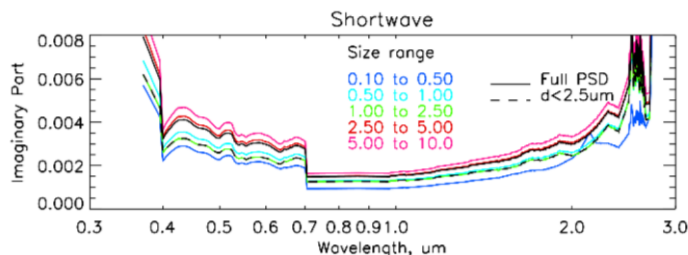
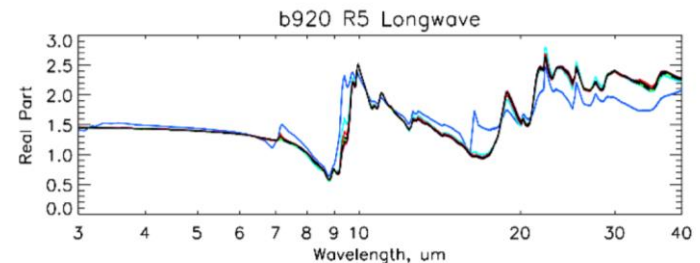
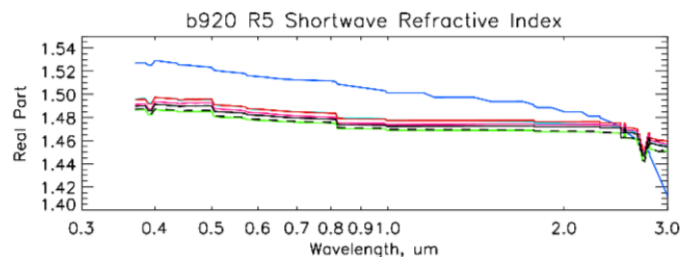


- Particle by particle SEM analysis
- $d < 0.5 \mu\text{m}$ dominated by sulfates & salts
- Giant mode MBL ($d > 10 \mu\text{m}$) is dust
- Dust dominates at $d > 0.5 \mu\text{m}$
- Fe-rich fraction small but important for refractive index; higher in SAL than MBL

AER-D Composition & Refractive Indices

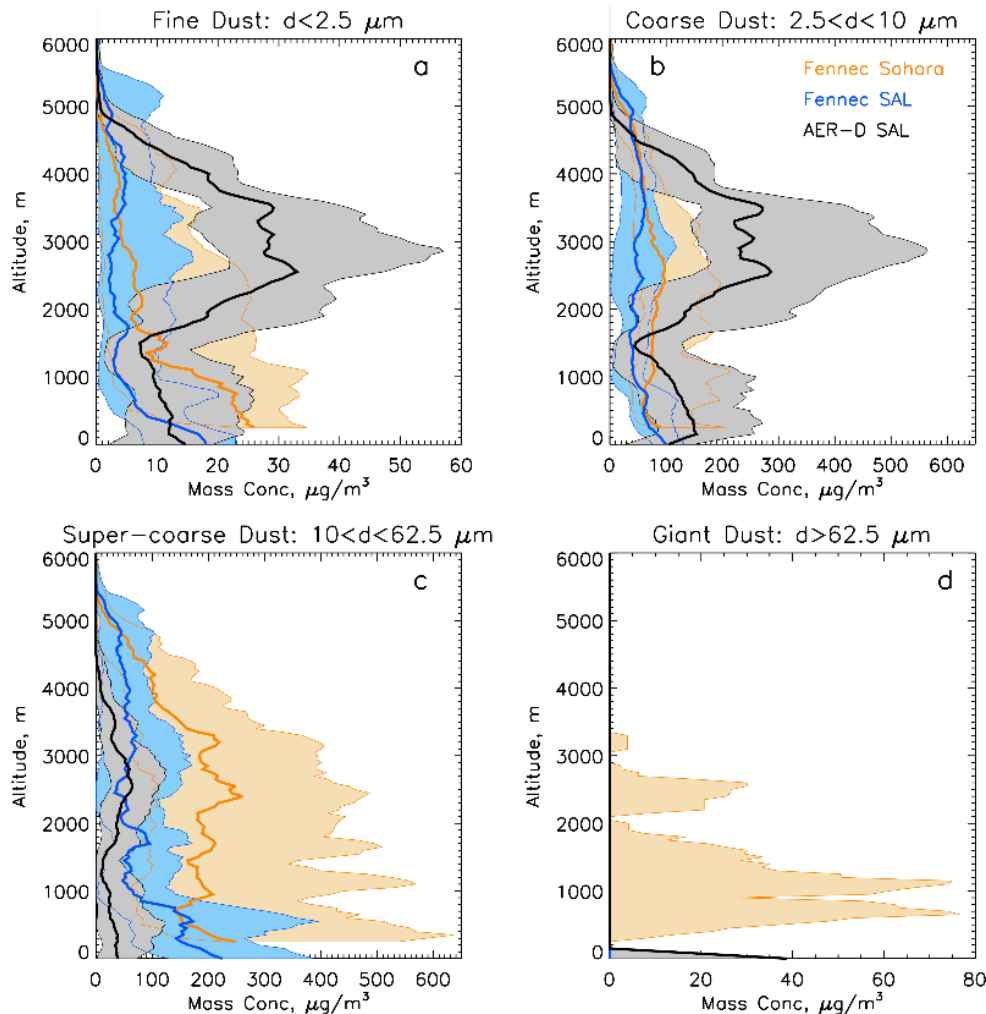


- Shortwave:
 - Real part $d=0.1-0.5\mu\text{m}$ high – sulfates
 - Increasing size \rightarrow higher imaginary part – Fe
 - Larger particles more absorbing (in this case)
 - Very sensitive to Fe content
- Longwave:
 - Sulfates content, and quartz vs silicate content is main control on RI
 - Differences for $d=0.1-0.5\mu\text{m}$ – sulfates
 - Quartz:silicate ratio fairly constant: not much variation with size



Combining Fennec-
Sahara, Fennec-SAL,
AER-D

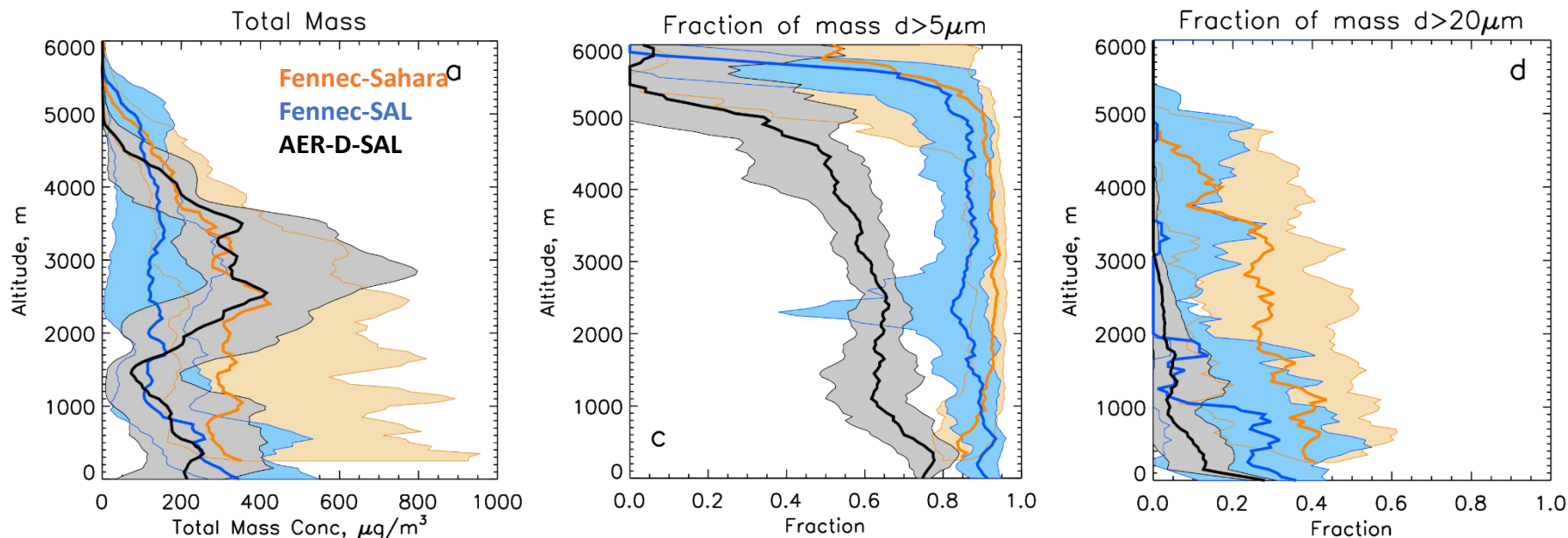
Dust Mass Concentration Profiles



- Largest mass over Sahara; Decreases with altitude; SAL well-mixed
- Desert higher in concentrations of super-coarse and giant dust
- AER-D: slightly different size distribution, more fine & coarse particles
- Aim to provide data suitable for model comparisons

Mass Concentration Profiles

Ryder et al., 2019



a) Fraction of mass $d > 5\mu\text{m}$

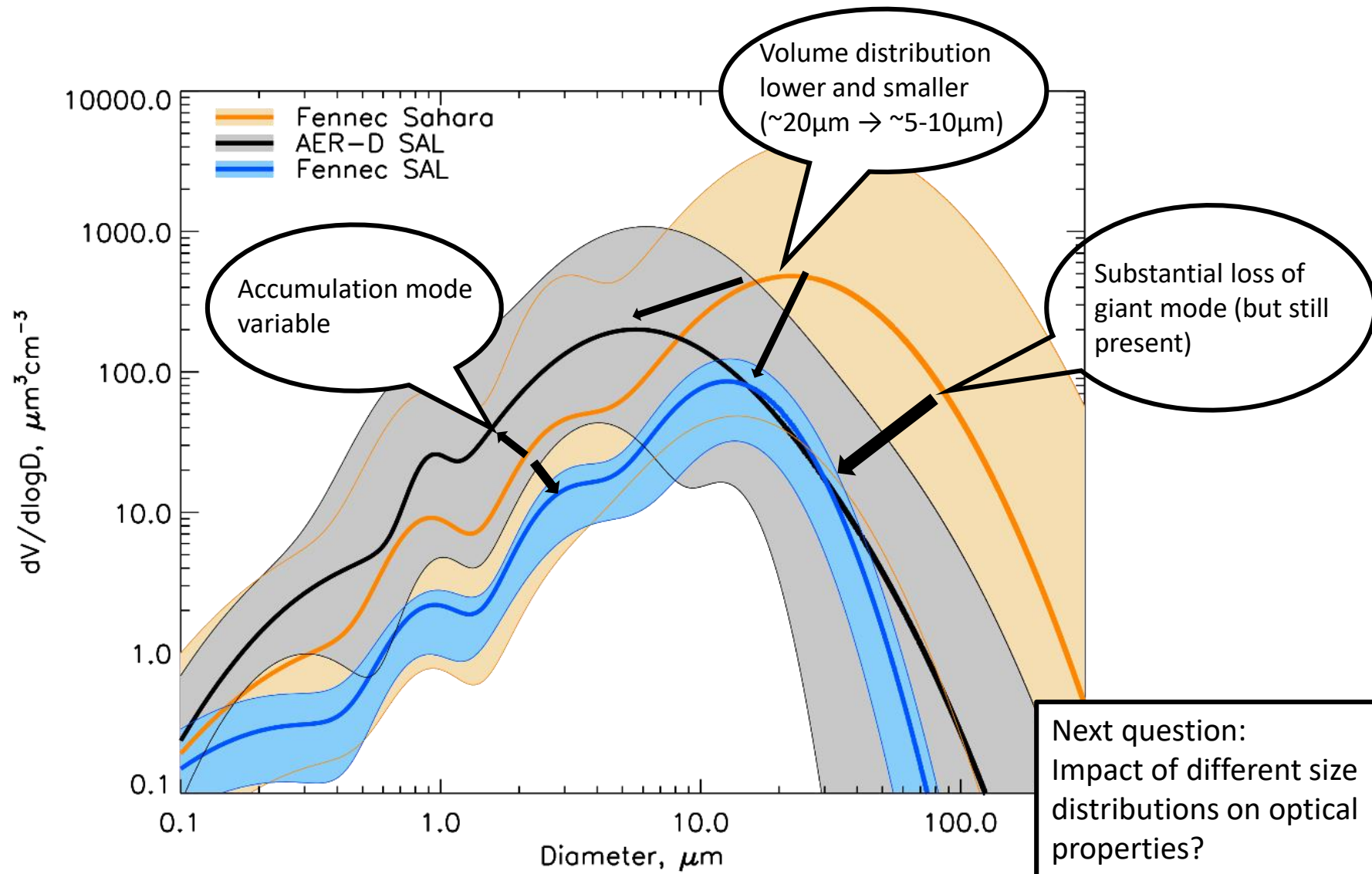
- Fennec-Sahara: 92% beneath 4.5 km
- SAL: 61-87%

b) Fraction of mass $d > 20\mu\text{m}$

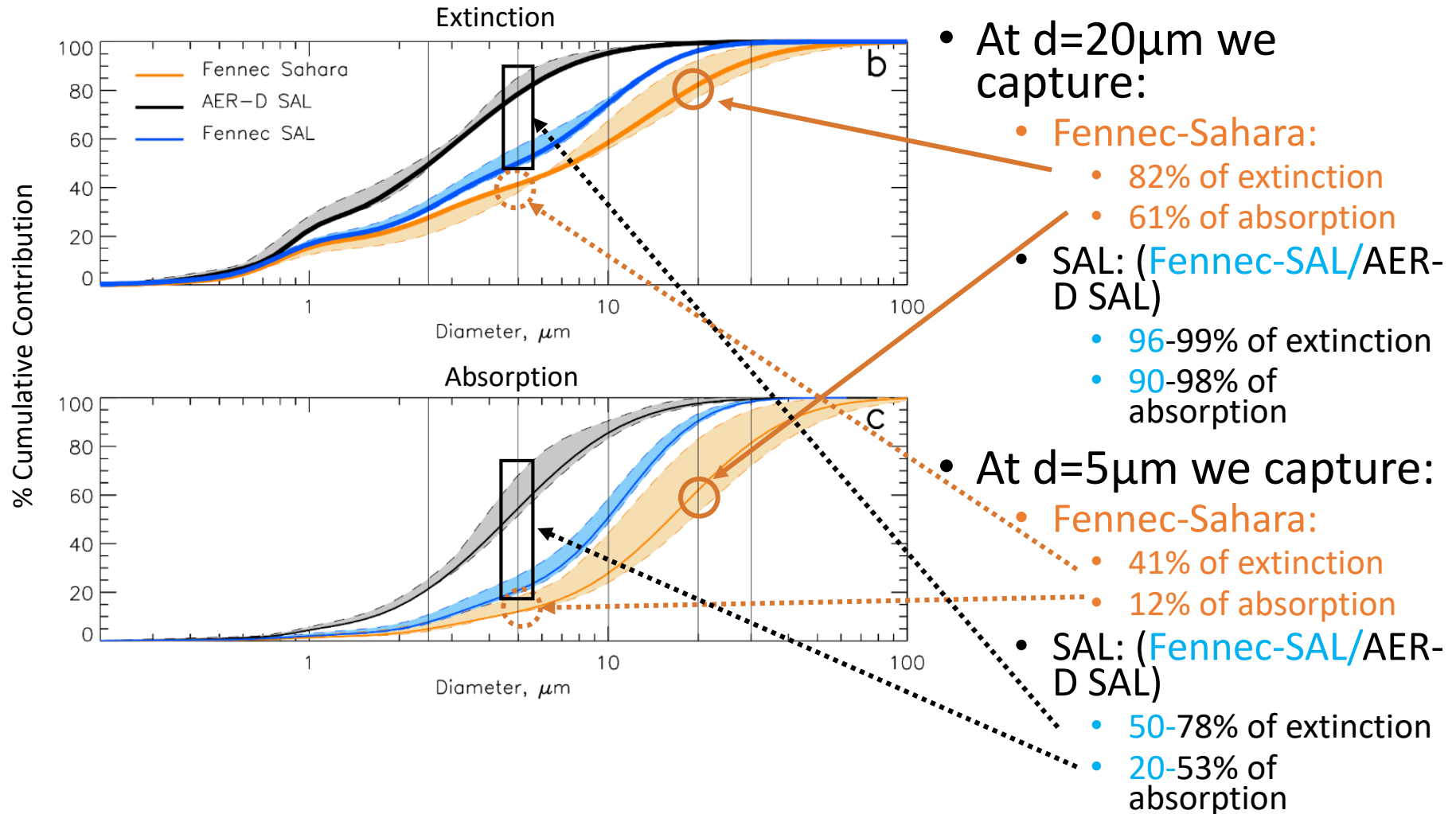
- Fennec-Sahara: 27% mass at $d > 20\mu\text{m}$
- SAL: 2%

- A significant amount of mass is being both completely excluded from models ($d > 20\mu\text{m}$) and underestimated by models ($d > 5\mu\text{m}$)

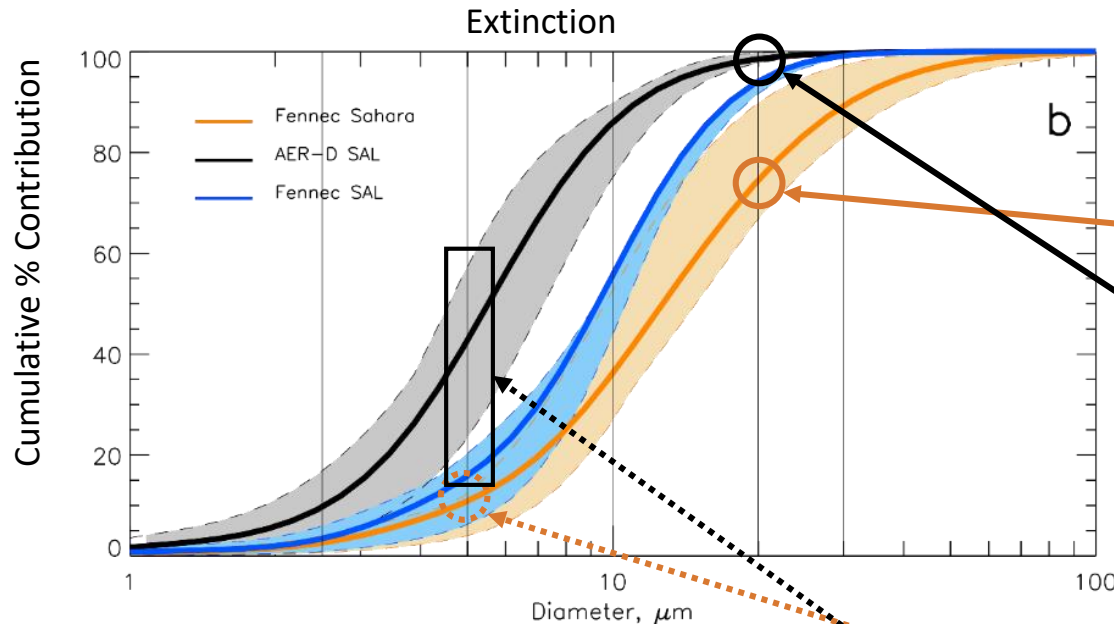
Size Distributions



Size Resolved SW Extinction & Absorption



Size Resolved LW Extinction



- At $d=20\mu\text{m}$ we capture:

- Fennec-Sahara:
 - 74% of extinction
- SAL: (Fennec-SAL/AER-D SAL)
 - 94-98% of extinction

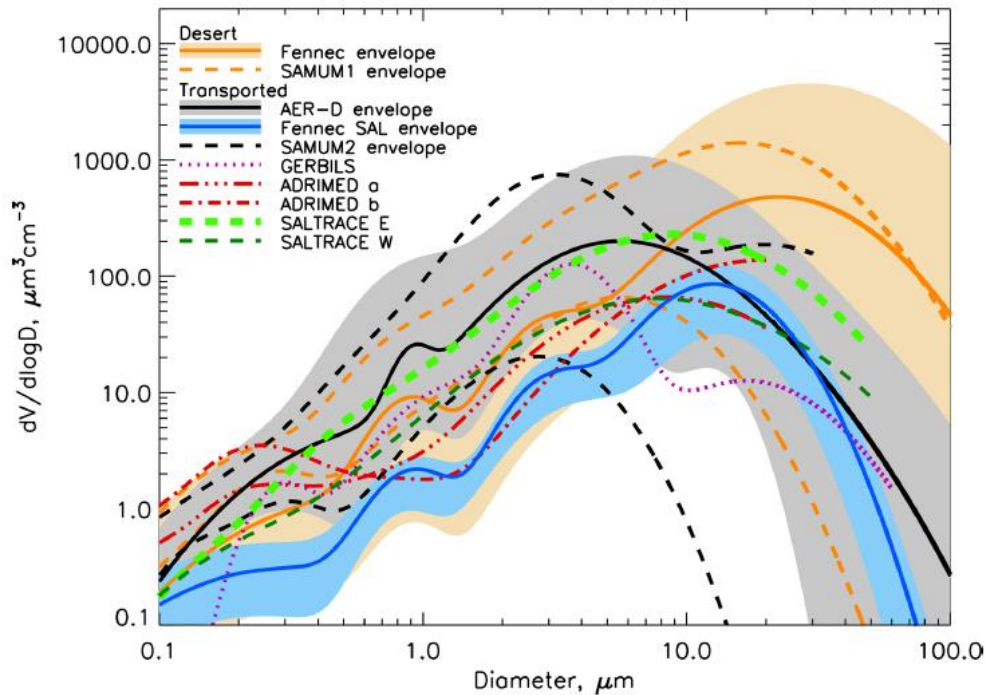
- At $d=5\mu\text{m}$ we capture:

- Fennec-Sahara:
 - 10% of extinction
- SAL: (Fennec-SAL/AER-D SAL)
 - 15-41% of extinction

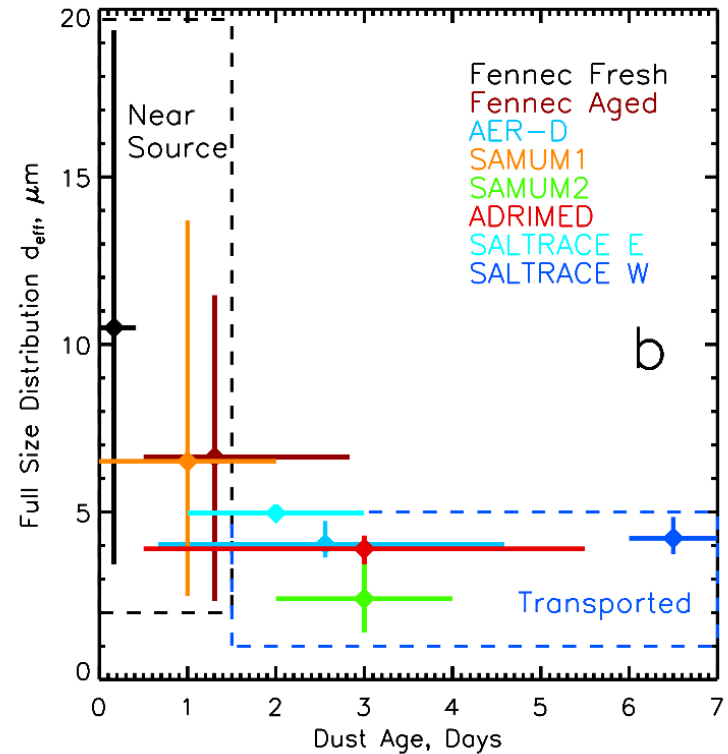
- **Models will significantly underestimate SW and LW extinction and absorption over the Sahara by excluding and/or underestimating the coarse dust concentrations**
- Greater underestimation of LW extinction than SW, shifts dust DRE to more positive values
- Changes to atmospheric heating from incorrect model dust properties may impact atmospheric circulation in dusty regions

Change in Dust Size with Age

Ryder et al., 2019



- Compilation of airborne observations measuring Saharan dust, including $d \geq 20 \mu\text{m}$
- Always a significant contribution from dust particles sized $d > 5 \mu\text{m}$
- Close to sources, there is also a strong contribution from particles larger than $20 \mu\text{m}$ diameter

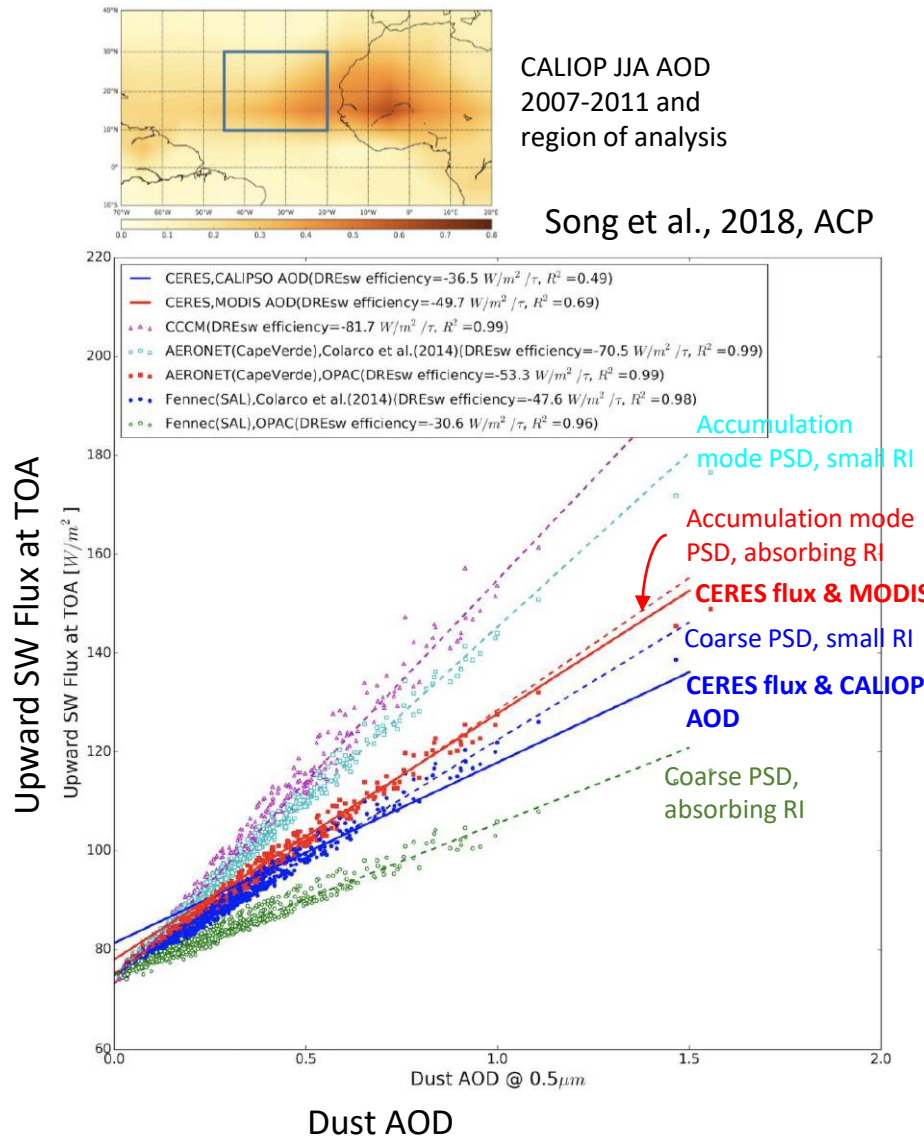


- Very large particles evident immediately after uplift with high d_{eff} values of 6 to $10 \mu\text{m}$
- d_{eff} decreases rapidly until around 1.5 days after uplift
- After this observations suggest little change in d_{eff}
- Size distribution stabilizes through transported regime

Summary

- Improved quantification of coarse & giant dust size distributions via:
 - Improved understanding of inlet behaviour
 - Use of Optical Array Probes
- Substantial coarse mode, and giant mode over desert
 - VMDs: Fennec-Sahara 21 μ m, Fennec-SAL 12 μ m, AER-D 6 μ m
- Impacts optical properties
 - E.g. 39% of absorption originates from $d > 20\mu\text{m}$ over desert
- Areas deserving further research:
 - Mechanisms & sensitivities for uplift, transport & deposition of coarse & giant particles

Satellite-based assessment of dust size



- Observed SW DRE efficiency (CERES, MODIS, CALIOP) vs RTM model values
- Sensitivity tests to dust properties:
 - PSD: Coarse (Fennec SAL) vs accumulation mode (AERONET)
 - Refractive index: OPAC (absorbing) vs Colarco 2014 (less absorbing)
- Most reflective dust → greatest DRE efficiency (smaller particles, less absorbing RI)
- Agreement with observations is possible with only 2 dust combinations:
 - Coarse PSD + low absorption
 - Accumulation mode PSD + high absorption
- LW sensitivity tests : Model minus CERES:
 - Coarse PSD, OPAC LW RI, $0.5 Wm^{-2}$
 - Accumulation mode PSD, OPAC LW RI, $1.6 Wm^{-2}$
- LW observations used to narrow down dust properties – only significant coarse mode presence can lead to reasonable agreement in both SW and LW
- High absorption (OPAC) of $\sim 0.008i$ deemed unlikely, $0.0015i$ more realistic (Balkanski et al., 2007)
- Significant coarse mode presence is required to match observations**